

Université Libre de Bruxelles
Faculté des Sciences
Service de Physique Théorique

Brane Physics in M-theory

Thèse présentée en vue de l'obtention
du grade de Docteur en Sciences
(grade légal)

Riccardo Argurio
Aspirant F.N.R.S.

Année académique
1997–1998

To my parents

Writing the acknowledgements for a thesis which represents four years of work is a difficult task, and this is why the acknowledgements are generally written few hours before printing the final version of the ‘book’. I apologize in advance for the dullness.

I wish to thank in the first place my advisor, or supervisor, or directeur de thèse, François Englert, for teaching me simply by his own way of working, what should be the attitude of any theoretical physicist. To be curious about the new developments, to question the foundations of any theory whatever ‘mainstream’ it is, to appreciate the subtle balance between mathematical structure and physical insight. It goes without saying that I am also grateful for his constant advices, for introducing me to quantum gravity (from black holes to string dualities), for the collaboration on most of my research, for carefully reading the manuscript of this thesis and for putting up with my moods and my naivety from time to time.

Special thanks are due to Laurent Houart, who has been indispensable for carrying on the research which led to this thesis, and who has been my main interlocutor on physics (and other) matters for the last two years. Thanks a lot also for the spectral reading of the manuscript.

It is also a pleasure to thank Paul Windey for the collaboration on one article, and all my colleagues at the Service de Physique Théorique, namely and in random order Marc Henneaux, Christiane Schomblond, Philippe Spindel, Robert Brout, Jean-Marie Frère, Glenn Barnich, André Wilch, Bernard Knaepen, Lubo Musongela, Abdelillah Barkallil, Marianne Rooman, Pietro Castoldi and Michel Tytgat, not to forget Peter van Driel and Ruud Siebelink with whom I had many discussion.

Particularly warm thanks go to Carmelo Iannuzzo and Karin Bautier with whom I shared the office at the beginning and at the end of these four years respectively. They have proven themselves skillful in enduring my changing moods.

I would also like to thank my (young) colleagues which I met during schools and conferences, and especially Edu Eyras and Dimi Matalliotakis for our virtual physics non-physics correspondence, and Walter Moretti.

Finally I think that an invaluable and possibly the strongest support has come from people outside the hep-th community, and since it would be too long and difficult to specialize this part of the acknowledgements, I will just say: thank all!!

Contents

1	Introduction	7
2	What is M-theory?	10
2.1	An historical perspective on string theory	11
2.2	Perturbative string theory	15
2.3	The philosophy of duality	21
2.4	String dualities and the appearance of M-theory	24
2.5	D-branes and their many uses	36
2.6	Matrix theory approach to M-theory	44
2.7	No conclusion	53
3	Classical p-brane solutions and their intersections	56
3.1	Why p -branes are interesting	57
3.1.1	The starting point: the supergravity actions	59
3.2	Singly charged extremal and black p -branes	63
3.2.1	The symmetries of the p -brane ansatz and the general equations of motion	64
3.2.2	The extremal dilatonic p -brane	69
3.2.3	The general non-extremal black p -brane	73
3.2.4	Semi-classical thermodynamics and the mass of the black p -branes	77
3.2.5	The branes of M-theory	84
3.3	Intersections of branes	89
3.3.1	Finding the classical solutions and the intersection rules	91
3.3.2	Further general considerations on intersecting brane solutions . .	98
3.3.3	Applications to M-theory	102
3.3.4	Supersymmetry and the branes	105
3.3.5	A non-supersymmetric, extremal solution	112
3.4	Conclusions and related work	116
4	The dynamics of open branes	122
4.1	A different light on some intersections	123
4.2	Charge conservation for the open branes	125
4.2.1	A D2-brane ending on a D4-brane	126
4.2.2	Charge conservation using the Chern-Simons terms	128

4.2.3	Electric-magnetic duality on the brane	130
4.3	Open branes and world-volume fields for the host brane	131
4.4	Two more cases of open branes	134
4.5	Conclusive speculations	136
5	Little theories on the world-volume of branes	138
5.1	Introduction	138
5.2	Theories with 16 supercharges	140
5.2.1	<i>iaa</i> theory	142
5.2.2	<i>ib</i> theory	145
5.2.3	<i>m</i> -theory	148
5.2.4	Some comments on the low-energy effective actions of the little theories	151
5.2.5	Some comments on the decoupling from the bulk	153
5.3	Theories with 8 supercharges	155
5.3.1	Type <i>i</i> , h_b and h_a theories	156
5.3.2	The low-energy effective actions and further comments	158
5.4	Application to Matrix theory compactification	159
5.5	Discussion	164
6	The four dimensional Schwarzschild black hole as an extremal brane configuration	166
6.1	Black hole entropy in M-theory	166
6.2	Mapping a neutral black brane on a configuration with four charges	167
6.3	Infinite boost limit, extremality, and statistical entropy	172
A	Kaluza-Klein reduction and the resulting action	175
B	Summary of the properties of the branes in 10 and 11 dimensions	179
C	Computation of curvature tensors	183
	Bibliography	189

Chapter 1

Introduction

M-theory is the name which designates the recent efforts to unify by non-perturbative dualities all the known superstring theories. It is a necessary step if string theories are to be related to a unique theory of all fundamental forces.

In the attempt to unify all the forces present in Nature, which entails having a consistent quantum theory of gravity, superstring theories seem promising candidates. They appear as theories achieving the unification ‘from above’. This is certainly appealing from a theoretical point of view, but it makes much more difficult the task to provide the concrete picture of how the unification occurs in the ‘real world’. The evidence that string theories could be unified theories is provided by the presence in their massless spectrum of enough particles to account for those present at low-energies, including the graviton. At the phenomenological level however, the difficulties associated with string theories show up when one tries to reduce the dimensionality of space-time from ten (the critical dimension for the superstrings) to four (the number of dimensions that are accessible to us, and to present-day accelerators), and when one also attempts to reduce the large amount of supersymmetry which characterizes the superstring spectrum.

However string theories have also theoretical drawbacks. First of all, string theories are properly speaking theories in which a single string is quantized perturbatively. There is no second quantized string theory, or string field theory. Our knowledge of string interactions, and string loop expansion, is entirely based on some sort of S-matrix theory. This means that we are able to compute scattering amplitudes between string states, but we are not able to formulate a path integral from which to derive such amplitudes. The inability to formulate a path integral is merely due to the inexistence of an action (or an Hamiltonian) describing string field theory. If we restrict our attention on the massless modes of the string, i.e. we focus on energies well below the scale fixed by the string tension, we can however use supersymmetry to argue that the low-energy effective action of these modes is given precisely by the corresponding supergravity. Such supergravities are non-renormalizable, but they are relevant essentially for the classical solutions that they contain, and which are to be identified with the solitons of the full string theory.

The second main drawback of superstring theories is their number. There are five string theories which all could pretend to be the unique theory. Some of them are more realistic than the others, but one would like to have a more theoretical reason to reduce

the number of consistent string theories. A different way to solve this problem is to show that all these five string theories are actually unified in a single theory.

The unification cannot however occur at the perturbative level, because it is precisely the perturbative analysis which singles out the five different string theories. The hope is that when one goes beyond this perturbative limit, and takes into account all non-perturbative effects, the five string theories turn out to be five different descriptions of the same physics. In this context, which is now called M-theory, a duality is a particular relation applying to string theories, which can map for instance the strong coupling region of a theory to the weak coupling region of the same theory or of another one, and vice-versa, thus being an intrinsically non-perturbative relation. In the recent years, the structure of M-theory has begun to be uncovered, with the essential tool provided by supersymmetry. Its most striking characteristic is that it indicates that space-time should be eleven dimensional.

Because of the intrinsic non-perturbative nature of any approach to M-theory, the study of the p -brane solitons, or more simply ‘branes’, is a natural step to take. The branes are extended objects present in M-theory or in string theories, generally associated to classical solutions of the respective supergravities. A 0-brane is simply a point-like object, a 1-brane is a string, a 2-brane a membrane, and so on. Leaving aside the fundamental strings which can be called abstrusely 1-branes, all the other branes are effectively non-perturbative objects, their mass (or tension) being proportional to an inverse power of the (string) coupling constant.

The study of the branes reveals itself very interesting and full of surprises. The branes first appear as solutions of the supergravities which are the low-energy effective theories of the superstring theories, and of eleven dimensional supergravity which is assumed to be the low-energy effective theory of M-theory. Most of these solutions were known before the dualities in string theory became the object of intensive research. However it is only recently that these objects have been treated as truly dynamical objects.

Already at the level of the classical supergravities, the analysis of solutions made up of several different branes gives interesting results. It is then possible to study some of these solitons directly in perturbative string theory, i.e. as D-branes. This gives an enormous amount of information on their dynamics, and some of these results can even be transposed by dualities to the other branes which do not benefit from such a ‘stringy’ formulation. Most of the branes turn out to have non-trivial low-energy effective theories defined on their world-volume, and these theories can be studied on their own. This latter property of the branes allows their use in addressing some very important problems. Indeed, the embedding in M-theory has led to serious advances in such problems as the black hole entropy and the study of non-perturbative aspects of field theories.

The aim of Chapter 2 is to give an introduction to M-theory, and to some of the objects that are relevant to its definition. Accordingly, the chapter contains a very brief introduction to (perturbative) string theory. The notion of duality and the various dualities in string theory and in M-theory are then reviewed. The chapter also contains a section on D-brane physics, and another on the Matrix theory approach to M-theory.

The chapters which follow contain material which is based on original results.

In Chapter 3 we concentrate on the classical p -brane solutions. The set up is actually more general than the one of a particular supergravity, mainly because for most of the time we only consider bosonic fields. Section 3.2 is devoted to reviewing and rederiving the known solutions carrying one antisymmetric tensor field charge. Both extreme and non-extreme (black) branes are considered, as well as a self-contained derivation of their semi-classical thermodynamics. In Section 3.3 we go on presenting the original results of [1]. General solutions consisting of several branes intersecting with definite pairwise rules, and forming extreme marginal bound states, are derived from the equations of motion. The intersection rules thus derived have a compact and very general form, which can be applied to all the cases of interest in M-theory. We then present an alternative and original way of deriving the intersection rules for the branes of eleven dimensional supergravity asking that the solution preserves some supersymmetries. We also derive a solution which is extreme but not supersymmetric, and thus turns out to be a different case from the previous ones.

In Chapter 4 we consider in detail the possibility that some of the intersection rules previously derived can reflect the property of some branes to open, with the boundaries attached to some other brane [2]. We thus check directly that the charge carried by the open brane is still conserved in the process. The mechanism by which the charge is conserved leads to the identification of the boundary of the open brane to a charged object on the world-volume of the brane which hosts it. Since a charged object couples to an antisymmetric tensor field, this procedure also fixes the field content of the world-volume effective theory of the host brane. All the relevant cases in M-theory are reviewed. This chapter and the previous one are also partially based on [3].

Chapter 5 is devoted to the study of the theories of extended objects that can be defined on the world-volume of some of the branes of M-theory, when they are decoupled from bulk effects [4]. These theories, which we call ‘little theories’, have extended objects corresponding to intersections between the host brane and some other brane. We define two different theories of little strings in six dimensions, which are obviously non-critical, and a theory in seven dimensions which possesses a membrane. Dualities can then be defined which relate all the above little theories. The low-energy effective theories of all the little theories are argued not to contain gravity. The definition of three more little string theories with fewer supersymmetry and additional group structure allows us to draw a parallel between these little theories and the ones in ten and eleven dimensions, based on the fact that the patterns of dualities are the same.

In Chapter 6 we study the entropy of the four dimensional Schwarzschild black hole in the context of M-theory [5]. By a repeated use of boosts along internal directions and of dualities, we map a neutral Schwarzschild black brane to a configuration carrying four charges. In the infinite boost limit, which coincides with extremality for the charged configuration, the entropy can be accounted for by a microscopical counting previously done in the literature, using D-brane techniques. We argue that the statistical interpretation of the entropy maps back to the Schwarzschild black hole.

Chapter 2

What is M-theory?

Rather than answering the question in the title of this chapter, we will try here to review how this question arose in the very last few years. The question above became a relevant one when it was realized that all superstring theories, and 11 dimensional supergravity, are actually related by an incredibly extended web of dualities, and can thus be actually unified in what goes now under the name of M-theory¹. The latter, which we will loosely refer to as a ‘theory’ even if it has not received a firm formulation yet, seems a very serious candidate for a unique theory which unifies all known interactions, since it resolves the problem posed by the existence of several consistent superstring theories.

Up to now, we are however very far from understanding how this unification concretely takes place. This is because the dualities which relate the known string theories are mostly intrinsically non-perturbative. Thus knowing what M-theory really is would be equivalent to having a complete mastery of any one of the string theories at all orders in the coupling expansion, including all non-perturbative effects. Moreover, the conjectured dualities should become manifest symmetries of the spectrum. It is worthless saying that the current research is very far from this point.

In this chapter, we begin by reviewing the history of string theory, which presents already many interesting aspects. This also betrays the fact that our current perception of M-theory is still intimately based on concepts pertaining to superstring theory. Accordingly, we then proceed to a very sketchy presentation of the five superstring theories, in their perturbative formulation. In section 2.3 the notion of duality is introduced, and the interplay with supersymmetry is emphasized. Then the history of the so-called ‘Second Superstring Revolution’ is reviewed, namely all the dualities relating the string theories are presented in the order of discovery (as far as a conjecture and its evidences can be discovered). Section 2.5 considers in more detail the D-branes, the interest in which stems from the fact that they are tractable in perturbative string theory. They turn out however to be crucial in confirming most of the duality conjectures, and moreover they also led to incredible advances in other interesting problems in theoretical

¹Our ignorance about the complete structure of M-theory already shows up dramatically in the fact that we still do not know exactly what M stands for. The extensive research about this particular problem, which requires notions of history, philology, and psychology, will not however be reviewed here.

physics. This is also briefly reviewed. In section 2.6, the Matrix theory attempt at a direct, non-perturbative formulation of M-theory is discussed. In order not to deceive the reader, we have to anticipate that the chapter will end with the same question with which it began.

All the material presented in this chapter is by no means original. This is intended to be a very quick introduction to the field of research, and it aims also at providing an historical situation to it. The presentation will necessarily be subjective, and the references are not intended to be exhaustive. The usual reference for (perturbative) superstring theory is Green, Schwarz and Witten [6], while a recent review which also discusses the dualities and M-theory is [7].

2.1 An historical perspective on string theory

The seed from which string theory grew up is commonly assumed to be the Veneziano amplitude, conjectured in 1968 in the context of dual models for hadrons.

At that time, there was no quantum theory for the strong interactions, but there was an experimental output revealing a much more complicated behaviour with respect to the recently formulated electro-weak interaction. Most notably, this intricacy showed up with a proliferation of particles (actually short-lived resonances) with increasing spin and mass. The regularity of the plot of the mass against the spin of these resonances (i.e. the Regge behaviour), and the conjecture of duality between the s - and the t -channels for the hadronic amplitudes, led Veneziano to propose his amplitude [8], which realizes explicitly both the duality and the Regge behaviour.

The most interesting aspect of the Veneziano amplitude is that it implies an infinity of poles of increasing mass and spin in both of the (dual) channels. This means that the underlying theory, whatever it is, must have a spectrum containing an infinite number of massive excitations (i.e. particles). This is a clear sign that the theory which produces such an amplitude is not a conventional quantum field theory.

Soon after Veneziano's proposition, a theory which reproduced the amplitude, and thus the infinite particle spectrum, was found in the form of the theory of a quantized string. This is certainly surprising; one may however consider the infinite number of particles as a hint to abandon locality, and the string as the most simple object which gives rise to an infinity of vibrations of higher and higher energy. In 1970, the Nambu-Goto action for the (bosonic) string was formulated [9], and string theory was officially born. It is amusing to see already at this stage the incredible series of theoretical bold conjectures from which string theory sprang up.

A year later, in 1971, the fermionic string was formulated by Ramond and by Neveu and Schwarz [10] in order to have a model that describes also space-time fermions. This implied the introduction of fermions on the world-sheet of the string that had to be related to the world-sheet bosons (the embedding coordinates) by a symmetry exchanging bosons and fermions, supersymmetry. These strings are accordingly called superstrings.

An important feature of string theories is that to consistently quantize them, the

dimensionality of space-time has to be fixed. This gives $D = 26$ for the bosonic string and $D = 10$ for the superstring. Moreover, the same procedure of quantization fixes completely the spectrum of particles. This spectrum has two very particular features. Leaving aside the infinite tower of massive string modes, we have a massless sector which always includes a spin 2 particle², and a tachyon (a particle with imaginary mass) is also necessarily present. It will turn out that one cannot get rid of the tachyon in the bosonic string, while one can consistently discard it in the superstring theories, as we will see shortly.

However, in the beginning of the 70s, there were two deathly blows for string theory in general: on the experimental side, the Veneziano amplitude did not quite fit to some new results in the region of high-energy scattering at fixed angle; on the theoretical side, quantum chromodynamics (QCD) was being formulated and revealed itself as a much more promising candidate for describing strong interactions.

String theory surprisingly survived to these blows, making a virtue out of its own problems. One of the most celebrated problems in the string theory attempt to describe strong interaction was the total lack of experimental evidence of a massless spin 2 particle. However rescaling the string parameter α' , measuring the inverse string tension, from the strong interaction scale to the Planck scale, i.e. making it much smaller, it may well be that the massless spin 2 particle is nothing else than the graviton. This idea was pushed forward by Scherk and Schwarz in 1974 [11]. At that moment, string theory ceased to be a candidate for a theory of hadrons, becoming instead a candidate for a theory describing, and unifying, all known interactions, including gravity.

That (super)string theories should describe all interactions is a consequence of the fact that if the $\alpha' \rightarrow 0$ limit is taken (also called the zero-slope limit), all the massive string modes decouple and we are left with the massless sector. This sector includes, besides the graviton, many other particles with lower spin, which could account (at least in some of the string theories) for all the known ‘light’ particles, matter and gauge bosons. Also, it is possible then to write a field theory of these ‘low-energy’ modes, and the Einstein gravity is correctly recovered.

An important step which led to the removal of all the remaining inconsistencies in superstring theories was undertaken in 1976 by Gliozzi, Scherk and Olive [12], who introduced the so-called GSO projection which at the same time removed the embarrassing tachyon and imposed space-time supersymmetry. This last property of superstring theories will turn out to be incredibly important for the present-day developments. It amounts to saying that, not only the world-sheet fields, but also the string spectrum in 10 dimensions obeys a symmetry relating bosons to fermions. It has to be said that contemporarily to the research in string theory, there had been in the 70s an extensive research on supersymmetry, and on its local generalization, supergravity. It suffices to say now that space-time (extended) supersymmetries severely constrain the form that an action can take. Also, a supergravity theory can be formulated in a space-time with at most 11 dimensions, a dimension more than for superstring theories.

At the beginning of the 80s, there were 3 superstring theories at one’s disposal: 2 the-

²For the open superstring, this particle only arises at one loop in string amplitudes.

ories of closed superstrings with extended $N = 2$ space-time supersymmetry (one chiral and one non-chiral), and one theory including open strings, with $N = 1$ supersymmetry and with a gauge group which was fixed to be $SO(32)$ for consistency. These theories were shown to be finite at one-loop, and there was the hope that the finiteness would remain at all loops. This was because all the massive string modes regulate in such a way the effective theories coming from the massless sector, that the ultra-violet behaviour was much softer than in ordinary field theories. These advances greatly benefited from the explicitly space-time supersymmetric formulation of the superstrings given by Green and Schwarz [13].

It is during 1985 that the ‘First Superstring Revolution’ takes place. A revival of interest in string theory takes places after the formulation during that year of two other kinds of superstrings, called the heterotic strings, by Gross, Harvey, Martinec and Rohm [14]. These closed string theories are built asymmetrically combining features of the toroidally compactified bosonic string, and of superstrings. The remaining $N = 1$ supersymmetry is enough to preserve all the nice features of the superstrings, while the bosonic sector allows the spectrum to fall into representations of a group, respectively $SO(32)$ and $E_8 \times E_8$ (hence the two heterotic string theories). The massless sector contains all the vectors in the adjoint representation of the group, which correctly translates into a low-energy effective theory consisting of $N = 1$ supergravity coupled to a super Yang-Mills (SYM) theory.

The interest raised by the heterotic strings is mainly because they are promising candidates to incorporate the Standard Model of strong and weak interactions, thus unifying most simply all the known interactions. The serious flaw of the previously known closed superstring theories was that in the low-energy sector, there was no chance to find, at the perturbative level, a non-abelian group. We will see that duality will predict, and evidence has now been found, that gauge symmetry enhancement nevertheless occurs in these theories, by non-perturbative effects. However at that time there was little hope to get, even after compactification, a viable non-abelian gauge group from the $N = 2$ superstrings. In order to fit the $SU(3) \times SU(2) \times U(1)$ gauge group of the Standard Model, it turned out that the $E_8 \times E_8$ group of one of the heterotic strings was the most suitable. This should also explain, besides maybe other more aesthetical reasons, the focus on heterotic strings more than, say, on the open-closed $SO(32)$ superstring.

Still, to make contact with phenomenology, one had eventually to compactify the space-time from 10 to 4 dimensions, and also possibly to reduce the supersymmetry from the unrealistic $N = 4$ (in 4 dimensions) to, say, the more workable $N = 1$ set up. These two problems could actually be addressed at the same time, and almost the day after the revolution the compactifications on Calabi-Yau manifolds [15] and on orbifolds [16] appeared as very promising tools. Accordingly, the end of the 80s saw a huge literature on string phenomenology and the hope that string theory was the sought for grand unified theory (we would think that ‘Theory Of Everything’ is too much of a metaphysical notion to be used in this context).

Nevertheless there were some serious conceptual problems left over, leaving aside the concrete technical complications of the theory. The main problem was that one had to choose first between the 5 consistent superstring theories, and secondly between the

way one compactifies, and it soon became clear that there were thousands of ways to do this. This means that the parameters one had to tune in order to recover the Standard Model physics were much more than the parameters of the Standard Model itself! This is clearly not an appealing feature for a theory that pretends to be unique. Also, even from a mathematical or aesthetical point of view, it was not satisfying that 4 superstring theories, though perfectly consistent, seemed to be ‘wasted’ by Nature. In this line of thought, 11 dimensional supergravity had the same uncertain status. These problems caused a pause in the enthusiasm towards string theory as a workable grand unified theory at the beginning of the 90s.

The date at which the ‘Second Superstring Revolution’ takes place can be situated at the beginning of 1995, when Witten showed in a very important paper [17] that most of the strong coupling limits of all the known string theories could be reformulated through dualities in terms of the weak coupling limits of some other string theory, or in terms of the 11 dimensional supergravity. With this paper M-theory was born. However it has to be stressed that the concept of duality that Witten uses was already introduced and used in string theory from the beginning of the 90s by some authors, in particular Sen [18] who studied and provided evidence for the S-duality of the heterotic string theory in 4 dimensions (the original conjecture was formulated in [19]) and by Hull and Townsend [20] who postulated the existence of U-duality, a key concept in Witten’s derivation.

The concept of duality (in the modern sense) already entered high energy physics when the electric-magnetic duality (which, being a strong-weak duality, can be realized only at the non-perturbative level) of 4 dimensional gauge theories was first conjectured by Montonen and Olive [21], and then shown to be more likely to happen in supersymmetric theories [22]. In string theory, it was already known in the late 80s that a duality, called T-duality, related some of the string theories. T-duality is visible at the perturbative level, and acts on the string theories when they are compactified on, say, a generic d -dimensional torus (see [23] for a review on T-duality). However the new, non-perturbative, dualities conjectured in 1994-1995 opened up the concrete possibility that all string theories were related to each other, and to 11 dimensional supergravity as well. It is this theory that unifies all string theories and is presumably 11 dimensional that we call here ‘M-theory’.

It is needless to say that since then, incredible support was gathered towards this conjecture, and the set up was refined to provide a very nice picture of the web of dualities, which will be discussed in Section 2.4. This same period saw also many other advances in the understanding of string theory (with the use of D-branes for instance), and an increasing interrelation between string theory (or, rather, M-theory) and other theoretical problems as black hole physics and supersymmetric gauge theories in various dimensions.

Another very recent attempt to address directly M-theory is the Matrix theory conjecture of Banks, Fischler, Shenker and Susskind [24], proposed in October 1996. This is however too much of the present day research to fit into this brief historical review of string theory, which thus stops here.

Let us end this section with a brief remark on the terminology. In the present section and in the following one, the expression ‘string theory’ refers strictly to its perturbative

formulation. In the subsequent sections, and in the more general considerations, ‘string theory’ will also more widely indicate both the perturbative and the non-perturbative aspects of the theory, which eventually lead to the concept of M-theory. Since there is clearly not a sharp frontier between the two interpretations of ‘string theory’, we hope it will be clear from the context which one is referred to.

2.2 Perturbative string theory

We review in this section the status of string theory before duality came into play. The main point will be to give an account of the massless particles in the perturbative string spectrum, and to show the low-energy effective actions that govern their dynamics. We will focus mainly on type II superstrings, since these are the ones which will be used in most of the following chapters. This section is essentially based on [6], where the derivation of all the results presented here should be found, unless otherwise quoted.

Let us start by writing the world-sheet action of the superstring. The easiest action one can write is one in a trivial (flat) background, in quadratic form, and with a world-sheet metric that has been gauged-fixed to a flat 2-dimensional Minkowski metric. The action for the string reads:

$$I_{WS} = -\frac{1}{4\pi\alpha'} \int d^2\sigma \eta_{\mu\nu} \left\{ \eta^{\alpha\beta} \partial_\alpha X^\mu \partial_\beta X^\nu + i\bar{\psi}^\mu \gamma^\alpha \partial_\alpha \psi^\nu \right\}, \quad (2.1)$$

where we have defined the following: $1/2\pi\alpha'$ is the string tension; σ^α (with $\alpha = 0, 1$) are the world-sheet coordinates; $X^\mu = X^\mu(\sigma)$ are the embedding coordinates, which are scalars from the world-sheet point of view; $\psi^\mu = \psi^\mu(\sigma)$ are world-sheet spinors carrying a vectorial index of the $SO(1, 9)$ space-time Lorentz group.

In 2 dimensions, one can impose at the same time the Majorana and the Weyl conditions (see for instance [25]). It is thus simplifying to choose a basis of γ matrices in which they are real. Here we take:

$$\gamma_0 = \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix}, \quad \gamma_1 = \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix}, \quad (2.2)$$

which verify $\{\gamma_\alpha, \gamma_\beta\} = 2\eta_{\alpha\beta}$. Since all the γ matrices are taken to be real, the matrix C such that $C\gamma_\alpha^T = -\gamma_\alpha C$ is simply $C = \gamma_0$. If the Majorana and the Dirac conjugate are defined respectively by $\bar{\psi}_M = \psi^T C$ and $\bar{\psi}_D = \psi^\dagger \gamma_0$, then the Majorana condition $\bar{\psi}_M = \bar{\psi}_D$ simply translates to a reality condition on the spinors, $\psi^* = \psi$. The matrix γ_2 can be defined by:

$$\gamma_2 = \gamma_0 \gamma_1 = \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix}, \quad (2.3)$$

and it correctly verifies $\gamma_2^2 = 1$, $\gamma_2^\dagger = \gamma_2$. Accordingly, a spinor ψ can be split into 2 chiral components, $\psi = \frac{1+\gamma_2}{2}\psi_R + \frac{1-\gamma_2}{2}\psi_L$; in two component notation, we have:

$$\psi = \begin{pmatrix} \psi_R \\ \psi_L \end{pmatrix}.$$

Note that the fermionic term of the action (2.1) can be rewritten as:

$$I_{WS} = \dots - \frac{i}{2\pi\alpha'} \int d^2\sigma \eta_{\mu\nu} \{ \psi_R^\mu \partial_+ \psi_R^\nu + \psi_L^\mu \partial_- \psi_L^\nu \}, \quad (2.4)$$

where we have defined $\sigma^\pm = \sigma^0 \pm \sigma^1$. The equations of motion thus imply that the fermions with positive and negative chirality are respectively right and left moving.

The action (2.1) has a very particular feature. Since both of the terms have the same overall coefficient, one can write the following symmetry relating the bosons and the fermions on the world-sheet:

$$\delta X^\mu = i\bar{\epsilon}\psi^\mu, \quad \delta\psi^\mu = \gamma^\alpha \partial_\alpha X^\mu \epsilon. \quad (2.5)$$

This is the simplest manifestation of supersymmetry. This symmetry can be seen to close (on-shell) to the momentum operator. Supersymmetry is particularly important to carry out properly the quantization of the superstring, and to fix consistently the dimensionality of space-time to $D = 10$.

When deriving the equations of motion from the action (2.1), one sees immediately that, both in the open and in the closed case, one can impose two different sets of boundary conditions on the world-sheet fermions ψ^μ . For the open string one has the choice between imposing at each boundary $\psi_R = \pm\psi_L$. Since the overall sign does not matter, the condition at one end is always fixed to have the positive sign. Then if at the other end there is also a positive sign, the fermionic fields ψ^μ will be integer moded. This is called the Ramond (R) sector. If on the contrary the second boundary condition has a negative sign in it, the ψ^μ will be half-integer moded. This is the so-called Neveu-Schwarz (NS) sector.

For a closed string, the discussion is the same, because we have now a choice in the periodicity condition for the fermions: $\psi_R^\mu(\sigma = 2\pi) = \pm\psi_R^\mu(\sigma = 0)$, and similarly for ψ_L^μ . Again, when choosing the + or the - sign, one is respectively in the R or in the NS sector. However now we have two independent choices to make. There will be thus four different sectors in the end, NS-NS, R-R, R-NS and NS-R.

In the R sector of, say, the open string, the vacuum energy is zero. The ground state has however to be a representation of the algebra of the zero modes of the fermionic fields ψ_0^μ , which is actually the Clifford algebra for the space-time Lorentz group. If we classify the states in terms of the representations of the little group $SO(8)$, we see that the ground state of the R sector can be chosen to be either in the $\mathbf{8}_+$ or in the $\mathbf{8}_-$ representation (which differ by their chirality).

On the other hand, the ground state of the NS sector has a negative vacuum energy of $-\frac{1}{2}$ (when the minimal gap for the bosonic fields is 1), and it is a singlet of $SO(8)$. This state is hence a tachyon. When acting on it with a $\psi_{-\frac{1}{2}}$ mode, a massless state is created with the indices of the $\mathbf{8}_v$ vectorial representation of $SO(8)$. The superstring represents space-time fermions and bosons when both sectors contribute to the particle spectrum.

Space-time supersymmetry, which requires the same number of on-shell bosonic and fermionic states at each mass level, is achieved, and the tachyon is consistently discarded,

when the GSO projection is performed. This projection has as a consequence that the chirality of the R sector ground state is fixed, and that in the NS sector, all the states created from the vacuum by an even number of fermionic operators ψ_{-r} are eliminated (a more precise definition of the GSO projection can be found e.g. in [6]). This in particular excludes the tachyon from the physical spectrum, and leaves us with a massless sector consisting of the states $\mathbf{8}_v + \mathbf{8}_+$.

The states in the massless sector realize 10 dimensional $N = 1$ chiral supersymmetry, and actually correspond to a gauge boson and to its fermionic superpartner (also called gaugino).

It is however too early to define a theory of open strings, since these ones necessarily couple to closed strings. Indeed, one-loop diagrams (e.g. the cylinder) for open strings can be cut in such a way that the internal propagating string is a closed string. Let us then review before the theories of closed strings.

The closed string theories are rather easy to build because the right moving and the left moving sectors decouple, and each one is quantized in exactly the same way as for the open string. The massless sector of a closed string theory is then a tensor product of two copies of the open string massless sector. There are however two different tensor products that one can make, depending on whether one takes the Ramond ground states of the opposite chirality on the two sides or of the same chirality.

If the right and left moving sectors are of opposite chiralities, the massless particles are classified along the following representations of $SO(8)$:

$$\begin{aligned}
 (\mathbf{8}_v + \mathbf{8}_+) \otimes (\mathbf{8}_v + \mathbf{8}_-) &= (\mathbf{1} + \mathbf{28} + \mathbf{35}_v)_{NSNS} + (\mathbf{8}_v + \mathbf{56}_v)_{RR} \\
 &+ (\mathbf{8}_+ + \mathbf{56}_-)_{NSR} + (\mathbf{8}_- + \mathbf{56}_+)_{RNS}.
 \end{aligned}
 \tag{2.6}$$

The bosonic particles of the NSNS sector are respectively a scalar, a two-index anti-symmetric tensor and a two-index symmetric traceless tensor; they are related to the dilaton field ϕ , the antisymmetric tensor field $B_{\mu\nu}$ and the metric $g_{\mu\nu}$ associated with the graviton. In the RR sector we find a vector and a 3-index antisymmetric tensor, which correspond to the fields A_μ and $A_{\mu\nu\rho}$. The space-time fermions are found in the RNS and NSR sectors, and the $\mathbf{56}_\pm$ representations are the two gravitini; this is thus a theory with $N = 2$ supersymmetry. Moreover, since the gravitini have opposite chiralities, the theory is non-chiral. This is usually called the type IIA string theory.

To write a low-energy effective action for the massless fields above is made easier by the constraint that we should be describing a space-time supersymmetric theory. Indeed, the action of 10 dimensional type IIA supergravity is completely fixed. Alternatively, one can compute the interactions between the massless fields from string theory amplitudes, and then extrapolate to the non-linear action that reproduces them. Still another successful way to obtain the effective action is to write the string world-sheet action (2.1) in curved background and then to ask that conformal invariance is not broken quantum mechanically, i.e. the β -function must vanish.³ One obtains the following action (we

³ Note that a proper definition of the low-energy effective action should actually only come from a string field theory perspective, which is still lacking.

only write its bosonic sector):

$$I_{IIA} = \frac{1}{16\pi G_N} \left[\int d^{10}x \sqrt{-g} \left\{ e^{-2\phi} \left(R + 4(\partial\phi)^2 - \frac{1}{12}H_3^2 \right) - \frac{1}{4}F_2^2 - \frac{1}{48}F_4'^2 \right\} + \frac{1}{2} \int B_2 \wedge F_4 \wedge F_4 \right]. \quad (2.7)$$

We have defined the field strengths in the following way: $H_3 = dB_2$, $F_2 = dA_1$, $F_4 = dA_3$ and $F_4' = F_4 + A_1 \wedge H_3$. The metric in the action above is the string metric, which appears in the string action in curved background $I_{curved} \sim \int g_{\mu\nu} \partial_\alpha X^\mu \partial^\alpha X^\nu$. Note that the fields coming from the NSNS sector have a $e^{-2\phi}$ factor in front of their kinetic term, while the RR fields not. This is discussed in [17, 26]. The only arbitrary dimensionful parameter of the type IIA string theory is α' , and accordingly the Newton constant appearing in (2.7) is expressed in terms of it, $G_N \sim \alpha'^4$. The string coupling constant, which governs the expansion in string loops, is actually the vacuum expectation value of the dilaton, $g_s = e^{\langle\phi\rangle}$. It is thus determined dynamically. The effective Newton constant will be affected by the asymptotic values of such moduli. Note that the action (2.7) is the leading term in both the α' and the g_s expansion.

An important remark on the closed string spectrum has to be made now. While there are perturbative string states carrying a charge with respect to the NSNS fields (i.e. the winding states and the Kaluza-Klein momentum states when there is a compact direction), there are no states carrying the charge of a RR field. This problem has a technical origin, and has been a long standing puzzle in perturbative string theory. We now know that the D-branes, which can be considered as non-perturbative objects, are the carriers of these charges.

We can now consider the second closed superstring theory, in which both left and right moving sectors have the same chirality. The massless particles are:

$$(\mathbf{8}_v + \mathbf{8}_+) \otimes (\mathbf{8}_v + \mathbf{8}_+) = (\mathbf{1} + \mathbf{28} + \mathbf{35}_v)_{NSNS} + (\mathbf{1} + \mathbf{28} + \mathbf{35}_+)_{RR} + (\mathbf{8}_- + \mathbf{56}_+)_{NSR} + (\mathbf{8}_- + \mathbf{56}_+)_{RNS}. \quad (2.8)$$

The NSNS sector is exactly the same as the one of the type IIA superstring. The RR sector on the other hand is different; this time we have a scalar (or 0-form), a 2-form and a 4-form potential. However for the 4-form potential to fit into the $\mathbf{35}_+$ representation, its field strength has to be self-dual. The fields are respectively χ , $A_{\mu\nu}$ and $A_{\mu\nu\rho\sigma}^+$. The space-time fermions of the mixed sectors now comprise two gravitini of the same chirality. We have thus a chiral $N = 2$ supersymmetry, and the theory goes under the name of type IIB string theory. Note that from the superalgebra point of view, one often refers to this theory as a (2,0) theory, while the type IIA theory is a (1,1) theory. This has as a consequence that the fields of the type IIB theory can be arranged into representations of $SO(2) \equiv U(1)$, i.e. some fields can be paired in complex fields.

Because of the presence of the self-dual 5-form field strength, the type IIB supergravity was originally formulated only in terms of its equations of motion, which are completely fixed by the (2,0) supersymmetry. We can however write an action for this theory if we do not ask that the self-duality condition follows from the variation of the

action, but is rather imposed afterwards. The bosonic part of the low-energy effective action of the type IIB superstring thus reads:

$$I_{IIB} = \frac{1}{16\pi G_N} \left[\int d^{10}x \sqrt{-g} \left\{ e^{-2\phi} \left(R + 4(\partial\phi)^2 - \frac{1}{12}H_3^2 \right) - \frac{1}{2}(\partial\chi)^2 - \frac{1}{12}F_3'^2 - \frac{1}{240}F_5'^2 \right\} + \int A_4 \wedge F_3 \wedge H_3 \right]. \quad (2.9)$$

Here the RR field strengths are defined by $F_3' = F_3 - \chi H_3$, $F_3 = dA_2$ and $F_5' = dA_4 + A_2 \wedge H_3$, and we impose that $F_5' = *F_5'$. The same remarks about the type IIA supergravity action (2.7) apply to this case.

The two actions above (2.7) and (2.9) are the actions of the only two maximal supergravities in 10 dimensions. Maximal supergravities exist because if one tries to extend further the supersymmetry, then one has more than one graviton in the spectrum, and also massless fields with spin higher than 2. This is clearly not physical, since we want gravity to be unique. The total amount of supercharges is thus fixed to be at most 32. This also constrains the dimension of the space-time in which one wants to build the supergravity, since the number of components of the smallest spinor representation increases exponentially with D (see e.g. [25]). It turns out that the highest dimensional supergravity one can formulate is in 11 dimensions. This supergravity is apparently not related to the superstrings, which live in 10 dimensions. Note however that the Kaluza-Klein reduction of the 11 dimensional supergravity yields exactly the action of the type IIA supergravity (2.7) (see Appendix A). This relation is now understood to have a very deep meaning. We will come back to it in Section 2.4.

We can now go back to discuss (briefly) the open superstring. We had already stated that the massless modes of the open strings have $N = 1$ supersymmetry, and that they must be coupled to a closed string sector. However the type II closed strings discussed above cannot be coupled consistently to the open strings, since they have $N = 2$ supersymmetry. One way to divide by two the supersymmetry of the closed theories is the following: one can identify the left and the right moving sectors of the type IIB superstring, thus making a theory of closed unoriented strings. Only half of the massless states in (2.8) survive, and these are the following:

$$\frac{(\mathbf{8}_v + \mathbf{8}_+) \otimes (\mathbf{8}_v + \mathbf{8}_+)}{\Omega} = (\mathbf{1} + \mathbf{35}_v)_{NSNS} + (\mathbf{28})_{RR} + (\mathbf{8}_- + \mathbf{56}_+)_{RNS-NSR}, \quad (2.10)$$

where we have called Ω the orientation reversal operator.

The open strings have however a richer structure than the closed ones, since they can carry charges at their ends. These are called Chan-Paton factors. If the factors are taken to be in the fundamental representation of, say, a $U(n)$ group, it follows that an open string state generically belongs to the adjoint representation of that group. If the string is unoriented, the group becomes either $SO(n)$ or $Sp(n)$ (see [26] for an exhaustive discussion). It turns out that the only anomaly-free theory is the one with the group $SO(32)$. This is consistent with the fact that the closed sector is unoriented as well.

The low-energy effective action of the $N = 1$ open-closed superstrings, also called type I superstrings, is then the action of $N = 1$ supergravity in 10 dimensions coupled

to a super Yang-Mills theory with gauge group $SO(32)$. We will not write the action here, however it is worth noting two features of the type I theory: first of all, as the type II theories, it has only one free parameter, which is α' . Secondly, the bosonic part of the action has the same structure as the type II ones, with a $e^{-2\phi}$ factor in front of the NSNS fields, and nothing in front of the RR one. Here however the kinetic term of the fields coming from the open sector has a $e^{-\phi}$ factor in front of it, and this can be traced back to the fact that this part of the action comes from the amplitude computed on the disk instead of the sphere.

Let us conclude with a very quick review of the heterotic string theories. The key idea in constructing these theories is to remark that in closed string theories the quantization of the right and left moving sectors can be carried out independently. For both sectors, the quantization is consistent if the central charge is either 15 if there is supersymmetry or 26 if there is no such symmetry. Knowing that each boson and each fermion contributes respectively 1 and $\frac{1}{2}$ to the central charge, we reach the conclusion that D must be 26 for the bosonic string, and 10 for the superstring. One can however consider that supersymmetry is present only in, say, the right moving sector. This amounts to take only the first term in the action (2.4). One directly sees that the world-sheet supersymmetry is realized in this case with a parameter ϵ of negative chirality, hence there is one half of the supersymmetry of the type II closed superstrings. This fixes already to 10 the dimensions of the target space-time.

In the left moving sector, we have now the freedom to choose the fields which will make up the $c = 26$ central charge. Ten of these fields are already fixed to be the (left moving part of the) embedding coordinates. Then the most natural choice for the other fields is to take 32 (anti)chiral fermions. One can easily guess that these fermions have an $SO(32)$ global symmetry. Carrying out the quantization of the left moving sector, one finds that the massless level contains, besides the vector $\mathbf{8}_v$ of the space-time $SO(8)$, also 496 states in the adjoint of $SO(32)$ or $E_8 \times E_8$ depending on which periodicity conditions have been taken for the fermions. The ground state of the left moving sector is actually tachyonic, but the level matching condition with the right moving sector excludes it from the physical sector. The supersymmetry of only one of the two sectors is thus enough to ensure a tachyon-free spectrum to the heterotic strings.

An alternative way to formulate the heterotic string is to take all the left moving fields to be bosons. However 16 of them must be compactified on a 16-dimensional torus T^{16} . Since these bosons are chiral, they satisfy slightly modified commutation relations, and their quantized momentum is fixed in terms of the winding number. Moreover, in order for them to contribute to the massless spectrum (a requirement for one-loop unitarity) the T^{16} has to be chosen very particularly, i.e. it must be defined by a self-dual lattice. This results in the same two possible spectra as the ones discussed in the previous paragraph. In the latter construction, however, we have the picture of the heterotic string being a hybrid of the superstring and of the bosonic string.

We can now recapitulate the massless spectrum of the heterotic strings. Let us first focus on the tensorial product of the right moving sector with the left moving space-time vector. We recover the same particles as in the closed sector of the type I string, though

they are obtained differently:

$$(\mathbf{8}_v + \mathbf{8}_+)_R \otimes (\mathbf{8}_v)_L = (\mathbf{1} + \mathbf{28} + \mathbf{35}_v)_B + (\mathbf{8}_- + \mathbf{56}_+)_F. \quad (2.11)$$

All the particles in the r.h.s. are singlets of the $SO(32)$ or $E_8 \times E_8$ group. The remaining left moving massless states are in the adjoint of one of the two groups above, and are $SO(8)$ scalars. When we take the tensorial product of these states with the right moving sector, we recover nothing else than the states which make up the $N = 1$ vector multiplet of an $SO(32)$ or $E_8 \times E_8$ gauge theory.

The massless particle content of the heterotic string is thus exactly the same as the one of the type I superstring⁴. There are however two versions of the heterotic string, since the construction is more ‘flexible’ and allows also for the $E_8 \times E_8$ group.

The low-energy effective action for the heterotic strings is also the one of $N = 1$ supergravity coupled to super Yang-Mills, but it differs from the low-energy effective action of the type I superstring in that the whole bosonic part of the action has now a factor of $e^{-2\phi}$ in front of it. Duality will eventually relate the $SO(32)$ heterotic string to the type I string.

This concludes this short review of all the five consistent supersymmetric string theories. We will now see how all of these string theories are related by perturbative and non-perturbative dualities, and are unified into M-theory, which is effectively an 11 dimensional theory.

2.3 The philosophy of duality

It is now time to introduce in this section the concept of duality, which is really the core of the so-called Second Superstring Revolution⁵. Roughly speaking, there is a duality when two seemingly different theories are actually physically the same. Since showing that two theories produce exactly the same physics is a tremendous task in most cases, duality has often the status of a conjecture.

To be more precise, two theories are related by a duality if there is a one-to-one map between their physical spectra, and if their dynamics are equivalent too. From the physical, on-shell, point of view, there should be no difference between the two theories. The latter are thus two different descriptions of the same physics. The two descriptions can be quite different as we will see. Note also that the two theories can also coincide with the same classical theory but in two different regimes, e.g. with weak and strong coupling interchanged.

What makes difficult to actually prove duality is that often a duality conjecture involves a map between a weak and a strongly coupled theory, and thus only the perfect knowledge of non-perturbative effects could firmly establish the existence of a duality.

⁴Starting from the first massive level, the spectra of the heterotic string and of the type I string strongly differ. This also implies that the α' corrections to the low-energy effective actions will be different.

⁵Amusingly, string theory originated from a conjecture also named duality, as explained in Section 2.1. This was however a duality of another kind with respect to the concept discussed in this section.

Until now, this has been possible only for the two-dimensional model proposed by Coleman [27], which is the first and unique case of a proof of a non-perturbative duality.

A duality should thus be established properly only by solving completely the two theories. However the duality conjecture, if supported by some evidence, can be used as a working hypothesis to focus alternatively on only one of the two theories, and to consider them as completely equivalent. Results in one theory can be derived in the dual one, where they can be more accessible. This is the power of duality.

Having seen how duality acts on two different theories, let us see the particular, and most interesting case, in which duality relates two different regimes of the same theory. The archetype of this duality is the simple observation that the Maxwell equations are invariant under electric-magnetic duality. This has been re-inserted in the context of modern quantum field theory by Montonen and Olive in 1977 [21]. That the electric-magnetic duality is a non-perturbative duality can be easily seen considering that it maps particles of the perturbative spectrum, such as the bosons of the broken gauge symmetry, to magnetic monopoles, which are classical, non-perturbative solutions (for this map to be truly consistent, one has to settle in the framework of $N = 4$ supersymmetric Yang-Mills (SYM) theory [22]). Moreover, under this map the coupling constant g is taken to its inverse, $g \rightarrow 1/g$. As it will also happen in string theories, one can show that the combined possibilities to have dyonic objects (carrying electric and magnetic charge), axionic θ -terms in the action, and the Dirac quantization of the electric charge altogether imply that the duality group is actually $SL(2, Z)$. We will review in more detail in the next section how this duality precisely acts.

What kind of symmetry is this $SL(2, Z)$ duality? It states that a multitude of states in the spectrum are actually the same, but it does not require that any physical state should be invariant under it. It can thus be seen as a broken discrete gauge symmetry [28]. One has to mod out the spectrum by the $SL(2, Z)$ symmetry before, say, counting the states. Also, the moduli space of all the theories which differ by the values of some (dynamical) parameters, generically has to be modded out by the duality group.

The resulting theory is however non-trivial. Let us describe simply how duality works in the case of $N = 4$ SYM. The usual formulation of this theory is in terms of its coupling g which becomes the electric coupling when the gauge group is broken down to a $U(1)$. We will call the W bosons particles with $(1, 0)$ charge, where a particle with (p, q) charge carries p units of electric charge and q units of magnetic charge. Quantization of this theory is usually performed assuming that g is small, $g \ll 1$. There are however objects in this theory which cannot be reliably treated in perturbation theory since they have a mass going like $1/g$. These are the 't Hooft-Polyakov monopoles, or particles with $(0, 1)$ charge. What the duality conjecture states is that the theory can be reformulated taking the monopole to be the fundamental object. This new theory is exactly the same SYM theory, with however a new coupling \tilde{g} given by $\tilde{g} = 1/g$. Furthermore, the configurations with one single W boson at weak coupling or with one single monopole at strong coupling are strictly equivalent. Of course, each one of these two configuration is not duality invariant, rather it breaks the $SL(2, Z)$ duality symmetry. Since they are mapped one to the other, they belong to the same duality orbit.

Consider now the object with charge $(1, 1)$, which is called a dyon. It belongs also

to the same duality orbit as the W boson and the monopole, and can thus be also considered as the starting point for another perturbation theory, which will be the same SYM theory. Note however that this dyonic object is a bound state, and that it differs from a configuration in which both a W boson and a monopole are present. The latter configuration would have an energy of the order $E \sim M_W + M_{mon}$, while the dyon has a mass going like $M_{dy} \sim \sqrt{M_W^2 + M_{mon}^2}$. This ensures that the theory is still non-trivial, because the situation in which a W boson is scattering off a magnetic monopole is definitely not equivalent to a single W boson elementary state. The scattering state described here is a configuration in which the knowledge about the non-perturbative effects introduced by the monopole is still indispensable. In this case, duality only allows us to reformulate the problem as, say, the scattering of two dyons with opposite magnetic charge.

The benefit of duality in this case of $N = 4$ SYM theory, is that we no longer need to develop tools to understand the behaviour at $g > 1$, i.e. at strong coupling, since we can always reformulate the problem in a set up in which the coupling constant is smaller than 1.

The picture above can be roughly transposed to all the other dualities, which hold in string theory. Sometimes however, as we will see, only some limits of the dual theories are known and can thus be described.

The non-perturbative nature of the duality that we have just described makes it very difficult to prove, since this would involve knowing the behaviour of the strong coupling dynamics. What we are left to do to establish and support a duality conjecture is to provide evidence for it. It turns out that the most powerful tool to provide such evidence is supersymmetry.

When one considers a supersymmetry algebra with central charges, it appears that several kinds of supermultiplets exist (see e.g. [29]). A massless, or lightlike, supermultiplet is the shortest one, while a generical massive multiplet has a dimension which is the square of the dimension of the massless multiplet. Extended supersymmetry however allows for the presence of central charges in the superalgebra, and these in turn imply the existence of massive states which nevertheless fit into short multiplets. This is so because these massive states still preserve some of the supersymmetries. The dimension of the multiplet depends on the amount of preserved supersymmetry, but it is always smaller than the dimension of the generic massive multiplet. The shortest massive multiplet has the same dimension of the massless multiplet, and preserves half of the supersymmetries.

For a massive state to preserve some supersymmetry, and thus to fit into a short multiplet, it has to verify a constraint on its mass, which has to be equal to some of the central charges. Short massive multiplets are thus associated with Bogomol'nyi-Prasad-Sommerfield (BPS) [30] states, which generically saturate a bound on their mass with respect to their charge(s). These states are thus often called BPS states, even if the original definition of the latter did not refer to any supersymmetry property.

The interest of the BPS states resides in the following argument, first given by Witten and Olive in 1978 [31]. Since the dimension of a multiplet cannot change when a continuous parameter, such as the coupling constant, is varied, a BPS state stays so at any value of the coupling. Now, the condition for a state to be BPS is that the mass

equals its charge. We have thus the certitude that even at strong coupling, where we have no handle on the dynamics, a BPS state of a given charge will have a fixed mass. In other words, this argument tells us that the mass-charge relation of a BPS object does not suffer any quantum correction. Moreover, in theories with enough supersymmetry also the continuous parameters are not renormalized, and this allows us to predict which BPS states become light at strong coupling.

The utility of BPS states is thus the following. If one restricts the evidence for a duality to the spectrum of (partially) supersymmetric states, one does not need any more to compute difficult strong coupling effects, but instead one can simply analyze the properties of the BPS states at weak coupling (or, at any rate, in the region where the computations are possible), and then extrapolate at strong coupling. This is in some way a second conjecture to support a first conjecture, but it has proven very efficient and indeed exact in some cases.

We have defined above the set of ideas that characterize what is meant nowadays by the word ‘duality’. Let us now emphasize what are the consequences of a conjectured duality between two theories.

Once one has acquired evidence in support of a duality, one can then start using the relation between the two theories (or between the strong and the weak coupling regimes of the same theory) to infer important and otherwise inaccessible results. In the case of dualities relating the strong coupling of one theory to the weak coupling of another, this gives a handle on the non-perturbative effects and on the strong coupling dynamics of the first theory simply studying the perturbative effects of the second theory. It is by the exclusive use of this tool that it has been possible to gain such an enormous insight into the non-perturbative aspects of string theory in the last few years.

We review in the next section the web of dualities relating all the known string theories and 11 dimensional supergravity.

2.4 String dualities and the appearance of M-theory

Historically, duality and related ideas already appeared in string theory before the revolution of 1994-1995. The $SL(2, Z)$ duality of field theory was transposed to the heterotic string (toroidally) compactified to 4 dimensions. This duality, which was basically an electric-magnetic duality, was conjectured in [19] and then thoroughly studied by Sen [18] who provided concrete evidence. Still in the context of the heterotic string, which attracted much attention after the First Superstring Revolution of the mid eighties, the perturbative duality known as target space duality, or T-duality, was established by Narain [32] and by Ginsparg [33] who showed that the two heterotic string theories are equivalent when compactified on a circle. Application of T-duality to type II string theories, which we will review hereafter, was carried out already in 1989 in [34, 35].

Supergravity theories were also intensely studied ever since their formulation. In particular, the study of the maximal supergravities uncovered a large symmetry group of the latter, as pioneered by Cremmer and Julia [36] for $N = 8$ supergravity in 4 dimensions (see also [37] for other cases). It is the confluence of these ideas which

produced the Second Superstring Revolution.

T-duality

Let us begin with the review of T-duality for type II theories (useful references are [23, 26]). T-duality is the most accessible duality, since it is visible at the perturbative level. Below we consider its most simple manifestation, namely the exchange of large and small radii.

In order to see T-duality at work, we have to go back to perturbative string theory and write the mode expansion of the bosonic fields X^μ . Their equations of motion imply that they split into left and right moving parts:

$$X^\mu = X_L^\mu(\sigma^+) + X_R^\mu(\sigma^-), \quad \sigma^\pm = \sigma^0 \pm \sigma^1 \equiv \tau \pm \sigma. \quad (2.12)$$

Taking into account that $\sigma \sim \sigma + 2\pi$ and that X^μ will have to obey to some periodicity condition to be defined shortly, the mode expansions for the closed string are the following:

$$\begin{aligned} X_L^\mu &= x_L^\mu + \alpha' p_L^\mu \sigma^+ + i\sqrt{\alpha'} \sum_{n \neq 0} \frac{1}{n} \alpha_n^\mu e^{-in\sigma^+}, \\ X_R^\mu &= x_R^\mu + \alpha' p_R^\mu \sigma^- + i\sqrt{\alpha'} \sum_{n \neq 0} \frac{1}{n} \tilde{\alpha}_n^\mu e^{-in\sigma^-}, \end{aligned} \quad (2.13)$$

where $x_{L,R}^\mu$ and $p_{L,R}^\mu$ are real and $(\alpha_n^\mu)^\dagger = \alpha_{-n}^\mu$, and similarly for $\tilde{\alpha}_n^\mu$. Note that the split of the center of mass $x_0^\mu = x_L^\mu + x_R^\mu$ is arbitrary and thus unphysical, while the relation between p_L^μ and p_R^μ will be fixed by the periodicity condition. The total momentum carried by the string is $P^\mu = p_L^\mu + p_R^\mu$.

Let us now analyze the periodicity conditions on X^μ . If the closed string is propagating in flat non-compact spacetime, then the only choice we have is to impose:

$$X^\mu(\tau, \sigma + 2\pi) = X^\mu(\tau, \sigma).$$

This implies that the sum (2.12) cannot depend linearly on σ , and thus fixes $p_L^\mu = p_R^\mu \equiv \frac{1}{2}P^\mu$.

Suppose now that the space is still flat, but one of the directions (e.g. X^9) is compact, with radius R . Space-time has now a topology of $R^9 \times S^1$, and the closed string can wind around this non-trivial 1-cycle. The periodicity condition for the compact coordinate can then be generalized to:

$$X^9(\tau, \sigma + 2\pi) = X^9(\tau, \sigma) + 2\pi w R. \quad (2.14)$$

Here the number w is an integer, and it is clear that it arises as a classical topological number. Quantum mechanically, the compact direction has another consequence: the momentum in this direction can no longer take any value, but is restricted to a set of discrete values in order for the wave function $e^{iP^9 X^9}$ to be single valued. We thus have $P^9 = \frac{n}{R}$, with n an integer (and a truly quantum number).

The center of mass and zero modes of X^9 are thus:

$$X^9 = x_0^9 + \alpha' \frac{n}{R} \tau + wR\sigma + \text{oscill. terms.}$$

This directly implies that:

$$\begin{aligned} p_L^9 &= \frac{1}{2} \left(\frac{n}{R} + \frac{wR}{\alpha'} \right), \\ p_R^9 &= \frac{1}{2} \left(\frac{n}{R} - \frac{wR}{\alpha'} \right). \end{aligned} \quad (2.15)$$

The classical, bosonic hamiltonian is the following:

$$H = \frac{\alpha'}{2} (p_L + p_R)^2 + \frac{\alpha'}{2} (p_L - p_R)^2 + \sum_{n \neq 0} (\alpha_{-n} \cdot \alpha_n + \tilde{\alpha}_{-n} \cdot \tilde{\alpha}_n).$$

All the products in the above expression are sums over the 10 space-time indices. We neglect the fermionic terms and the (possible) quantum ground state energy since they are not relevant to the present discussion.

Inserting the expressions (2.15) and the 9 dimensional non-compact space-time momentum P^μ , we get:

$$H = \frac{\alpha'}{2} P^2 + \frac{1}{2} n^2 \frac{\alpha'}{R^2} + \frac{1}{2} w^2 \frac{R^2}{\alpha'} + \text{oscill. terms.} \quad (2.16)$$

Using that $P^2 = -M^2$ and that the Hamiltonian must vanish on-shell, one can correctly rederive that the mass of an unexcited wound string is its length divided by its tension.

We have now the full set-up to uncover T-duality. The expression used to derive the masses of the perturbative string spectrum (2.16) is obviously invariant under the combined exchange:

$$\begin{aligned} R &\leftrightarrow \frac{\alpha'}{R}, \\ n &\leftrightarrow w. \end{aligned} \quad (2.17)$$

This is a very interesting symmetry in several respects. Although it is visible perturbatively in string theory, it has some striking features which have to be attributed to its stringy nature. The first consequence of (2.17) is that there seems to be a minimum length in string theory, since whenever a string tries to probe a distance smaller than $l_s \equiv \sqrt{\alpha'}$, then we know that we could instead reformulate the problem as a (possibly different) string probing a length scale larger than l_s (note however that there is now evidence that non-perturbative objects as D-particles can probe smaller distances). Furthermore, the second line of (2.17) indicates that quantum numbers are exchanged with classical, topological ones. We see already at this stage that string theory is a natural framework for the duality symmetries to appear. Indeed, T-duality is possible because of the existence of winding modes, which are absent in an ordinary Kaluza-Klein theory.

In terms of the left and right moving momenta (2.15), the T-duality transformation (2.17) becomes:

$$p_L^9 \leftrightarrow p_L^9, \quad p_R^9 \leftrightarrow -p_R^9.$$

Transforming also $\tilde{\alpha}_n \leftrightarrow -\tilde{\alpha}_n$ (which does not change the spectrum), the T-duality can be rewritten in the following way:

$$X^9 = X_L^9 + X_R^9 \quad \leftrightarrow \quad X'^9 = X_L^9 - X_R^9. \quad (2.18)$$

We can now see that because of the world-sheet supersymmetry (2.5), the fermionic superpartner ψ^9 also has to transform under T-duality, as $\psi_L^9 \leftrightarrow \psi_L^9$ and $\psi_R^9 \leftrightarrow -\psi_R^9$.

The transformation (2.18) acts like a space-time parity reversal restricted to the right moving modes. This is an important remark because it implies that the chirality of the right-moving Ramond ground state also changes, and this in turn means that T-duality maps type IIA superstring theory to type IIB and vice-versa.

We have thus established that type IIA and type IIB theories compactified on a circle are equivalent at the perturbative level. The equivalence has of course to be verified also at the non-perturbative level, and indeed it will be possible to check that the map extends also to all the non-perturbative objects of the two theories.

At the level of the low-energy effective actions, there is only one supergravity in 9 dimensions and thus it is obvious that the reduction of the two 10 dimensional supergravities is equivalent. The concrete map between the fields of IIA and IIB supergravities (with an isometry) has been shown in [38].

A last but important remark on T-duality is that the string coupling constant is actually not inert under it. This can be seen as follows. For the physics in 9 dimensions to be invariant under T-duality (this is required by the notion of duality discussed in the previous section), the Newton constant (which is the only independent coupling constant in the low-energy physics) has also to be invariant. Now the effective 9 dimensional Newton constant derived from, say, (2.7), is given by:

$$G_9 \sim \frac{g_A^2 \alpha'^4}{R_A},$$

where the subscripts indicate that we settle in type IIA theory. The constants g_A and R_A are simply the asymptotic values of, respectively, the fields e^ϕ and $(g_{99})^{1/2}$, and the way they fit into G_9 is dictated by the form of the action (2.7). We now require that the same G_9 is given by the IIB theory compactified on a circle of radius $R_B \equiv \frac{\alpha'}{R_A}$. Since α' cannot change in this problem (the string scale does not change), we have that:

$$g_B = g_A \frac{\sqrt{\alpha'}}{R_A}. \quad (2.19)$$

The same is of course true if IIA and IIB are interchanged. This formula will turn out to be very useful in determining how the non-perturbative objects transform under T-duality.

S-duality

We now turn to analyze a symmetry which is particular to type IIB supergravity in 10 dimensions, and that can have far-reaching consequences if extended to the whole type IIB string theory.

As it can be seen in the massless spectrum (2.8), the scalars and the 2-form potentials come into pairs. Due to the particular rôle played by the dilaton in the action (2.9), it can be seen that the equations of motion of type IIB supergravity are invariant under a $SL(2, R)$ symmetry [39]. The fields transform as follows. If we arrange the RR scalar χ and the dilaton in a complex scalar $\lambda = \chi + ie^{-\phi}$, then the $SL(2, R)$ symmetry acts like:

$$\lambda \rightarrow \frac{a\lambda + b}{c\lambda + d}, \quad (2.20)$$

where the real parameters are such that $ad - bc = 1$. If B_2 and A_2 are respectively the NSNS and the RR 2-form potentials, then they also transform according to:

$$\begin{pmatrix} B_2 \\ A_2 \end{pmatrix} \rightarrow \begin{pmatrix} d & -c \\ -b & a \end{pmatrix} \begin{pmatrix} B_2 \\ A_2 \end{pmatrix} = \begin{pmatrix} dB_2 - cA_2 \\ aA_2 - bB_2 \end{pmatrix}. \quad (2.21)$$

Since the NSNS field B_2 couples to the fundamental string and the corresponding charge is the winding number, which is quantized, the symmetry group is broken to its discrete subgroup $SL(2, Z)$ when one asks that the charges also transform in a way similar to (2.21). The quantization of the charges can also be seen to arise in 10 non-compact dimensions because of the possible presence of magnetically charged 5-branes.

Let us focus on a particular case of the above symmetry. If the RR scalar χ is taken to vanish and recalling that $g = e^{\phi_\infty}$, one of the above $SL(2, Z)$ transformations is such that:

$$g \rightarrow \frac{1}{g}, \quad B_2 \rightarrow A_2, \quad A_2 \rightarrow -B_2. \quad (2.22)$$

This particular transformation is often referred to as S-duality. It is now trivial to see that this duality is non-perturbative, since it exchanges weak and strong coupling. However, instead of exchanging at the same time an electric field with its dual magnetic field (as in the 4 dimensional dualities), it exchanges a NSNS field with a RR one, both electric. Now we had already pointed out that there are no RR-charged states in the perturbative string spectrum. We see that if the S-duality conjecture is true, then we must find non-perturbative objects carrying RR-charge.

Note that this $SL(2, Z)$ duality of type IIB strings is a duality relating different regimes of the same theory. A general $SL(2, Z)$ transformation maps the fundamental string, carrying one unit of B_2 charge, into a general (p, q) string carrying p units of NSNS charge and q units of RR charge (with p and q relatively prime). What the duality conjecture tells us is that this (p, q) string can be quantized and should reproduce, in its own variables, the same type IIB theory (see [20, 40]).

Contrary to T-duality, in the present case the fundamental string is mapped to a ‘solitonic’ string, which must have a tension going like an inverse power of g (in order not to be seen in perturbation theory). Since the tension of this dual string gives the

string scale of the dual theory, we see that α' should not be invariant under S-duality. We can actually determine how it transforms asking that the 10 dimensional effective Newton constant is invariant. Since we have $G_{10} \sim g^2 \alpha'^4$, we see that under S-duality (2.22) the string scale must transform as:

$$\alpha' \rightarrow g\alpha'. \quad (2.23)$$

An immediate consequence of the relation above is that the tension of the (0,1) string goes like $1/g\alpha'$. This is not what we would expect from a typically solitonic effect, like the energy of the electro-magnetic dual of the fundamental string which should go like $1/g^2$ because of the charge quantization condition.

U-duality

In fall 1994, the first signal of the Second Superstring Revolution was sent out with the conjecture by Hull and Townsend [20] that S-duality and T-duality actually fit into a larger group of duality symmetries, called U-duality.

Take for instance any of the two type II string theories, compactified on a d -dimensional torus T^d . The resulting theory is invariant under two sets of dualities. Firstly, T-duality extends now to a wider group, which is $O(d, d; Z)$ [32, 23]. For $d = 1$, we recover the only element described above, namely Z_2 . The invariance under $O(d, d; Z)$ can be seen as follows. If there are d compact directions, then we have d relations similar to (2.15). The ‘compact’ momenta can be shown [32, 23] to verify:

$$\vec{p}_L^2 - \vec{p}_R^2 = \frac{n_i w_i}{\alpha'}, \quad i = 1 \dots d,$$

which clearly displays $O(d, d; R)$ invariance. This group is broken to its discrete subgroup asking that the spectrum is mapped to itself. Note that the larger group denotes the fact that now one has the freedom to change some of the moduli of the torus T^d , still keeping the $10 - d$ dimensional physics invariant.

Secondly, since as soon as there is one compact direction there is no longer distinction between the two type II theories, any one of these string theories in less than 10 dimensions inherits the S-duality of type IIB theory in 10 dimensions. As we have already seen, S-duality actually extends to a full $SL(2, Z)$ duality group.

The conjecture of Hull and Townsend is that the duality group of type II string theories compactified on T^d is:

$$G_d(Z) \supset SL(2, Z) \times O(d, d; Z), \quad (2.24)$$

where $G_d(Z)$ is a discrete subgroup of the Cremmer-Julia group of the $10 - d$ dimensional maximal supergravity. A table with such groups can be found for instance in [20]. If we take for example type II theories compactified down to 4 dimensions, then the group is $E_7(Z)$ (a discrete and maximally non-compact version of E_7 , also written $E_{7,7}(Z)$).

The group $G_d(Z)$ is in general larger than $SL(2, Z) \times O(d, d; Z)$, and this implies that the orbits of any object are also much larger, i.e. U-duality maps fundamental string states to any one of the non-perturbative objects that can be predicted.

In [20], the authors have been able to find all the objects belonging to an orbit of a particular class of states, the BPS saturated supersymmetric states. What they actually find are some extreme dilatonic black hole solutions of the dimensionally reduced supergravity. However some of these black holes are identified with the winding and momentum modes of the fundamental string. Furthermore, all the other black holes are reinterpreted as wrapping modes of some higher dimensional extended objects, called p -branes⁶. Since all of these branes can be mapped by a U-duality to the fundamental string, the idea that any one of these objects can be treated as a fundamental one [41, 42] appears now natural. The string coupling constant is also mixed through U-duality with all the other moduli of the compactified theory, thus mapping strong coupling physics to some other region in the moduli space, possibly weakly coupled.

There is however something more which comes out of this conjecture, besides unifying the two type II theories and putting on the same footing all the perturbative and non-perturbative objects. It is well known that all the maximal supergravities (except the 10 dimensional type IIB) can be derived from the 11 dimensional supergravity [43]. The Cremmer-Julia groups arise from this point of view as the symmetries of 11 dimensional supergravity compactified on T^{d+1} . A remnant of this origin can be seen from the fact that each of the G_d groups has a subgroup which is the modular group of the $d + 1$ dimensional torus T^{d+1} . In the discretized version, we have:

$$SL(d + 1, Z) \subset G_d(Z).$$

If $G_d(Z)$ is a true symmetry of string theory, and accordingly its spectrum fits into representations of this group, then the same spectrum can be decomposed into representations of $SL(d + 1, Z)$.

The existence of 11 dimensional supergravity has always been intriguing in a string theory perspective, but the fact pointed out above seems a strong hint that something truly eleven dimensional is actually happening in string theory itself.

11 dimensions and M-theory

In his famous paper of spring 1995 [17], Witten uses the U-dualities to map all the strong coupling regimes of the known string theories (with the exception of the $E_8 \times E_8$ heterotic string in 10 dimensions) to some weakly coupled limit of another theory. The most astonishing surprise is that for the picture to be self-contained, one has to include in the possible ‘dual’ theories also the 11 dimensional supergravity. It has to be stressed that this is not a mathematical artifact, but a clear indication that string dynamics at strong coupling effectively sees 11 dimensions. This is not in contradiction with the critical dimension of the superstrings being 10 dimensions. In the weak string coupling, perturbative limit, the space-time in which the strings propagate can be seen to become again 10 dimensional.

Let us now briefly see how 11 dimensional supergravity really comes into play. It is well known that type IIA supergravity is the dimensional reduction of 11 dimensional

⁶ The p -branes are extensively studied in the next chapter. We thus postpone until then a thorough introduction to these objects.

supergravity (see Appendix A). It is however instructing to see how the fields of these two theories are related. In 10 dimensions, we have only one scalar, the dilaton, which will control the strength of the string interactions. In 11 dimensions, there are no scalars, and the only one which is produced by Kaluza-Klein reduction is the metric component relative to the direction along which one is reducing. This latter scalar gives rise, by its asymptotic value, to the compactification modulus called the radius of the ‘11th direction’, R_{11} . It is thus clear that the string coupling g and the compactification radius R_{11} are related.

It is actually easy to derive the exact relation between g and R_{11} using the results of the Appendix A. One finds that if $g_{11,11} = e^{2\sigma}$, the relation between the two scalar fields is $e^\phi = e^{\frac{3}{2}\sigma}$, where ϕ is the same which appears in (2.7). The relation between the string coupling and the compactification radius (in 11 dimensional Planck units) is thus:

$$g \sim \left(\frac{R_{11}}{l_p} \right)^{\frac{3}{2}}. \quad (2.25)$$

Note that the precise numerical factor in this relation and in the following ones are not one but can be determined by consistency with the dualities (see for instance Appendix B). The relation above can be rewritten as:

$$R_{11} \sim g^{\frac{2}{3}} l_p, \quad (2.26)$$

and the 11 dimensional Planck length is simply given in terms of the 11 dimensional Newton constant by $G_{11} \equiv l_p^9$.

We would like now to relate the 11 dimensional units to the string units, given by the length $l_s = \sqrt{\alpha'}$. This is straightforward since the 10 dimensional effective Newton constants have to coincide. We have thus the following relation:

$$g^2 l_s^8 \sim G_{10} \sim \frac{G_{11}}{R_{11}} = \frac{l_p^9}{R_{11}},$$

which implies:

$$l_p \sim g^{\frac{1}{3}} l_s, \quad R_{11} \sim g l_s. \quad (2.27)$$

Thus we see that at weak string coupling, $g \ll 1$, both the 11 dimensional Planck scale and the radius of the 11th direction are small compared to the string scale⁷. Perturbative string theory is thus correctly described by a 10 dimensional theory.

However, when we now consider strongly coupled type IIA superstrings, we see that the radius of the 11th direction R_{11} becomes large in both string and 11 dimensional units. The 11th direction thus effectively decompactifies. This is the core of this duality conjecture and the first glimpse at an 11 dimensional theory underlying all string theories, M-theory. It has however to be noted that for the present duality, we only have a low-energy description of the dual theory, since we do not know what is the quantum theory that reduces to 11 dimensional supergravity at low energies.

⁷Note that the two quantities above are also small compared to the 10 dimensional Planck scale, which is given by $L_p = G_{10}^{\frac{1}{4}} \sim g^{\frac{1}{4}} l_s$.

To convince oneself that the strong coupling dynamics of type IIA string theory is truly 11 dimensional, it can be shown that all the branes living in 10 dimensions have an 11 dimensional ascendant. The most celebrated example is the 11 dimensional supermembrane which gives the type IIA superstring upon dimensional reduction along its world-volume [44, 41].

An even more direct evidence of this duality can be shown following Witten [17]. In the limit of a large radius R_{11} , the KK modes propagating in this direction should become light degrees of freedom. What are the corresponding states which become light on the type IIA string side? Their mass has to go like:

$$M = \frac{1}{R_{11}} \sim \frac{1}{gl_s}. \quad (2.28)$$

They are thus solitonic objects, the mass of which becomes vanishing in the strong coupling limit. That the mass has the same expression for all values of the string coupling is ensured by the fact that these objects can be shown to preserve half of the 32 supersymmetries, and thus they fit in the shortest massive supermultiplet.

Moreover the KK 2-form field strength under which these 0-branes (since they are pointlike objects) are charged maps in the IIA language to the RR 2-form field strength. Thus we find again that a RR-charged object has a mass going like $1/g$.

The RR-charged particle discussed here and the RR-charged string discussed in the previous subsection are actually T-dual to each other. The quickest way to see it is to suppose that the type IIA theory has a compact direction and to transform accordingly the mass formula (2.28) under the T-duality (2.19). We obtain:

$$M' = \frac{R_B}{g_B l_s^2}.$$

This is exactly the mass formula for a string-like object with a tension going like $1/g\alpha'$, i.e. like the RR-charged (0,1) string of type IIB theory.

Performing a chain of T-dualities, one can find a series of brane-like objects, of even dimension in type IIA theory and of odd dimension in type IIB, which all have a tension going like $1/g$, and which all have the right dimension to couple to a RR potential. The supergravity solutions relative to these branes were found in [45]. It is amusing to see that their tensions perfectly agree with a Dirac-like quantization condition of their charges (see [46, 47]):

$$T_p T_{6-p} = \frac{2\pi}{16\pi G_{10}}, \quad (2.29)$$

where p and $6-p$ are the extensions of the electric and the magnetic brane, respectively. The exact value of the Newton constant in 10 dimensions is consistently fixed to be:

$$G_{10} = 8\pi^6 g^2 \alpha'^4. \quad (2.30)$$

A major breakthrough in string theory was the discovery by Polchinski [48] that these RR-charged branes were the Dirichlet branes, or D-branes [35]. The D-branes, and their many applications, will be reviewed in the next section. It suffices here to say that the

particular $1/g$ behaviour of the D-brane tensions is the product of the disk amplitude for the open strings that end on the D-brane.

The way in which all the dualities above and the compactification from 11 to 10 dimensions act on the various branes is summarized in Appendix B.

We can now pause a little bit and ask what is M-theory from the point of view of its duality with type IIA string theory. What we know about M-theory is that it lives in 11 dimensions, that its low-energy limit is 11 dimensional supergravity, and that when it is compactified on a vanishingly small circle it reproduces perturbative type IIA string theory. We could also say, as for every duality, that M-theory is strictly equivalent to type IIA string theory when all non-perturbative effects are taken into account. We will now see that more can be said about M-theory, and indeed many strong coupling regimes of string theories in lower dimensions can be resolved by M-theory [17].

The simplest non-trivial example of the ‘power of M-theory’ is to see how M-theory and type IIB theory are directly related [40, 49]. Consider for instance M-theory compactified on a 2-torus T^2 . The membrane of M-theory, or M2-brane, can now wrap on the T^2 to give finite energy modes in the effective 9-dimensional space-time. We also have a set of KK modes for each direction of the torus. Now, when the volume of the torus uniformly shrinks to zero size, we have one family of modes which become massless, and two sets of modes which become very massive. The appearance of massless modes is the signal of a new direction which opens up and decompactifies, these modes being identified to the KK modes of this new direction. The two sets of infinitely massive modes indicate on the other hand that there are two different string-like objects in the dual theory which cannot wind any more on the new direction. We are clearly describing type IIB string theory on a circle of growing size. Moreover, the $SL(2, Z)$ duality of the latter theory is here given a geometrical interpretation as the modular group of the original torus. The duality between M-theory on T^2 and type IIB theory on S^1 can be made more explicit working out the relations between the couplings and between the other BPS states, but the discussion above gives already a strong evidence in favour of it.

Dualities in lower dimensions [20, 17] can also involve compactification on manifolds breaking some of the supersymmetries, like the Calabi-Yau manifolds in 6 and 4 dimensions (in this latter case they go under the name of $K3$ surfaces). This allows for dualities between type II and heterotic string theories in 4 and 6 dimensions, and between M-theory and heterotic string theory in 7 dimensions. A crucial fact in establishing these dualities is that the enhanced gauge symmetry points in the heterotic string moduli space are reproduced non-perturbatively on the type II and M-theory side by branes wrapping on cycles of vanishing area or volume. This is another example of the branes becoming really elementary particles in some regions of the moduli space [50]. We will not review further these dualities here.

More dualities

Until now, we have discussed the dualities relating the superstring theories with maximal supersymmetry ($N = 2$) and 11 dimensional supergravity. We now wish to see how the

other three string theories fit into the same unified picture, and can thus also be covered by the concept of M-theory.

It was known already in the late 80s that the two heterotic string theories were related by target space duality whenever at least one of the directions was compact [32, 33]. In other words, the 10 dimensional $E_8 \times E_8$ and $SO(32)$ heterotic string theories were connected by a continuous path in the moduli space of the heterotic string theory compactified on a circle.

The discovery of the rôle played by the D-branes allowed for a reformulation of type I string theory which displayed its close relationship with type IIB strings, and at the same time provided evidence for a duality with the $SO(32)$ heterotic string.

The description of type I string theory that comes out of the work of Polchinski [48] is the following. Since D-branes are the only loci in space-time where superstrings can end, the only way to have open strings propagating in the bulk of space-time is to fill it with D9-branes. The difference with respect to the traditional understanding of type I theory is that now the D9-branes are accompanied by a RR-charge that has to be cancelled in some way.

It turns out that there is an object, called orientifold (see e.g. [51] and the next section), which implements an orientation reversal on the world-sheet. When the orientifold fills space-time (i.e. it is an orientifold 9-plane), then the procedure is equivalent to the one performed in Section 2.2 to define open superstrings. What one can show is that this object also carries a RR-charge, which is fixed once for all. Thus, since we want to have a vanishing RR-charge associated to these 9-dimensional objects, we have to take an orientifold of SO type and 16 D9-branes, which together yield a gauge group $SO(32)$ (see [26] for the details).

We have thus obtained a somewhat more heuristic understanding of type I string theory: it can really be seen as type IIB theory with some additional objects in it, and the anomaly cancellation condition which fixes the group to $SO(32)$ is reduced to the cancellation of D9-brane RR-charge.

The duality with the $SO(32)$ heterotic string can now be studied in detail in this framework, as proposed by Polchinski and Witten [52]. We know already that the two theories have the same low-energy field content. What is new in this picture is that type I theory has a soliton (which survives the projection from IIB theory) which is the D1-brane, or D-string, or (0,1) string. Contrary to the fundamental type I strings which, having the possibility to be open, do not have a winding number, the D-string is closed and can wind on compact directions. Moreover, the fundamental open strings which end on it on one side and on one of the D9-branes at the other side endow it with a world-sheet structure identical to the one of the $SO(32)$ heterotic string. For the map between the two theories to work, one also has to exchange weak and strong coupling (as already noted in [17]), the couplings being related by $g_I = \frac{1}{g_{het}}$. This heterotic-type I duality can actually be seen (with some caution) as a remnant of the S-duality of type IIB theory, after the orientifold projection [53].

We have thus related type I theory to type IIB, the $SO(32)$ heterotic string to the type I strings, and the two heterotic strings when they are compactified on a circle. All the string theories are thus related to each other, but we are still lacking of an alternative,

weakly coupled description for the strong coupling regime of the $E_8 \times E_8$ heterotic string in 10 dimensions. Here M-theory comes again into play.

It was shown by Horava and Witten [54], still before the end of 1995, that the strong coupling limit of the $E_8 \times E_8$ heterotic string can be interpreted as M-theory compactified on a segment, the length of the segment being related to the heterotic string coupling by a relation similar to (2.26). This compactification of M-theory can be considered as an orbifold compactification on S^1/Z_2 (see e.g. [6] for a description of the orbifolds in string theory), and thus it correctly breaks half of the original 11 dimensional supersymmetry to give $N = 1$ supersymmetry in 10 dimensions. 10 dimensional physics is actually associated to the physics on the ‘ends of the world’, i.e. on the hypersurfaces determined by the two endpoints of the segment. It can be seen by anomaly cancellation arguments (see [54] for all the details on this particular duality) that each of the two 10-dimensional boundaries must have an E_8 gauge theory defined on it. Moreover, the orbifold procedure allows open membranes of M-theory to stretch from one end-of-the-world to the other, thus giving effective closed strings in 10 dimensions. These are to be identified with the fundamental heterotic strings. When the size of the segment shrinks to zero size, we have a closed string theory at weak coupling carrying a gauge group which is $E_8 \times E_8$. This is thus the explanation of the duality between M-theory and the $E_8 \times E_8$ heterotic string. Much like the type IIA string, there is an 11th direction opening up at strong coupling.

Let us conclude this section by some comments on the picture that comes out of this web of dualities.

We seem to know 6 limits in the moduli space of what we can now call M-theory: these are the 5 string theories supplemented by 11 dimensional supergravity. As far as the string theories are concerned, we seem to know something more than the low-energy effective action because we can analyze them perturbatively. On the other hand, we only know the low-energy effective action of the conjectured 11 dimensional theory. Note however that this 11 dimensional theory has no natural dimensionless parameter to define a perturbative expansion, such as the string coupling g for the 10 dimensional theories.

M-theory represents what stands in the middle, and allows us to relate all these different limits. It can thus be considered as the region of intermediate coupling, generically a much wider region than the others. It is difficult to have an idea of what this unified theory really is, since it does not have to reproduce all the known string theories at the same time, but rather its existence implies that all the string theories describe the same physics.

When we are at intermediate coupling, it is almost an option to say that we are considering an originally 11 dimensional theory on a circle of a finite radius, or rather a string theory at finite coupling. It is however fair to think about M-theory as an 11 dimensional theory, since there are effectively regions in the moduli space where the physics is truly 11 dimensional. On the other hand, nothing prevents us in principle to choose our favourite string theory and try to solve it completely, with the goal of knowing what M-theory is. This approach has already proven to be hopelessly hard. Nevertheless, there is the hope that M-theory itself can be addressed directly, short-

cutting the traditional string theory path. We will indeed review in Section 2.6 a proposal going along this way.

2.5 D-branes and their many uses

The construction leading to the concept of D-brane is actually very simple: instead of imposing the usual Neumann boundary conditions to the open strings, one imposes Dirichlet boundary conditions in some directions. The D-branes are then the hypersurfaces defined by these conditions. It turns out that the open strings constrained to end on D-branes, couple consistently to the closed type II strings. Furthermore, what makes the D-branes interesting from the string duality point of view is that they precisely couple to the RR bosons of the closed string, which means that the D-branes are the sought for RR-charged objects. The D-branes are thus solitonic objects in the sense that they are not part of the perturbative string spectrum, but their properties can nevertheless be studied in the framework of perturbative string theory. Most notably, one finds that the D-branes are dynamical objects.

Let us define the D-branes simply by the boundary conditions one imposes to the open string sector (the standard reference for D-branes is [26]).

In order for the action (2.1) to give the correct equations of motion for $X^\mu(\tau, \sigma)$ in the case of an open superstring (we now take $0 \leq \sigma \leq \pi$), one generally imposes Neumann boundary conditions:

$$\partial_\sigma X^\mu|_{\sigma=0,\pi} = 0. \quad (2.31)$$

These are the conditions usually imposed when defining type I string theory, and they imply that the ends of the string are freely moving. In terms of the left and right moving pieces of X^μ (2.12), the conditions rewrite:

$$\partial_+ X_L^\mu - \partial_- X_R^\mu|_{\sigma=0,\pi} = 0. \quad (2.32)$$

The mode expansion of the left and right moving fields X_L^μ and X_R^μ is much similar to (2.13), except that now there is only one kind of oscillator, $\alpha_n^\mu = \tilde{\alpha}_n^\mu$, and, most notably, $p_L^\mu = p_R^\mu$ is imposed regardless of the fact that a direction is compact or not. The complete expansion of the bosonic field X^μ is thus:

$$X^\mu = x_0^\mu + 2\alpha' p^\mu \tau + 2i\sqrt{\alpha'} \sum_{n \neq 0} \frac{1}{n} \alpha_n^\mu e^{-in\tau} \cos n\sigma. \quad (2.33)$$

Note that x_0^μ and p^μ are canonical conjugate variables, and thus x_0^μ correctly describes a dynamical variable, i.e. the position of the string.

Let us now impose Dirichlet boundary conditions instead of the Neumann ones in the $9 - p$ directions X^i , with $i = p + 1, \dots, 9$. We simply impose:

$$\partial_\tau X^i|_{\sigma=0,\pi} = 0, \quad (2.34)$$

or, in terms of the left and right moving fields:

$$\partial_+ X_L^i + \partial_- X_R^i|_{\sigma=0,\pi} = 0. \quad (2.35)$$

The remaining $p + 1$ directions (including the timelike one) still obey the Neumann boundary conditions (2.31) or (2.32). Now again the mode expansion for X_L^i and X_R^i will be similar to (2.13), but in the present case the identifications will be $\alpha_n^i = -\tilde{\alpha}_n^i$ and $p_L^i = -p_R^i$. Going back to the expression (2.15), this last condition allows only for a winding number but not for any momentum in the direction X^i . This is fully consistent with the heuristic picture: the Dirichlet boundary conditions constrain the ends of the open string to live on a particular $p + 1$ dimensional hypersurface (the D-brane). However, since now the open string can no longer break in two pieces in an arbitrary locus of space-time, it can indeed wind on a compact direction. Here one has even an additional possibility, in that the open strings can stretch between two D-branes separated by some distance (thus giving some kind of ‘fractional winding’).

A consequence of (2.34) is that we can choose the two fixed values indicating where the ends of the open string lie:

$$X^i(\tau, \sigma = 0) = x^i, \quad X^i(\tau, \sigma = \pi) = y^i.$$

The mode expansion for X^i is then:

$$X^i = x^i + \frac{1}{\pi}(y^i - x^i)\sigma + 2\sqrt{\alpha'} \sum_{n \neq 0} \frac{1}{n} \alpha_n^i e^{-in\tau} \sin n\sigma. \quad (2.36)$$

Since $p^i = p_L^i + p_R^i \equiv 0$, x^i is now no longer a dynamical variable, and can thus be taken to vanish for simplicity. If there is only one D-brane and the direction X^i is compact with radius R_i , then we can replace y^i by $2\pi w R_i$. Note however that only the mode $w = 1$ is topologically conserved, since for higher w there is always the possibility for the string to break when it crosses the D-brane.

The fact that the zero modes of the open string in the X^i directions are not dynamical implies that the low-energy fields corresponding to the massless modes of the open sector do not depend on these coordinates. Accordingly, they should be arranged into representations of $SO(1, p)$ instead of $SO(1, 9)$. This means that the low-energy effective theory of these modes is $p + 1$ dimensional. It is however very easy to determine this effective theory starting from the knowledge of the effective theory for the open strings with Neumann boundary conditions, which is $N = 1$ super Yang-Mills (SYM) in 10 dimensions. Indeed, the quantization of the open string with Dirichlet boundary conditions is exactly the same as for the open sector of the type I superstring. For instance, the Dirichlet boundary conditions change an overall sign for ψ_R^i by world-sheet supersymmetry, and thus do not change the integer and half-integer moding in the Ramond and Neveu-Schwarz sectors respectively. The net result is that one only has to decompose the $\mathbf{8}_v + \mathbf{8}_+$ states under the $p + 1$ dimensional little group $SO(p - 1)$. The low-energy effective theory one obtains is (to leading order in α') the SYM theory in $p + 1$ dimensions, which is actually the dimensional reduction of the 10 dimensional one. The gauge field components with indices in the X^i directions are now scalars.

The D-brane, as we will argue shortly, couples to the closed string modes of the type II strings. However, by its presence, it breaks half of the supersymmetries since this is the maximum space-time supersymmetry that the open strings (with Dirichlet boundary

conditions in this case) can support. The situation is thus that, far away from the D-brane, there seem to be two space-time supersymmetries, however their generators are related to each other on the world-sheet of the open strings, which are localized near the D-brane (see the discussion in the original paper of Polchinski [48]).

The relation between the two supersymmetry generators can actually be derived considering how the D-branes behave under T-duality.

Recall that T-duality acted for the closed string as a parity reversal restricted to the right moving modes (2.18). If we do the same operation on an open string, we see from (2.32) and (2.35) that Neumann and Dirichlet boundary conditions are interchanged. If we thus start from a theory of open superstrings and T-dualize on the $9 - p$ directions X^i , we obtain the same Dp -brane theory as defined above (p is the number of space-like directions longitudinal to the D-brane). This is actually one of the ways the D-branes are usually introduced [26]. Note however that in this formulation, the X^i directions have to be compact, while as we have seen above the D-branes can be defined in 10 non-compact directions.

It is now straightforward to see how D-branes transform under T-duality. If the T-duality is performed along a direction transverse to the Dp -brane, then the latter transforms into a $D(p+1)$ -brane longitudinal to the T-dualized direction, while if the T-duality is performed along a direction longitudinal to the world-volume of the Dp -brane, then it becomes a $D(p-1)$ -brane transverse to the T-dualized direction.

The supersymmetry projection associated with a D-brane can now be found. If we start with type I open strings, the Neumann boundary conditions impose that the two generators of space-time supersymmetry Q_L and Q_R must be equal, $Q_L = Q_R$ [6, 26] (up to a conventional sign). For this to be consistent, both generators have to have the same space-time chirality. Type I open superstrings can then be (formally) seen as type IIB superstrings coupled to a D9-brane (which fills space-time).

We can now introduce Dirichlet boundary conditions by the implementation of space-time parity reversal on the right moving modes. The effect on the supersymmetry generators of this operation in, say, the X^9 direction is the following:

$$Q'_L = Q_L, \quad Q'_R = P_9 Q_R. \quad (2.37)$$

P_9 is an operator which anticommutes with Γ_9 and commutes with all the others. It can be taken to be $P_9 = \Gamma_9 \Gamma_{11}$, where $\Gamma_{11} = \Gamma_0 \dots \Gamma_9$ is the matrix defining chirality.

A D-brane lying in the $X^1 \dots X^p$ directions can thus be seen to impose the following relation between the supersymmetry generators:

$$Q_L = \pm P_{p+1} \dots P_9 Q_R. \quad (2.38)$$

Using the fact that Q_R is chiral or anti-chiral, the relation above can be rewritten as:

$$Q_L = \pm \Gamma_0 \dots \Gamma_p Q_R. \quad (2.39)$$

Note that since T-duality interchanges type IIA and IIB theories, we only have supersymmetric D-branes of odd dimension in type IIB theory and of even dimension in type IIA theory.

The computation of Polchinski [48] can now be simply stated. The static force between two parallel D-branes is given by the amplitude corresponding to a cylinder stretching between the two D-branes, without any insertion. This is a one-loop amplitude from the open string point of view, but it is a tree-level amplitude in the closed channel which goes from one D-brane to the other. The result of the computation is that the amplitude vanishes, thus implying that there is no static force between parallel D-branes. The D-branes are truly BPS states, as their supersymmetric properties might have suggested.

More can be said [48] about this vanishing amplitude. In the closed channel it can be seen that it actually vanishes because two terms cancel: the exchange of NSNS particles on one side (the graviton and the dilaton), and the exchange of RR gauge bosons on the other. This is the central result, i.e. the proof that the D-branes carry a RR-charge, and that their charge is equal to their tension. Moreover, the tension can be extracted and it matches exactly with the one which was guessed by duality arguments in the previous section:

$$T_{Dp} = \frac{1}{(2\pi)^p g l_s^{p+1}}. \quad (2.40)$$

Let us now review some generalizations of the theory of D-branes as described above.

The most straightforward generalization was elaborated on by Witten [55], and consists in taking a stack of N D-branes of the same kind. As just noted above, the D-branes are BPS objects and can thus stay statically at any distance from each other. The presence of N D-branes has the direct consequence of introducing a ‘quantum number’ associated to each end of the open strings and which distinguishes between the various D-branes on which the end can be attached. This quantum number is nothing else than a Chan-Paton factor, as discussed in Section 2.2. Here however the situation is slightly different. If the N D-branes coincide (i.e. they lie on top of each other), then the effective theory on the branes is $U(N)$ SYM in $p+1$ dimensions (in the case of oriented strings). On the other hand, if all the D-branes are separated by a finite distance, the effective theory at low-energies becomes the same SYM but now with a group $U(1)^N$.

The mechanism by which the group is broken from $U(N)$ to $U(1)^N$ exactly maps to spontaneous breaking of local symmetry in the low-energy effective action. It can be seen that the scalars of the SYM theory actually represent the location of the branes. In the case of a $U(N)$ gauge theory, the scalars belong to the vector multiplet of SYM and are thus $U(N)$ matrices. The diagonal values can be associated to the locations in transverse space of the N D-branes. Pulling the D-branes apart is equivalent to giving an expectation value to these scalars. Maximal supersymmetry on the effective theory side, and the BPS property in the D-brane picture, ensure that there are flat directions in the potential, along which such expectation values can be arbitrary. Since the scalars are in the adjoint representation, the gauge group can be broken maximally to $U(1)^N$.

The gauge bosons which get masses because of the expectation values of the scalars are, in the D-brane picture, nothing else than the lowest lying, now massive, modes of the open strings connecting two different, separated, D-branes. The phenomenon of spontaneous breaking of gauge symmetry thus earns a very beautiful heuristic picture in the D-brane framework.

Let us consider in more detail the discussion above introducing the reduction to $p+1$ dimensions of the 10 dimensional SYM theory. This is done straightforwardly. The ten dimensional Yang-Mills field strength is given by:

$$F_{MN} = \partial_M A_N - \partial_N A_M + [A_M, A_N], \quad M, N = 0, \dots, 9.$$

The 10 coordinates are split into two sets, $x^M = \{\xi^\mu, y^i\}$ ($\mu = 0, \dots, p$ and $i = p + 1, \dots, 9$), the fields being now independent of the second set. We thus have:

$$\begin{aligned} F_{\mu\nu} &= \partial_\mu A_\nu - \partial_\nu A_\mu + [A_\mu, A_\nu], \\ F_{\mu i} &= \partial_\mu A_i + [A_\mu, A_i] \equiv D_\mu A_i, \\ F_{ij} &= [A_i, A_j]. \end{aligned}$$

If we rename $A_i \equiv \phi_i$ since they are now scalars, the bosonic part of the action of SYM in $p + 1$ dimensions writes:

$$I_{p+1} = -\frac{1}{g_{YM_{p+1}}^2} \int d^{p+1}\xi \left(\frac{1}{4} \text{tr} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} \text{tr} D_\mu \phi_i D^\mu \phi^i + \frac{1}{4} \text{tr} [\phi_i, \phi_j]^2 \right). \quad (2.41)$$

Because of the presence of a total amount of 16 supercharges, all the matter fields present in this SYM theory are actually part of the vector multiplet containing the gauge field. Therefore all the scalars, as well as the spinors, belong to the adjoint representation of $U(N)$. The third term in the action (2.41) is the potential, and it has flat directions for commuting vacuum expectation values of the scalars. Moreover, due to the large amount of supersymmetry, non-renormalization theorems exist (or are conjectured) which ensure that these directions remain flat even at the quantum level. Any vacuum configuration with diagonal ϕ_i 's has a vanishing potential, and for generic values of the diagonal elements the gauge group is broken down to $U(1)^N$. In the D-brane picture, this corresponds to widely separated D-branes.

The SYM coupling constant $g_{YM_{p+1}}$ is fixed in terms of string variables when one considers that the action (2.41) is the leading term in the expansion of the Born-Infeld action for a D-brane [56]. Since the prefactor of the Born-Infeld action is simply the tension of the D-brane T_{Dp} as given in (2.40), we have the following relation (where all numerical factors have been dropped):

$$g_{YM_{p+1}}^2 = \frac{1}{T_{Dp} \alpha'^2} = g_s^{p-3}. \quad (2.42)$$

Note that the α'^2 factor comes in because the fields A_μ and ϕ_i both have a mass dimension of one. The action (2.41) is thus the leading term for the dynamics of D p -branes both in the g and in the α' expansion.

Another kind of generalization of the D-brane physics is to take, for some directions, Neumann boundary conditions on one side and Dirichlet boundary conditions on the other side (these directions will be called ND directions). This physically corresponds to having two D-branes of different dimensionality at the same time, and open strings going from one to the other.

The ND directions now introduce a change of sign in the right moving (bosonic and fermionic) fields only at one boundary, thus interchanging integer and half-integer mode expansions. The energy of the NS sector ground state now depends on the number ν of these ND directions. Aside from the previous (trivial) $\nu = 0$ case, let us only state here that the case $\nu = 4$ gives a vanishing energy for the NS ground state, and a reduced number of massless states which correspond to a further breaking of supersymmetry. Indeed, each of the two D-branes comes with a different projection of the kind (2.39). The whole configuration preserves 1/4 of the original supersymmetry. The other cases are trickier to treat, it suffices here to say that in the $\nu = 2$ and $\nu = 6$ case there is respectively an attractive and a repulsive static force, while in the $\nu = 8$ case there is also a vanishing static force and the configuration turns out to be supersymmetric (note that ν cannot of course be odd because the two D-branes must belong to the same type IIA or IIB theory).

There is an object very similar to the D-branes that can also be defined, that is the orientifold [51, 26]. Briefly stated, the orientifold procedure amounts to accompanying an orientation reversal on the world-sheet by a parity reversal in space-time, and then modding out by this symmetry. When no parity reversal is taken, then we are simply defining unoriented strings. When on the other hand we take a parity reversal in $9 - p$ directions, the orientifold procedure Ωp acts as follows on the closed strings (it acts similarly on the open ones):

$$\begin{aligned} X^\mu(\tau, \sigma) &= X^\mu(\tau, 2\pi - \sigma), & \mu &= 0 \dots p, \\ X^i(\tau, \sigma) &= -X^i(\tau, 2\pi - \sigma), & i &= p + 1 \dots 9. \end{aligned} \quad (2.43)$$

In the bulk of space-time, i.e. far away from the orientifold Ωp -plane located at $X^i = 0$, the strings are oriented, but (2.43) states that they have a ‘mirror’ partner on the other side of the Ωp -plane. On the Ωp -plane itself, the strings are unoriented. The introduction of such an object can be seen to impose a relation on the supersymmetry generators which is exactly the one for a Dp -brane (2.38). Actually, a one-loop computation like the one performed by Polchinski for the D-branes can show that an orientifold Ωp -plane carries 2^{p-5} units of the same RR-charge carried by a Dp -brane [26].

An orientifold plane can be combined with D-branes, for instance if one wants to cancel the RR-charge of the orientifold. If one puts N D-branes on one side of the orientifold, then one has also to take into account the existence of the ‘mirror’ D-branes on the other side⁸. The gauge theory one recovers on the world-volume of the D-branes is $SO(2N)$ or $Sp(2N)$, depending on how the world-sheet orientation reversal has been defined on the string states. The main physical difference between an orientifold and a D-brane is that the former is not dynamical, since it has no low-energy effective action associated to its world-volume.

Let us now briefly review some of the applications where the D-branes have proven to be very useful. We single out here three main successes of the D-branes: proving that the U-duality orbits could be filled; making the microscopic counting of the entropy of

⁸ Even if there is a total amount of $2N$ different Chan-Paton factors, it is important to note that there are only N dynamical D-branes.

some black holes accessible; providing a pictorial way to understand some gauge theory phenomena, such as dualities.

In the preceding section we have reviewed the dualities which led to the concept of M-theory. Many of these dualities predicted dual objects with mass going like $1/g$, and carrying RR-charge. The effective presence of these objects in the string spectrum, even if non-perturbative, is crucial in providing evidence for the conjectured unification of string theories under M-theory.

The existence of RR-charged solitons was already proved in [45] at the level of the classical solutions of the supergravities. This has the same status as the existence of the NSNS 5-brane soliton (or NS5-brane), which is also a matter of having a consistent background for the fundamental string [57]. In this latter case however, the solitonic 5-brane is the magnetic dual of the fundamental string, for which a perturbative definition obviously exists. The status of the RR-charged solitons is somewhat different, in that neither the electric nor the magnetic object existed in perturbation theory. With the advent of D-branes, the nature of the RR-charged object has been made more precise: they appear indeed as solitons, but their quantum mechanical properties can be analyzed quite in detail from string perturbation theory, much more than what can be said about the NS5-brane. Actually, the particular behaviour of their mass in the string coupling constant permits to study them in a flat background, because the strength of their gravitational interaction is given by $G_N M_{Dp} \sim g$, contrary to the NS5-brane where the same strength is of order 1 and cannot be neglected. The D-branes can thus be thought of as the weak coupling description of the RR-charged solitonic supergravity solutions.

Another feature of D-branes that allows a thorough study of their properties is the knowledge of their effective action. For instance, one can in this context compute very accurately the scattering of two D-branes (performing a computation either in string perturbation theory, or in the low-energy effective field theory). An interesting result [58, 59, 60] supporting the M-theory conjecture is that the scattering of D0-branes, or D-particles, displays a characteristic length scale of the order of the 11 dimensional Planck scale (2.27).

The application of D-brane physics to the problem of the entropy of black holes has been spectacular [61, 62]. The idea is simple and is the following. In the context of the low-energy effective supergravity theories, one can find all sorts of black holes, of any space-time dimension. Most notably, any one of the solitonic solutions discussed above (and treated at length in the next chapter) can be seen as a black hole in the space transverse to its world-volume. Now we have just seen that at weak string coupling, we have a fairly accurate quantum description of some of these solitons, i.e. those carrying RR-charge. It is thus tempting to study whether some of the properties of the D-branes can be transposed to properties of quantum black holes.

If one focuses on the problem of the microscopical origin of the black hole entropy, there are two more ingredients that one needs. First of all, if the computation at weak coupling is to teach us something about the physics of black holes, which are generally defined when gravitational effects are important, then we need to be confident that the weak coupling computation is not renormalized when extrapolated to strong coupling. This is done restricting our attention to supersymmetric BPS configurations, correspond-

ing to extreme black holes. It is actually the main limitation of this computation. The second requisite is that the area of the horizon, which gives the entropy, is still finite in the extreme limit. This is quite restrictive, so one has to focus only on 5 and 4 dimensional extreme black holes.

The most accessible setting for the computation of the microscopical entropy is the one presented in [62]. Here a five dimensional black hole carrying 3 charges is considered. Two of the charges are of RR type, and the third one is a KK momentum in an internal direction (all the directions longitudinal to the D-branes are now taken to be compact). The massless modes of the ($\nu = 4$) system consisting of N_1 D1-branes inside N_5 D5-branes are identified, and are taken to contribute to a total amount of N_{KK} units of internal momentum (in the direction of the D-string). In order for the configuration to preserve some supersymmetry and to be extremal, the massless modes can only be, say, right moving. The statistical entropy is then given by the logarithm of the degeneracy of such a configuration, and it exactly matches (including all the numerical factors) the semiclassical Bekenstein-Hawking entropy of the corresponding black hole solution.

Note that the entropy, being a physical quantity, must be U-duality invariant. It can thus be computed once for all in any one of the configurations which belong to the same U-duality orbit.

Even if away from extremality it is not clear how the strong and weak coupling pictures can exactly be related, the D-brane picture gives a very nice heuristic picture of how the Hawking radiation could be explained in an evidently unitary way. Consider the same 5 dimensional configuration as before, but now with the massless modes going in both directions along the D-string. This corresponds to a non-extremal black hole, with non-vanishing Hawking temperature, and which thus tends to radiate its energy. The mechanism by which it radiates translates in the D-brane picture by a process in which two open string modes travelling in opposite directions collide and produce a closed string which propagates away from the D-brane compound. The corresponding amplitude is given by the disk with two insertions at its boundaries and one insertion in the bulk. The same amplitude gives the reverse process, a particle infalling into the black hole. These amplitudes have been computed, and matched to the corresponding semi-classical amplitudes of particle emission and absorption by black holes [62, 63].

This microscopic explanation of the black hole entropy with the use of D-branes is a very important result, but there are still some obscure points in our understanding of black holes. For instance, it is still not clear what is the real nature of the event horizon, which only appears at strong coupling, and is simply not present in the D-brane picture used to compute the entropy. We will come back to the problem of black hole entropy in M-theory in Chapter 6.

The last topic we will quickly overview is the application of D-branes to the construction of gauge theories in several dimensions [64, 65, 66]. As we have already discussed, the theory on the world-volume of the branes is at low-energies SYM with gauge group $U(N)$, where N is the number of parallel branes. The amount of supersymmetry is equivalent to $\mathcal{N} = 4$ in 4 dimensions (we now use \mathcal{N} for the number of supersymmetries in order not to confuse it with the number N of D-branes). One way to reduce the supersymmetry is to suspend the D-branes between two NS5-branes. It can be seen that

the boundary conditions on the D-brane world-volume break one half of the supersymmetries. Furthermore, the direction transverse to the NS5-branes is of finite extent, and it results that D p -branes produce a p -dimensional theory with $\mathcal{N} = 2$ supersymmetry. If the two NS5-branes are rotated with respect to each other, one can further break the supersymmetry down to $\mathcal{N} = 1$. Additional D-branes perpendicular to the original D p -branes constituting the gauge theory provide matter fields arising from the open strings connecting the two kinds of D-branes.

It can then be shown that moving around the branes of the set-up described above, and using the dualities of the underlying string theory (including the possibility to view the configuration as an 11 dimensional set up [67]), lots of informations on the classical and quantum moduli space and on the dualities of the gauge theories can be determined and matched to field theory results, which can thus be extended in this way. This is actually only one example of how one can learn about the non-perturbative aspects of field theories by embedding them into M-theory.

To conclude with, we can mention that one of the applications of D-branes has also been the possibility to formulate Matrix theory, which is the subject of the next section.

2.6 Matrix theory approach to M-theory

In this section we discuss the proposal of Banks, Fischler, Shenker and Susskind [24] to treat M-theory directly at the non-perturbative level by formulating it in the infinite momentum frame. Some time after the original conjecture of [24], which appeared in fall 1996, the model was given an alternative, but similar, formulation in the discrete light cone quantization by Susskind [68]. Subsequently, its rigorous definition emerged from the two contemporary papers of Sen [69] and Seiberg [70]. Some reviews on Matrix theory already exist [71]. It is important to note that unlike the material that has been discussed until now, Matrix theory as a description of M-theory is still at a highly conjectural level, and we will indeed mention some of the problems which might prevent the establishment of Matrix theory as the ‘standard’ approach to M-theory.

Let us preliminarily discuss what are the infinite momentum frame (IMF) and the discrete light cone quantization (DLCQ) for a general theory.

The basic idea leading to the formulation of a theory in the IMF is the following. One takes a particular direction in the (flat) space-time, say, \tilde{x}_{11} . An arbitrary distribution of relativistic particles will be described, in particular, by momenta in the 11th direction \tilde{p}_{11} of both signs and of all magnitudes. Let us now perform a boost of parameter s in \tilde{x}_{11} :

$$\begin{aligned} t &= \cosh s \tilde{t} - \sinh s \tilde{x}_{11} \\ x_{11} &= -\sinh s \tilde{t} + \cosh s \tilde{x}_{11}, \end{aligned} \tag{2.44}$$

The momenta transform accordingly:

$$\begin{aligned} p_0 &= \cosh s \tilde{p}_0 + \sinh s \tilde{p}_{11} \\ p_{11} &= \sinh s \tilde{p}_0 + \cosh s \tilde{p}_{11}, \end{aligned} \tag{2.45}$$

Fixing the convention to have $p_0 > 0$, and choosing the boost parameter s to be also positive, we see that for a sufficiently large boost, p_{11} can always be made positive

because for any particle of mass m and transverse momentum p_\perp , the mass-shell relation $-\tilde{p}_0^2 + \tilde{p}_{11}^2 + p_\perp^2 = -m^2$ implies $\tilde{p}_0 \geq |\tilde{p}_{11}|$. Furthermore, we can also make p_{11} arbitrarily large compared to any other scale in the problem. When we have done so, the energy of a particle in the IMF can be extracted from p_0 :

$$E = p_0 - p_{11} = \sqrt{p_{11}^2 + p_\perp^2 + m^2} - p_{11} = \frac{p_\perp^2 + m^2}{2p_{11}}, \quad (2.46)$$

in the limit in which p_{11} is very large. We thus see that the energy acquires a non-relativistic form, with p_{11} playing the rôle of the non-relativistic mass.

We see that the bulk of the particles of the theory we are considering have an energy in the IMF which is of $\mathcal{O}(1/p_{11})$, with $p_{11} \rightarrow \infty$ denoting now the generic scale of the momenta in the 11th direction (as e.g. the total momentum of all the particles). We might be worried about the particles which after the infinite boost still have a vanishing or negative p_{11} . Their IMF energy as defined in (2.46) will be however of $\mathcal{O}(1)$ or more in p_{11} , which means that they will be very energetic with respect to all the others. It is in this sense that they can be neglected, even if the procedure of integrating them out should be by no means trivial.

Further boosts in the 11th direction of the kind (2.45) simply produce a concomitant rescaling $p_{11} \rightarrow e^s p_{11}$, $E \rightarrow e^{-s} E$. A boost in the transverse directions x_\perp has to be small with respect to the scale determined by p_{11} , and the consequence is that the system has Galilean invariance instead of Lorentz invariance in the transverse directions. For a small boost parameter v , the transverse momenta can be seen to transform like $p_\perp \rightarrow p_\perp + p_{11} v$, with p_{11} acting again as the non-relativistic mass.

In order to regulate the theory, we can further take the direction x_{11} to be compact of radius R . The momentum p_{11} is accordingly quantized,

$$p_{11} = \frac{N}{R}. \quad (2.47)$$

N is the number of partons, i.e. of units of momentum. Since any particle in the IMF formulation of the theory has a non-vanishing momentum in the 11th direction, it can be said to be made of partons. The dynamics of partons then describes the whole theory in uncompactified space-time, when p_{11} is taken to infinity, along with the radius R . This implies that one has to consider the large N limit of the theory of partons.

The principle of the IMF formulation of a theory is thus the following: if one has the theory governing the dynamics of partons, then its extrapolation to large N (in order to have large p_{11}) and large R should give the complete non-perturbative description of the original theory.

The basic requirement on the theory of partons is that it has to have Galilean invariance in the transverse directions. This invariance is promoted to its supersymmetric version if the original theory was supersymmetric, however note that generically the theory of the partons has half of the supersymmetries of the full theory because the other half is broken by the presence of the non-vanishing momentum in the 11th direction.

Before going on to see what is the theory of partons for M-theory, let us briefly discuss the DLCQ formulation for a general theory. In this case one sets up from the

beginning in the light cone coordinates $x^\pm = \frac{1}{2}(t \pm x_{11})$, and takes x^+ to act as the ‘new’ timelike coordinate. Accordingly, the hamiltonian will be given by its canonical conjugate variable, p_+ (where we have $p_\pm = p_0 \pm p_{11}$). From the mass-shell relation $-p_+p_- + p_\perp^2 = -m^2$, one extracts an expression for the light-cone energy:

$$p_+ = \frac{p_\perp^2 + m^2}{p_-}. \quad (2.48)$$

The momentum p_- now acts as the non-relativistic mass, and we also see that in this formulation the modes which originally had $p_- = 0$ are neglected (the only such modes are massless ones propagating along the x^- direction; see [72] for a discussion of the possible problems in discarding these modes).

We can now periodically identify the lightlike coordinate x^- , with period proportional to R (we do not care about numerical factors throughout this section). The conjugate momentum is thus quantized, $p_- = \frac{N}{R}$, playing thus the same rôle as p_{11} in the IMF.

Note that performing a boost similar to (2.44)–(2.45), the momenta are rescaled by $p_\pm \rightarrow e^{\pm s} p_\pm$. This might seem the opposite behaviour with respect to E and p_{11} in the IMF. It has however the same physical meaning. In the IMF, performing a further boost amounted to increase the number of partons N . Only after these rescalings x_{11} could be compactified, when the scale of p_{11} had already been fixed. On the other hand, in the DLCQ the number of partons is fixed, and one can still perform a boost which has the effect of rescaling x^- and thus R , which goes to $e^s R$. From this follows the behaviour of p_- and p_+ under the boosts in the light cone direction.

The DLCQ makes contact with the original theory in non-compact space-time when we take $R \rightarrow \infty$ and $N \rightarrow \infty$ while keeping p_- fixed.

The same theory of partons can thus describe the original theory when the latter is formulated in the DLCQ. The DLCQ is a finite N formulation of the theory in the sense that in the case of the IMF one has to take $N \rightarrow \infty$ before taking the uncompactified limit (i.e. even the regulated theory needs the $N \rightarrow \infty$ limit), while in the present case the large N limit should only be taken to recover the unregulated theory.

Let us now see why it is convenient to formulate M-theory in the IMF or in the DLCQ. The main step in the two latter formulations is having a good description of the parton dynamics. In the case of M-theory, we are particularly lucky because even if M-theory itself does not have a formulation, an elementary knowledge of string dualities allows us to determine completely the theory of partons!

The key remark, which crucially uses the fact that one has to go through a regulated, compactified theory, is that M-theory with a compact direction is nothing else than type IIA string theory. Moreover, momentum in the compact 11th direction maps to the electric charge of the RR 2-form field strength of type IIA theory. Now we have already argued in the preceding sections that there are no perturbative string states carrying such charge. Thus, the theory of N partons reduces to the theory of N D0-branes of type IIA theory. It is important to note that since p_{11} (or p_-) is strictly positive, anti-D0-branes are excluded from the model to begin with. The type IIA theory in which the D0-branes live is a totally auxiliary theory, not to be confused with the theory that

should be given by M-theory (as described by Matrix theory) when compactified on a direction other than x_{11} (or x^-).

M-theory is then recovered when the number of D0-branes is taken to infinity, along with the radius R . This should provide a complete and non-perturbative description of M-theory in 11 non-compact directions (we will see shortly that there are some non-trivial complications in describing compactifications of M-theory). In particular, there are other objects in type IIA string theory that can carry a D0-brane charge, most notably the other D-branes (this is so because their world-volume action couples to the RR 1-form potential through Wess-Zumino terms, see e.g. [73]). The D0-brane theory should then reproduce by itself the theory of all the other D-branes, which are interpreted as M-branes (or branes of M-theory) upon decompactification.

The theory of N D0-branes is easily derived from the D-brane physics described in the preceding section. In the limit of small string coupling g and large string tension $\alpha' \rightarrow 0$, it is given by the reduction to 0+1 dimensions of the 10 dimensional super Yang-Mills (SYM) theory with $U(N)$ gauge group. It is thus some sort of supersymmetric quantum mechanics of matrices, hence the name Matrix theory, also spelled M(atr)ix theory to indicate that it should describe M-theory. We will present this SYM theory hereafter. The non-trivial step in the Matrix theory conjecture is that one can consistently consider the theory of N D0-brane as decoupled from the rest of the type IIA theory, and that this limit correctly reproduces the IMF or DLCQ limit.

Let us note that Matrix theory is a quite elaborated parton model, since the set of N parton coordinates is actually promoted to a set of $N \times N$ non-commuting matrices, thus introducing a lot of extra modes giving the non-trivial interactions between the partons.

Before going to the precise definition of the model, it is worthwhile mentioning the concrete facts presented already in [24], and which supported this bold conjecture.

The first fact is that the D0-brane scattering reproduces the scattering of gravitons (and their supersymmetric partners) in 11 dimensions. That the D0-branes interact at a length scale which is the 11 dimensional Planck length was already shown in [58, 59, 60], but in [24] the scattering of 11 dimensional massless particles was reproduced with the exclusive use of the SYM quantum mechanics. When in the scattering process there is no transfer of momentum in the 11th direction, then the computation is performed in the limit in which the matrices are commuting, which corresponds to the limit of widely separated clusters of D0-branes. Note that interactions with p_{11} transfer are more difficult to treat because they involve rearrangements of D0-branes between different clusters. This first evidence is an important result, even if one could have foreseen it simply by duality arguments, for instance noting that after all the D0-brane and the 11 dimensional massless particles are the same objects but in two different, dual, pictures. The non-trivial step is really that the parton model is enough to reproduce the result.

The second, and maybe more astonishing fact, is that the Matrix model reproduces also the membranes of M-theory, or M2-branes. The fact that D2-branes could be collective excitations of D0-branes was already suggested by Townsend [74], and both remarks are actually based on the attempt by de Wit et al. [75] to quantize the 11 dimensional supermembrane by formulating it in a regularized set-up. This regularized supermem-

brane theory exactly coincides with the supersymmetric quantum mechanics governing the dynamics of the D0-branes. The smooth supermembrane is however recovered when in the $N \rightarrow \infty$ limit the matrices have a fundamentally non-commuting behaviour. In the context of Matrix theory, the non-trivial check is that the tension of the matrix membrane exactly matches the one which is derived for the M2-brane by duality arguments. The membrane formalism also provides us with a proof that, at least at the classical level, Matrix theory has 11 dimensional Lorentz covariance.

At this stage, one might be puzzled by the following remark: the model used to describe the dynamics of the D0-branes, i.e. the 0+1 dimensional SYM, is derived in the string theory context in the limit in which $g \rightarrow 0$ and $m_s \rightarrow \infty$, where $m_s = 1/l_s = 1/\sqrt{\alpha'}$ is the string mass. Now if the D0-branes are identified with the partons of M-theory in the IMF (or in the DLCQ), the relations between g and m_s on one side and $R \equiv R_{11}$ and $m_p = 1/l_p$ (the 11 dimensional Planck mass) on the other are:

$$g = (Rm_p)^{\frac{3}{2}}, \quad m_s = (Rm_p^3)^{\frac{1}{2}}. \quad (2.49)$$

These are simply the relations (2.27) but expressed the other way round. In the end, when one is making contact with the original uncompactified M-theory that the parton model is supposed to describe, one has to take the limit $R \rightarrow \infty$ while keeping m_p fixed. This seems to imply that it corresponds to taking the D0-brane theory in the limit $m_s \rightarrow \infty$ and $g \rightarrow \infty$, in apparent contradiction with the limit in type IIA theory discussed above. Moreover, when one is actually doing computations with the parton model, one generally wants to take m_p and R fixed at a finite value. This also implies finite values for g and m_s , and not the limiting values which give the SYM theory. How can the Matrix theory conjecture be correct?

The argument which solves this apparent problem comes actually in two steps, each one of which has been emphasized respectively by Sen [69] and by Seiberg [70] in two contemporary papers in fall 1997, one year after the original Matrix theory proposal.

The first step has been essentially analyzed by Sen, and amounts to finding the exact limit in which the SYM quantum mechanics of the D0-branes is a good approximation, but nevertheless yields non-trivial dynamics. In other words, one is searching for the limit in which the D0-branes decouple from the rest of the type IIA string theory, but still produce an interacting theory. As in the two papers [69, 70], we also take the occasion to discuss the case in which some of the transverse directions are compact.

Let us specialize the general action for a Dp -brane (2.41)–(2.42) to our case of interest, that is the action for the D0-branes. In this case, actually one can fix the gauge in such a way that only the scalars survive. The residual gauge symmetry then enters through a constraint on the dynamics of the model. The bosonic action we are left with is:

$$I_{D0} = \frac{1}{gm_s^3} \int dt \left(\frac{1}{2} \text{tr} \dot{\phi}_i^2 - \frac{1}{4} \text{tr} [\phi_i, \phi_j]^2 \right). \quad (2.50)$$

We are now at the core of the problem of how the dynamics of the D0-branes decouples from the rest of the physics of type IIA string theory. For the action above to be a good description of the D0-brane dynamics, we have to take the limits

$$g \rightarrow 0, \quad m_s \rightarrow \infty \quad (2.51)$$

in the string theory. However for the same action to yield non-trivial informations, one has to keep fixed the value of the coupling constant $g_{YM_{0+1}}$, and the scale given by the expectation values of the scalars. This second requirement implies that if there are some compact directions, we want the scale of the interval over which the ϕ_i are defined to be fixed. If the compact directions have radii \tilde{R}_i , then the scale to be fixed is $\tilde{R}_i m_s^2$.

Since $g_{YM_{0+1}}^2 = g m_s^3$, fixing the SYM coupling constant is consistent with the limit (2.51). In order to keep $\tilde{R}_i m_s^2$ fixed in the same limit, we see that we must take $\tilde{R}_i \rightarrow 0$, and actually \tilde{R}_i has to take substringy values, $\tilde{R}_i \ll l_s$.

Let us now view the string theory parameters as coming from 11 dimensional ones, which we call for the moment R_s and \tilde{m}_p to distinguish them with the ones which define the Matrix model, R and m_p , and which must be finite. From (2.27) we see that the limit (2.51) implies:

$$R_s \rightarrow 0, \quad \tilde{m}_p \rightarrow \infty. \quad (2.52)$$

The SYM coupling however is the only combination of these two quantities which enters in the action (2.50), and is given by:

$$g_{YM_{0+1}}^2 = g m_s^3 = (R_s \tilde{m}_p^2)^3.$$

The quantity to be fixed is thus $R_s \tilde{m}_p^2$. As for the compact directions, one can also see that:

$$\tilde{R}_i m_s^2 = \tilde{R}_i R_s \tilde{m}_p^3 = (R_s \tilde{m}_p^2) \tilde{R}_i \tilde{m}_p,$$

so that one fixes $\tilde{R}_i \tilde{m}_p$, the transverse radii in 11 dimensional Planck units, when taking the limit (2.52).

Sen argues in his paper [69] that the Matrix theory conjecture amounts to postulating that the same action (2.50) describes the physics of the DLCQ partons of M-theory with finite parameters R , m_p and R_i , provided one makes the following identifications:

$$R_s \tilde{m}_p^2 = R m_p^2, \quad (2.53)$$

$$\tilde{R}_i \tilde{m}_p = R_i m_p. \quad (2.54)$$

Seiberg has actually shown in his paper [70] how the two pictures are physically related. The basic result is that the theory in the DLCQ is equivalent, by a large boost, to the same theory in the IMF and compactified on a vanishingly small circle.

Let us see, following Seiberg, how we can reach the DLCQ formulation by boosting a theory compactified on a space-like circle. The ‘space-like’ theory is defined by the following identifications:

$$\tilde{t} \sim \tilde{t}, \quad \tilde{x}_{11} \sim \tilde{x}_{11} + R_s. \quad (2.55)$$

When we perform a boost like (2.44), the identifications on the boosted coordinates become:

$$t \sim t - \sinh s R_s, \quad x_{11} \sim x_{11} + \cosh s R_s.$$

This gives for the light like coordinate $x^- \sim x^- + e^s R_s$. Choosing the boost parameter to be:

$$e^s = \frac{R}{R_s}, \quad (2.56)$$

we recover the DLCQ identifications in the limit $R_s \rightarrow 0$:

$$t \sim t - \frac{1}{2}R, \quad x_{11} \sim x_{11} + \frac{1}{2}R, \quad (2.57)$$

with R finite. Note that in this limit the velocity of the boost is:

$$v = \tanh s = 1 - 2\frac{R_s^2}{R^2} + \dots$$

We have thus shown that the theory we started with has to be compactified on a vanishingly small space-like circle.

Remember now that we had already argued that it was still possible in the DLCQ to perform boosts in the lightcone direction, which amounted to rescale the radius by $R \rightarrow \lambda R$, and accordingly the DLCQ momentum and energy by $p_- \rightarrow \frac{1}{\lambda}p_-$ and $p_+ \rightarrow \lambda p_+$ respectively. This means that p_- correctly scales as $1/R$, and that the typical DLCQ energy must scale linearly in R , i.e. it is of the order:

$$p_+ \sim Rm_p^2, \quad (2.58)$$

as suggested also by (2.48).

When we ‘undo’ the boost from the light cone set up (2.57) to the space-like compactification (2.55), the momentum and the energy are accordingly rescaled by the factor (2.56). This gives at leading order in R_s the following behaviour:

$$|\tilde{p}_{11}| \simeq \frac{R}{R_s}p_- = \frac{N}{R_s},$$

$$\tilde{p}_+ = \tilde{p}_0 - |\tilde{p}_{11}| \equiv E = \frac{R_s}{R}p_+ \sim R_s m_p^2. \quad (2.59)$$

We thus observe that all the relevant energies of the DLCQ formulation are mapped to vanishingly small energies in the space-like set up (2.55).

The benefit of considering the original DLCQ M-theory in the equivalent formulation (2.55) is that now the compactification gives perturbative string theory. Since however we are keeping m_p fixed, the string theory parameters given by (2.49) are such that $g \rightarrow 0$ but also $m_s \rightarrow 0$. This is not what we really want in order to make contact with the D0-brane action (2.50). Note however that from (2.59) we learn that we are actually focusing on a very particular range of energies.

One way to simplify the problem is the following. Instead of fixing the Planck scale and looking at very small energies, we can take the energies to be finite and rescale the Planck mass to be very large. In doing this, we are actually changing the theory in which we are working, in some ways we are transposing the problem to an auxiliary M-theory. There is however a very controlled way to do this transposition. We simply choose the rescaled Planck mass \tilde{m}_p in such a way that the energies in the auxiliary theory are matched to the energies in the DLCQ of the original theory:

$$\tilde{E} \sim R_s \tilde{m}_p^2 = Rm_p^2 \sim p_+.$$

Also we want the transverse geometry to be invariant under the rescaling, so that if there are compact directions with radii R_i in the original theory, these are rescaled to the values \tilde{R}_i which satisfy $\tilde{R}_i \tilde{m}_p = R_i m_p$.

We are obviously repeating the procedure which led to the identifications (2.53)–(2.54), but we have now a clear picture of how the two theories are related and why should they be equivalent.

Let us recall the steps which lead to Seiberg's formulation of Matrix theory and which we have discussed above:

- We start with M-theory in a DLCQ formulation, with finite parameters R , m_p and R_i , and with energies which scale as $p_+ \sim R m_p^2$.
- We can see the preceding formulation as the infinite boost limit of the same M-theory (thus characterized by the same m_p and R_i), but this time compactified on a space-like circle with radius R_s which is vanishingly small, $R_s \rightarrow 0$. The energies are also very small, since they now scale like $E \sim R_s m_p^2$.
- The relevant states of the M-theory of the previous step are identified to the states of a new, auxiliary \tilde{M} -theory, with parameters \tilde{m}_p and \tilde{R}_i satisfying (2.53)–(2.54). The energies are now finite.
- Since now in \tilde{M} -theory we have $R_s \rightarrow 0$ and $\tilde{m}_p \rightarrow \infty$, we can turn to the string theory picture and describe the dynamics of the partons by the D0-brane mechanics in the limit $g \rightarrow 0$, $m_s \rightarrow \infty$, which is correctly given by the action (2.50).

Using the relations (2.49) together with (2.53)–(2.54), we can now precisely derive the behaviour of the parameters in the string theory in which the D0-branes live when, say, $R_s \rightarrow 0$. For the string coupling and the string mass, we find:

$$g = (R_s \tilde{m}_p)^{\frac{3}{2}} = R_s^{\frac{3}{4}} (R m_p^2)^{\frac{3}{4}}, \quad (2.60)$$

$$m_s = (R_s \tilde{m}_p^3)^{\frac{1}{2}} = R_s^{-\frac{1}{4}} (R m_p^2)^{\frac{3}{4}}. \quad (2.61)$$

The radii of the transverse directions behave like:

$$\tilde{R}_i = \frac{1}{\tilde{m}_p} R_i m_p = R_s^{\frac{1}{2}} \left(\frac{R_i^2}{R} \right)^{\frac{1}{2}}, \quad (2.62)$$

and thus are smaller than the string length $l_s = 1/m_s$, as already noted.

What the dualities have taught us to do in this case, is to perform a T-duality on all the compact directions (we now assume that M-theory was compactified on a p -torus T^p). The string scale (2.61) is not changed under T-duality. On the other hand, the radii of the dual torus are given by (2.17), which translates here into:

$$\tilde{R}_i \rightarrow \Sigma_i = \frac{1}{\tilde{R}_i m_s^2} = \frac{1}{R_i R m_p^3}. \quad (2.63)$$

The dual torus T'^p has thus a finite volume. Note that the finite values of the Σ_i are the inverse of the scale of the vacuum expectation values of the scalars in (2.50), which we wanted to keep fixed.

Under T-duality, the string coupling also changes according to (2.19). Performing the transformation in the p directions, one obtains:

$$g \rightarrow g' = g \prod_i \frac{1}{\tilde{R}_i m_s} = R_s^{\frac{3-p}{4}} (R m_p^2)^{\frac{3-p}{4}} \prod_i (R_i m_p)^{-1}. \quad (2.64)$$

The behaviour of the dual string coupling thus depends on the number of compact directions. Note also that the p T-dualities change the D0-branes into D p -branes wrapped on the dual torus T'^p . We thus see that for $p \leq 3$, we have a theory of D-branes at weak or finite string coupling, but as soon as we consider compactifications with $p > 3$, we have a theory of D-branes at strong coupling. We postpone until Chapter 5 the detailed discussion on how to make sense of these D-brane theories embedded in a strongly coupled string theory. Indeed, string dualities and elevation to (yet another auxiliary) M-theory will allow us to give a proposal for the cases $p = 4, 5$ and 6.

We can still compute the coupling of the $p + 1$ dimensional SYM theory, using (2.42). We obtain:

$$g_{YM_{p+1}}^2 = g' m_s^{3-p} = (R m_p^2)^{3-p} \prod_i (R_i m_p)^{-1}. \quad (2.65)$$

This finite result, as the one for the dual radii (2.63), is a direct consequence of, and actually the motivation for, the prescription (2.53)–(2.54). Note also that the two finite values (2.65) and (2.63) are the same as the ones that would have been obtained if one had naively computed these values directly from the DLCQ set up, without caring about the limit (2.51).

As a last remark on the Matrix description of M-theory compactifications, we should point out that new light BPS states arise from the wrapping modes of the branes of M-theory. This is a non-trivial test for the Matrix model, because if it is to describe M-theory in the full non-perturbative regime, then all these modes have to be contained in Matrix theory compactifications as given by the prescription above. We will see in Chapter 5 that indeed the theories on the D-branes reproduce all these BPS states, at the price of being rather unusual theories for $p \geq 4$.

Let us conclude this section on Matrix theory by mentioning the problems that are still unresolved, and which might prevent it from being the favourite approach to M-theory.

A problem that has been addressed ever since the original proposal, but that nevertheless has still not been settled is the proof of 11 dimensional Lorentz invariance at the quantum level. More recently, there have been two papers stating that some of the predictions of Matrix theory did not reproduce some non-trivial but standard supergravity results, namely three graviton scattering [76] and two graviton scattering on a (weakly) curved background [77].

Another problematic issue is the background dependence of the model. For instance, if one wants to adapt the Matrix model to the presence of a M5-brane wrapped on the IMF or DLCQ direction (or longitudinal 5-brane), then in [78] it was shown that

the Hamiltonian has to be supplemented by new pieces. In the string picture, these new terms are related to the fields corresponding to the open strings connecting the D0-branes to the D4-brane (which is the longitudinal 5-brane in the IIA picture). A somewhat orthogonal approach was undertaken in [79], where the central charges of the supersymmetry algebra of the Matrix model were computed and shown to reproduce the charges of the membrane and of the longitudinal 5-brane. This result would thus imply that the description of the branes is contained in the Matrix model from the beginning. Yet another problem about the branes in Matrix theory is that contrary to the longitudinal one, the transverse 5-brane (i.e. an M5-brane not wrapping on the 11th direction) still remains very elusive.

2.7 No conclusion

This section is not a conclusion since there is obviously no definite answer to the question at the head of the chapter, ‘What is M-theory?’. We will instead try to collect and summarize the facts presented in this chapter which suggest that string theories are unified at the non-perturbative level under an 11 dimensional theory.

The 5 consistent superstring theories presented in Section 2.2 are all related to each other in the following way: type IIA and type IIB theories are related by T-duality, which acts at the perturbative level as soon as one of the 9 space-like directions is compact. These two theories can thus be seen as two different limits in the moduli space of compactifications of a unique, type II theory. Then we have seen that type I theory can be seen as type IIB theory with the addition of a set of D9-branes and an orientifold 9-plane. In this respect type I theory is ‘contained’ in type IIB theory, even if this requires the knowledge of the non-perturbative aspects of the latter, such as the physics of D-branes. Type I theory is further related to the $SO(32)$ heterotic string theory by a duality interchanging strong and weak coupling. To end with, again T-duality relates the two heterotic string theories as it did for the two type II theories. Further direct relations between, say, type II and heterotic string theories exist when these theories are compactified to lower dimensions, possibly on manifolds breaking some of the supersymmetries.

The above dualities already tell us that there is a unique theory with many perturbative string theory limits, with the fundamental string of one theory often being a complicated soliton in the dual theory. There are however two more dualities which relate 10 dimensional string theories to an 11 dimensional theory, the low energy effective action of which is 11 dimensional supergravity. For both type IIA string theory and the $E_8 \times E_8$ heterotic string theory, the strong coupling limit can be reinterpreted as the low-energy limit of a theory with one more dimension. The unique theory behind all string theories, i.e. M-theory, appears thus to be an 11 dimensional theory.

There is little we can say directly about M-theory. The main outcome of the string dualities is that its low-energy limit should be 11 dimensional supergravity. This is appealing since this latter theory is totally constrained and furthermore it is the endpoint of the supergravity program, i.e. no supergravity in more than 11 dimensions can be

constructed. In order to reproduce also the non-perturbative spectrum of the string theories, M-theory should contain some BPS extended objects, or branes. There is however little hope that M-theory could come out of the quantum theory of one of these objects. Leaving aside the technical reasons, it appears that in M-theory the choice between the M2-brane and the M5-brane as the fundamental object of the theory is somewhat arbitrary, simply because there is actually no perturbative coupling constant in M-theory, such as the string coupling constant g in string theories.

It should also be mentioned that M-theory has recently been shown to possess more exotic branes, such as the KK6 monopole (which is properly defined only when one of the directions is compactified on a circle) and the M9-brane, which is still at a very conjectural level (see [53] for a discussion), and which should be related to the Horava-Witten construction of M-theory compactified on the orbifold S^1/Z_2 . The M-branes have been a powerful tool to provide evidence for dualities involving directly (the 11 dimensional manifestation of) M-theory.

Matrix theory is a direct approach to M-theory, even if string theories are used as auxiliary theories in which to find the appropriate tools for this approach. It has however some serious flaws, such as the difficulty to recover all the BPS branes, and the indirect way in which the supergravity low-energy limit is recovered. For instance, it seems unlikely or very difficult to prove that Matrix theory could reproduce the equations of motion of 11 dimensional supergravity, as string theories do in 10 dimensions.

After all, string theories as we know them now, are very nice constructions, despite the lack of a second quantized string theory. In their almost 30 years of living, they have been cast in a well-defined set of operating rules which gives beautiful results. If there was only one consistent realistic string theory instead of five, the construction described in Section 2.2 is really what a theoretical physicist would call a unified theory based on first principles. What M-theory teaches us is that if non-perturbative effects are taken into account, this has a chance to be true, even if we have five (and maybe more) starting points.

The Second Superstring Revolution, besides promoting M-theory to the leading rôle, has also made the string theories much more powerful. The exploration of their non-perturbative aspects has been enormously facilitated by the accessibility of the physics of some of their solitons through D-brane physics, and by the dualities which allow to study the strong coupling limit of a string theory in terms of the perturbative regime of a dual theory. It is maybe this enhanced understanding of the non-perturbative aspects of string theories that can be generically called M-theory.

What we should not forget when discussing a theory unifying all interactions, is that one should indeed be able to make contact with ‘all the interactions’.

The aim of a unique theory of all the interactions should thus be to meet the basic facts that are accessible in everyday’s life. In other words, M-theory should produce a picture consistent with the world as we know it. A minimal requirement is that M-theory should predict that the world in which we live is 4 dimensional, and that the low-energy physics should be given by the Standard Model. Moreover, the constants entering in the Standard Model should also be predicted by M-theory. Now, string phenomenology, and its M-theory extensions, even if they can reproduce some of the features of the Standard

Model, suffer from an initial problem: in order to reproduce the phenomenology, one has to introduce many things by hand, first of all compactification to 4 dimensions, which includes supersymmetry breaking.

There are thousands of ways in which one can compactify string or M-theory to 4 dimensions, and one is generally reduced to study the more realistic model. The problem which is however much more difficult to address is how M-theory chooses the particular compactification which is closest to the Standard Model. String dualities and M-theory have led to the picture in which all the different compactifications are related. This would mean that all the vacua of M-theory are connected, and that it would thus be possible, in principle, to find the global minimum, or at least some (long-lived) local minima. Then showing that the Standard Model indeed corresponds to one of these minima should be the ultimate evidence that M-theory predicts the world as we see it. Present-day research is clearly still very far from this ultimate goal.

M-theory could well not be the Theory Of Everything, or at least we are still not convinced of it. The study of M-theory is however highly interesting in two respects. First of all, M-theory has a deep structure that still waits to be uncovered, and this can be approached from the several sides and aspects of the theory. Secondly, M-theory reveals itself as an incredibly powerful framework to address many other relevant problems in high energy physics.

Chapter 3

Classical p -brane solutions and their intersections

The aim of this chapter is to give a classical description of the solitons that are relevant to the study of the non-perturbative aspects of string theories and of M-theory. These solitons are identified with the classical p -brane solutions of the supergravities which are the low energy effective actions of the above-mentioned theories. A large part of the derivation of the p -brane solutions and of their intersections will be actually done in the more general setting of a purely bosonic theory with arbitrary couplings of the antisymmetric tensor fields to the dilaton. Then for some particular values of these couplings we will reduce to the cases of 10 and 11 dimensional supergravities. In these latter cases, one can perform an analysis of the supersymmetric properties of the solutions previously found, and sometimes find an alternative derivation of the solutions, the main drawback of this second approach being its lack of generality. Supersymmetry is however often helpful in motivating the form of the sought for solutions, and above all it is the main ingredient in allowing us to consider these solitons of the supergravities as solitons of the full string or M-theories.

The chapter is organized as follows. In Section 3.1 the problem of finding classical solutions of some supergravity theories is put in the perspective of addressing issues pertaining to the non-perturbative aspects of superstring theories. Section 3.2 is devoted to finding singly charged, spherically symmetric black p -brane solutions. This derivation is interesting with two respects: it will allow us to classify all the most simple p -brane solutions in 10 and 11 dimensional supergravities, and thus all the basic solitons of M-theory; we will use techniques and find some characteristics of these solutions that will guide us in finding more involved solutions carrying several charges. In Section 3.3 extreme solutions carrying several charges are found. The supersymmetry of the 11 dimensional solutions is also carefully examined. In addition, a different kind of solution carrying two charges is also derived, and it is shown to break all the supersymmetries. In the Section 3.4 we end with a discussion and with the review of some more general solutions that can be found.

The results of the paper [1] (see also [3]) are included in Section 3.3.

3.1 Why p -branes are interesting

The name ‘ p -brane’ actually covers a good deal of different objects. Most generally, it should be used to indicate, in the context of a theory containing gravity, a classical solution which is extended in p directions, i.e. which has p spacelike translational Killing vectors. In general p -branes will carry the charge of an antisymmetric tensor field, and those for which the mass saturates a lower bound given by the charge are called extremal. Black holes and, extrapolating to theories without gravity, magnetic monopoles are prototypes of p -branes for the case $p = 0$. As they are classical solutions, p -branes are often called ‘solitons’. This is essentially in view of the rôle they play in the quantum theory which is behind the classical theory of which they are solutions.

p -branes gain a lot of interesting properties when supersymmetry is a symmetry of the theory. When enough supersymmetry is present, we are allowed to consider that some relevant properties of the (extremal) p -brane solitons do not change when the coupling constant of the theory is varied. At strong coupling the mass, or the tension, of these objects becomes very small and the solitons become the fundamental quantum objects of a new, dual theory at strong coupling. The properties of the classical p -brane solutions can be used to identify in this way some of the properties of this new quantum objects and of their quantum theory. This is most interesting when one finds hints that the dual strong coupling theory is actually an already known and well-defined theory, thus giving a handle on the strong coupling dynamics of the original theory.

The first example of a theory for which a strong coupling formulation was found owing to the properties of its solitons was $N = 4$ Super Yang-Mills (SYM) theory in 4 dimensions. This theory has magnetic monopoles with mass $M_{mon} \sim M_W/g^2$, where g is the coupling constant of the SYM theory and M_W is the mass of the massive gauge bosons in the broken gauge symmetry phase. These monopoles preserve half of the supersymmetries of the $N = 4$ SYM and can be shown to fall into supermultiplets exactly similar to the ones to which belong the massive W gauge bosons [22]. The conjecture, following the ideas of [21], is that the strong coupling dynamics of $N = 4$ SYM is exactly reproduced by the same theory where however the coupling constant is the ‘magnetic’ charge $g_{dual} = 1/g$. The main ingredient of this conjecture is that $N = 4$ supersymmetry prevents the renormalization of the relation between the mass of the monopole and its magnetic charge, thus allowing its interpretation as a ‘magnetic’ massive gauge boson.

Although the interplay between supersymmetry and duality has produced incredible advances in the understanding of 4 dimensional gauge theories (since the work of Seiberg and Witten [80]), it is really in the higher dimensional theories that the pattern of solitonic extended objects becomes crucial in determining the properties of the underlying quantum theory.

The superstring theories, which have a critical dimension $D = 10$, have as low energy effective actions of their massless modes the $N = 1$ or $N = 2$ ten dimensional supergravities¹. The supergravity theories have a well-defined set of (abelian) antisymmetric

¹ Though non-renormalizable, these latter theories should give a fairly accurate description of the dynamics at energies well below the string scale, which is fixed by $\alpha'^{-1/2}$.

tensor fields with rank which can be taken to be ≤ 5 . To each of these n -form fields correspond two charged objects, one electric and the other magnetic.

Suppose we have in D dimensions an antisymmetric tensor field strength F_n of rank n , deriving from a potential such that $F_n = dA_{n-1}$. Then this potential can couple minimally to the world-volume of an object which thus has to be a p -brane with $p = n - 2$ spatial extended directions:

$$\mu_{n-2} \int_{W^{n-1}} A_{n-1}. \quad (3.1)$$

Here μ_p denotes the charge of the p -brane under the $(p + 2)$ -form field strength, and W^{p+1} is the world-volume swept by the p -brane. The term displayed in (3.1) is an electric coupling. However one can always define the Hodge dual of the form F_n which will be the $(D - n)$ -form field strength $\tilde{F}_{D-n} = *F_n$. Outside electric sources one can always define a dual potential by $\tilde{F}_{D-n} = d\tilde{A}_{D-n-1}$. This magnetic potential can now couple to a p' -brane, with $p' = D - n - 2$, by the following term:

$$\mu_{D-n-2} \int_{W^{D-n-1}} \tilde{A}_{D-n-1}. \quad (3.2)$$

To summarize, every n -form field strength should imply the existence of an electric p -brane and a magnetic p' -brane, with $p = n - 2$ and $p' = D - n - 2$. Note that $p + p' = D - 4$, and that the existence of both of these objects imposes a Dirac-like quantization condition on their charges [46, 47].

According to the remarkable conjecture of Hull and Townsend [20] and of Witten [17], all string theories, together with 11 dimensional supergravity, are related by dualities, and therefore all strong coupling limits are reformulated in terms of the perturbative dynamics of a dual theory [17]. The essential assumption leading to this conjecture is that the full non-perturbative duality group of a string theory in 10 dimensions or lower, coincides with the discretized version of the symmetry group of the corresponding supergravity. This enlarged duality, generalizing and unifying the strong-weak S-duality (much similar to the electric-magnetic duality in 4 dimensions) and the perturbative T-duality exchanging momentum and winding modes in the string spectrum, has been called U-duality [20] (see Section 2.4).

One of the main consequences of U-duality is to put on an equal footing all the gauge fields appearing in an effective (maximal) supergravity obtained by trivial (i.e. toroidal) compactification. These gauge fields arise from the off-diagonal parts of the metric and from partial compactification of the antisymmetric tensors of the higher dimensional supergravities. The relation between the gauge fields extends to their charged objects, and thus to their 10 (or 11) dimensional p -brane ascendents. All p -branes should then be considered on an equal footing, i.e. equally ‘potentially’ fundamental [42].

There is, in string theories, a subtlety in the solitonic ‘spectrum’, which is most easily seen in type II string theories. In the latter theories, the fields coming from the Ramond-Ramond (RR) sector have the peculiarity that there are no states carrying their charge in the perturbative string spectrum. However U-duality relates this sought for states to elementary string states, thus requiring the presence of the former. This means that solitons of the effective supergravities must exist, which correspond to both electric

and magnetic RR-charged states. Now classical U-duality in turn implies the existence of solitons carrying the same charge as some perturbative string states. The consistency of the U-duality picture imposes the identification of these classical solutions to the corresponding perturbative string states. In conclusion, all possible charged objects must exist as classical solutions in the effective supergravities.

One of the interests in finding all the p -brane solutions is thus to corroborate the duality conjectures and to properly find all the objects belonging to the U-duality multiplets. It is worth saying that there is by now overwhelming evidence for U-duality.

Even more interesting, searching for the way the classical solutions interact, for instance by building solutions with intersecting p -branes, is relevant for checking the consistency with, and sometimes deriving, the interactions of the quantum objects. It is also important that all these intersecting solutions realize the duality conjectures.

Another important issue involving intersections of p -branes is the search for supersymmetric extremal black holes with a non-vanishing horizon area. It will turn out that this is possible only in 4 and 5 non-compact spacetime (including time). Then going on to the identification of these p -brane configurations with systems of objects in perturbative string theory, i.e. mainly systems involving D-branes [26], it has been possible to give a microscopic counting of states strictly reproducing the semiclassical black hole entropy (see [61, 62] and the following, considerable, literature).

3.1.1 The starting point: the supergravity actions

We now go on to the formulation of the problem of finding p -brane solutions in classical supergravities.

We will be primarily interested in the p -branes of the maximal supergravities in 10 and 11 dimensions, because in principle all lower dimensional configurations can be constructed by trivial compactification. Note that compactifications which break some of the supersymmetries have a richer structure, and it is more involved to analyze which brane configurations they are compatible with. We will not consider such compactifications here.

Despite the differences between the supergravity theories discussed above, namely 11 dimensional supergravity, type IIA non-chiral and type IIB chiral 10 dimensional supergravities, searching for solutions of the equations of motion will be essentially the same. It is actually a matter of selecting particular values of some parameters. It is thus worthwhile attempting to formulate these equations of motion for a general theory, and then specialize to the particular cases of interest after the solutions have been found. At the end of this subsection we will present a general action in D dimensions, with general couplings to the dilaton and with an arbitrary number of antisymmetric tensor fields of arbitrary rank n , which can be straightforwardly reduced to any one of the cases above.

This generalization to arbitrary dimension, for instance, is of course possible because we will consider only a bosonic action. In the case of the supergravities, we thus consider the truncation to their bosonic sector. This is what is ordinarily done when one is searching for classical solutions, since in the latter the expectation values of the fermionic fields are generically taken to vanish. This fact actually allows for an alternative way to

find a particular class of solutions of the supergravity theories, namely solutions which preserve some of the supersymmetries: since all the fermionic fields must vanish, the condition of preserved supersymmetry reduces to the condition that the supersymmetric variation on the fermionic fields must vanish also, $\delta_{susy}\psi = 0$. However the equations one obtains in this way depend crucially on all the parameters of the theory, and thus cannot be straightforwardly generalized. In Section 3.3 we will perform this calculation in the context of 11 dimensional supergravity.

Let us now review the actions in $D=10$ and 11 and see how they can be captured in a single general action.

The bosonic part of the action of $D=11$ supergravity is the simplest one, because there are only two fields present: the metric and a 4-form field strength $F_4 = dA_3$. We use conventions for the n -forms such that $F_n = \frac{1}{n!}F_{\mu_1\dots\mu_n}dx^{\mu_1} \wedge \dots \wedge dx^{\mu_n}$. The action is the following [43]:

$$I_{11} = \frac{1}{16\pi G_{11}} \left(\int d^{11}x \sqrt{-g} \left\{ R - \frac{1}{48} F_4^2 \right\} + \frac{1}{6} \int A_3 \wedge F_4 \wedge F_4 \right). \quad (3.3)$$

The Newton constant G_{11} defines the 11 dimensional Planck length by $G_{11} = l_p^9$. The second term in (3.3) is the Chern-Simons-like term necessary for supersymmetry to hold [43]. We will disregard it in most of the rest of this chapter, but it can be checked ‘a posteriori’ that the solutions we will find are not affected by the reintroduction of this term.

The (bosonic part of the) action of $D=10$ type IIA supergravity [81] is, written in the so-called string frame (see for instance [7]):

$$I_{IIA} = \frac{1}{16\pi G_{10}} \int d^{10}x \sqrt{-g_S} \left\{ e^{-2\phi} \left(R_S + 4\partial_\mu\phi\partial^\mu\phi - \frac{1}{12}H_3^2 \right) - \frac{1}{4}F_2^2 - \frac{1}{48}F_4'^2 \right\} \\ + \frac{1}{16\pi G_{10}} \frac{1}{2} \int B_2 \wedge F_4 \wedge F_4. \quad (3.4)$$

Here we have used the following definitions: $H_3 = dB_2$, $F_2 = dA_1$, $F_4' = F_4 + A_1 \wedge H_3$ and $F_4 = dA_3$. The fields which have a factor of $e^{-2\phi}$ in front of their kinetic term are those arising from the Neveu-Schwarz-Neveu-Schwarz (NSNS) sector of the closed string, while the others arise from the Ramond-Ramond (RR) sector. The 10 dimensional Newton constant is defined in terms of the string tension $G_{10} \sim \alpha'^4$ and we recall that $e^{\phi_\infty} = g$, the string coupling constant. It is in general convenient to subtract from the dilaton field its constant part at infinity ϕ_∞ and to insert it into the value of G_{10} , which becomes $G'_{10} \sim g^2\alpha'^4$. If we want to keep a unique coupling constant in front of the whole action, we also have to rescale the RR fields. The new dilaton $\phi' = \phi - \phi_\infty$ is thus vanishing at infinity. It is always this dilaton and Newton constant that we will use hereafter, and we will accordingly drop the primes.

We now rewrite the action (3.4) in the Einstein frame, where the gravitational term is canonical. In order to do this we have to perform a Weyl rescaling of the metric:

$$g_{\mu\nu}^S = e^{\frac{\phi}{2}} g_{\mu\nu}^E. \quad (3.5)$$

Note that, following the discussion of the previous paragraph, this Einstein metric actually coincides with the string metric at infinity. Using the formula (A.10) of Appendix A, we can rewrite the action in the following way:

$$I_{IIA} = \frac{1}{16\pi G_{10}} \int d^{10}x \sqrt{-g_E} \left\{ R_E - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{12} e^{-\phi} H_3^2 - \frac{1}{4} e^{\frac{3}{2}\phi} F_2^2 - \frac{1}{48} e^{\frac{1}{2}\phi} F_4'^2 \right\} + \frac{1}{16\pi G_{10}} \frac{1}{2} \int B_2 \wedge F_4 \wedge F_4. \quad (3.6)$$

It can now be seen, using again Appendix A, that this action is the dimensional reduction of the 11 dimensional action (3.3). This is actually the simplest way to determine the action of type IIA supergravity.

Let us now turn to the action of the chiral type IIB theory in $D=10$. In the string frame it reads:

$$I_{IIB} = \frac{1}{16\pi G_{10}} \int d^{10}x \sqrt{-g_S} \left\{ e^{-2\phi} \left(R_S + 4\partial_\mu \phi \partial^\mu \phi - \frac{1}{12} H_3^2 \right) - \frac{1}{2} \partial_\mu \chi \partial^\mu \chi - \frac{1}{12} F_3'^2 - \frac{1}{240} F_5'^2 \right\} + \frac{1}{16\pi G_{10}} \int A_4 \wedge F_3 \wedge H_3. \quad (3.7)$$

Here the RR fields have the following definitions: $F_3' = F_3 - \chi H_3$, $F_3 = dA_2$ and $F_5' = dA_4 + A_2 \wedge H_3$. To recover the true equations of motion of type IIB supergravity [39], one has to impose a self-duality condition on the 5-form field strength: $*F_5' = F_5'$. Strictly speaking, the equations of motion are thus not derived from the action (3.7) alone. It will be however convenient in the following to work with such an action, and to impose the self-duality condition only at the end (actually on the solutions).

If we go to the Einstein frame with the same transformation above (3.5), we obtain:

$$I_{IIB} = \frac{1}{16\pi G_{10}} \int d^{10}x \sqrt{-g_E} \left\{ R_E - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{12} e^{-\phi} H_3^2 - \frac{1}{2} e^{2\phi} \partial_\mu \chi \partial^\mu \chi - \frac{1}{12} e^\phi F_3'^2 - \frac{1}{240} F_5'^2 \right\} + \frac{1}{16\pi G_{10}} \int A_4 \wedge F_3 \wedge H_3. \quad (3.8)$$

We can now easily write an action in D dimensions which generalizes the ones above (3.3), (3.6) and (3.8). However, in order to do this we have to neglect all the Chern-Simons-like terms in the actions above and in the definitions of the field strengths. Luckily, they will turn out to be completely irrelevant in the case of the solutions we will be searching for.

Consider a theory including gravity, a dilaton field and \mathcal{M} antisymmetric tensor fields of arbitrary rank n_I , $I = 1 \dots \mathcal{M}$. Then the most general action, written in the Einstein frame, is the following [1, 3]:

$$I = \frac{1}{16\pi G_D} \int d^D x \sqrt{-g} \left\{ R - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} \sum_I \frac{1}{n_I!} e^{a_I \phi} F_{n_I}^2 \right\}. \quad (3.9)$$

Note that some of the forms can be of the same rank, with nevertheless a different coupling to the dilaton a_I which distinguishes them.

The reduction to the cases discussed above is straightforward: for $D=11$ supergravity, we only have a 4-form and $a_4 = 0$, which means that the dilaton is irrelevant in this theory. For the type II supergravities in $D=10$, we have $a_3 = -1$ for the NSNS 3-form, and $a_n = \frac{5-n}{2}$ for any RR n -form.

We can also specialize to the other string theories: the low energy effective action of type I string theory can be obtained with the same value of the coupling for the RR 3-form, and with a coupling $a_2 = \frac{1}{2}$ for the 2-form field strengths coming from the open sector (we only consider the 16 fields relative to the abelian subgroup). This coupling derives from the fact that the latter fields are multiplied, in the string frame action, by a factor of $e^{-\phi}$, characteristic of the disk amplitude. For the heterotic string theories, all the fields arise from a closed NSNS-like sector and the couplings are thus $a_3 = -1$ for the 3-form and $a_2 = -\frac{1}{2}$ for the 16 (abelian) 2-forms.

Note that if we wanted to reproduce also lower dimensional supergravities, we should have added more than one scalar, i.e. instead of only one dilaton we should have had several moduli. However all the solutions in higher dimensions can be trivially compactified to reproduce solutions in lower dimensions. Though straightforwardly generalizable, the action (3.9) is a compromise in favour of the treatability of the equations and of the readability of their solutions.

The equations of motion one derives from (3.9) are:

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R - \frac{1}{2}\left(\partial_\mu\phi\partial_\nu\phi - \frac{1}{2}g_{\mu\nu}\partial_\lambda\phi\partial^\lambda\phi\right) - \frac{1}{2}\sum_I \frac{1}{n_I!}e^{a_I\phi}\left(n_I F_{\mu\lambda_2\dots\lambda_{n_I}}F_{\nu\lambda_2\dots\lambda_{n_I}} - \frac{1}{2}g_{\mu\nu}F_{n_I}^2\right) = 0, \quad (3.10)$$

$$\frac{1}{\sqrt{-g}}\partial_\mu\left(\sqrt{-g}g^{\mu\nu}\partial_\nu\phi\right) - \frac{1}{2}\sum_I \frac{a_I}{n_I!}e^{a_I\phi}F_{n_I}^2 = 0, \quad (3.11)$$

$$\frac{1}{(n-1)!}\frac{1}{\sqrt{-g}}\partial_\mu\left(\sqrt{-g}e^{a_I\phi}F^{\mu\nu_2\dots\nu_{n_I}}\right) = 0. \quad (3.12)$$

They are better recast in the following form:

$$R^\mu{}_\nu = \frac{1}{2}\partial^\mu\phi\partial_\nu\phi + \frac{1}{2}\sum_I \frac{1}{n_I!}e^{a_I\phi}\left(n_I F^{\mu\lambda_2\dots\lambda_{n_I}}F_{\nu\lambda_2\dots\lambda_{n_I}} - \frac{n_I-1}{D-2}\delta^\mu_\nu F_{n_I}^2\right), \quad (3.13)$$

$$\square\phi = \frac{1}{2}\sum_I \frac{a_I}{n_I!}e^{a_I\phi}F_{n_I}^2, \quad (3.14)$$

$$\partial_\mu\left(\sqrt{-g}e^{a_I\phi}F^{\mu\nu_2\dots\nu_{n_I}}\right) = 0. \quad (3.15)$$

The statement that the n -form derives from a potential can be replaced by the imposition, at the level of the equations of motion, of the Bianchi identities:

$$\partial_{[\mu_1}F_{\mu_2\dots\mu_{n_I+1]}} = 0. \quad (3.16)$$

The four sets of equations above, (3.13)–(3.16), are the ones which we will use extensively in this chapter to find the p -brane solutions.

3.2 Singly charged extremal and black p -branes

In this section we focus on a particular case of the action (3.9), where there is a single n -form field strength. In this framework, we can implement several symmetries in the problem, which will allow us to constrain the fields in such a way that it will be possible to find a solution to the equations derived from the action. The solutions found in this way will then turn out to be useful guides when we will search for more general solutions in problems with less symmetries.

The material presented in this section is well-known, and one can find most of these results in [82, 83, 45, 84]. We have included it because it is very useful to be familiar with the single brane solutions before considering the configurations with several branes, i.e. the multiply charged solutions. The presentation followed here is however original, and takes into account the structure of the whole chapter.

We can rewrite the equations of motion and Bianchi identities for this simpler case, containing only gravity, a n -form field strength and the dilaton:

$$R^\mu{}_\nu = \frac{1}{2} \partial^\mu \phi \partial_\nu \phi + \frac{1}{2 n!} e^{a\phi} \left(n F^{\mu\lambda_2 \dots \lambda_n} F_{\nu\lambda_2 \dots \lambda_n} - \frac{n-1}{D-2} \delta^\mu_\nu F_n^2 \right), \quad (3.17)$$

$$\square \phi = \frac{a}{2 n!} e^{a\phi} F_n^2, \quad (3.18)$$

$$\partial_\mu \left(\sqrt{-g} e^{a\phi} F^{\mu\nu_2 \dots \nu_n} \right) = 0, \quad (3.19)$$

$$\partial_{[\mu_1} F_{\mu_2 \dots \mu_{n+1}]} = 0. \quad (3.20)$$

Let us note that the equations above have an interesting property. One can reformulate them in terms of a Hodge dual $(D-n)$ -form field strength. However to correctly have a generalized electric-magnetic duality, the definition of the Hodge dual must contain a dilaton dependent factor. A good definition of the dual field strength is thus the following:

$$\sqrt{-g} e^{a\phi} F^{\mu_1 \dots \mu_n} = \frac{1}{(D-n)!} \epsilon^{\mu_1 \dots \mu_n \nu_1 \dots \nu_{D-n}} \tilde{F}_{\nu_1 \dots \nu_{D-n}}, \quad \epsilon^{0 \dots D-1} = 1. \quad (3.21)$$

It is easy to see that using this definition of \tilde{F}_{D-n} , the equations (3.17)–(3.20) rewritten in terms of it have exactly the same form with however F_n replaced by \tilde{F}_{D-n} , n replaced by $D-n$ and a replaced by $-a$ (or equivalently ϕ replaced by $-\phi$). In this process, the Bianchi identity for F_n becomes the ‘Maxwell-like’ equation for \tilde{F}_{D-n} and vice-versa. Note also that the replacement $a \rightarrow -a$ is valid everywhere since F_n^2 is proportional to $-\tilde{F}_{D-n}^2$.

The electro-magnetic duality described above is a symmetry of the equations of motion supplemented by the Bianchi identity, but not of the action. We shall exploit this duality to give a unified treatment of electrically and magnetically charged p -branes.

3.2.1 The symmetries of the p -brane ansatz and the general equations of motion

We now consider the symmetries of the problem, in order to simplify the field equations above by restricting us to some particular field configurations.

A p -brane solution living in a D dimensional space-time is characterized by the fact that p space-like directions can be considered ‘longitudinal’ to the brane, while the remaining $d \equiv D - p - 1$ space-like directions are taken to be ‘transverse’ to the p -brane. Since we are considering solutions carrying a single charge, and thus by definition composed of a single p -brane, it is natural to suppose that the p longitudinal directions are all equivalent, i.e. there is nothing in the solution that distinguishes between them. The timelike direction, which can be considered as being also longitudinal to the world-volume of the brane, will not be considered as equivalent to the other longitudinal directions, at least in the general case.

Since the p -brane is taken to be a uniform object, i.e. nothing should single out particular points or regions of its volume, the p space-like longitudinal directions define as many translational invariant directions of the solution. When one considers static objects, as we will do here, another translational invariant direction is the timelike one.

In the transverse space, the p -brane is taken to be localized at a particular point. Invariance under translations is thus broken. If however the brane is truly static and not endowed with any angular momentum, it is natural to postulate spherical symmetry in the d dimensional transverse space. These are the maximal symmetries we can postulate for a single p -brane ansatz.

Let us start with a general metric in D dimensions, namely:

$$ds^2 = g_{\mu\nu} dz^\mu dz^\nu, \quad \mu, \nu = 0 \dots D - 1. \quad (3.22)$$

Having in mind the discussion above, we split the coordinates accordingly in three sets, $z^\mu = \{t, y^i, x^a\}$, with $i = 1 \dots p$ and $a = 1 \dots d$ (thus $1 + p + d = D$). The y ’s span the directions longitudinal to the brane, and the x ’s span the transverse space. We have singled out the time-like direction t .

It is now easy to translate the discussion above on the expected symmetries of the solutions in terms of Killing vectors of the p -brane geometry.

We can define the following sets of Killing vectors:

- The geometry should be invariant under time translations, i.e. it should describe a static configuration. The corresponding Killing vector is $\xi_t = \partial_t$ (which in components means $(\xi_t)^\mu = \delta_0^\mu$).
- There should be invariance under translations in the y^i directions, along the extension of the brane. The Killing vectors are $\xi_i = \partial_i$.
- There should be invariance under $SO(d)$ rotations in the transverse space of the x ’s. The Killing vectors are $\xi_{ab} = x^a \partial_b - x^b \partial_a$, or in components $(\xi_{ab})^\mu = x^a \delta_b^\mu - x^b \delta_a^\mu$.
- Invariance under $SO(p)$ rotations in the longitudinal space leads to the Killing vectors $\xi_{ij} = y^i \partial_j - y^j \partial_i$.

A word of caution has to be said about the last set of Killing vectors. Generally, we would like eventually to be able to compactify the longitudinal space, and take it to have the topology of a p -torus T^p . In this case these last Killing vectors can only be defined locally but not globally. It is however useful to assume that they exist, at the price of considering at a first stage truly infinite branes.

The statement that a vector field is a Killing vector means that the Lie derivative of any tensor field along this vector field must vanish. The Lie derivative along ξ^μ is defined by:

$$\begin{aligned} \mathcal{L}_\xi T_{\mu_1 \dots \mu_r}^{\nu_1 \dots \nu_s} &= T_{\mu_1 \dots \mu_r}^{\nu_1 \dots \nu_s} \xi^\rho{}_{,\rho} + T_{\rho \mu_2 \dots \mu_r}^{\nu_1 \dots \nu_s} \xi^\rho{}_{,\mu_1} + \dots + T_{\mu_1 \dots \mu_{r-1} \rho}^{\nu_1 \dots \nu_s} \xi^\rho{}_{,\mu_r} \\ &\quad - T_{\mu_1 \dots \mu_r}^{\rho \nu_2 \dots \nu_s} \xi^{\nu_1}{}_{,\rho} - \dots - T_{\mu_1 \dots \mu_r}^{\nu_1 \dots \nu_{s-1} \rho} \xi^{\nu_s}{}_{,\rho}. \end{aligned} \quad (3.23)$$

The translational Killing vectors ξ_t and ξ_i , being constant vector fields, imply as expected that an invariant tensor (and particularly the metric tensor $g_{\mu\nu}$) should not depend on the t and y^i coordinates respectively.

The Killing vectors of $SO(p)$ symmetry ξ_{ij} , because of the simultaneous presence of the ξ_i symmetries, are very restrictive. In particular, they imply that an invariant symmetric 2-index tensor must be proportional to the identity δ_{ij} , and that the only invariant antisymmetric tensor is the one proportional to the Levi-Civita p -form $\epsilon_{i_1 \dots i_p}$.

The ξ_{ab} Killing vectors of the $SO(d)$ symmetry in the transverse space give rise to more complicated equations since the tensor components still depend on the x^a coordinates.

For a 2-index symmetric tensor these equations can be solved without much effort and when we apply the result to the metric we obtain:

$$\begin{aligned} ds^2 &= -B^2(r)dt^2 + C^2(r)\delta_{ij}dy^i dy^j + h(r)\delta_{ab}x^a dx^b dt + k(r)\delta_{ab}dx^a dx^b \\ &\quad + j(r)\delta_{ac}\delta_{bd}x^a x^b dx^c dx^d, \quad r^2 = \delta_{ab}x^a x^b. \end{aligned} \quad (3.24)$$

We can now go to spherical coordinates, for which:

$$\delta_{ab}dx^a dx^b = dr^2 + r^2 d\Omega_{d-1}^2, \quad d\Omega_{d-1}^2 = d\theta_1^2 + \sin^2 \theta_1 d\theta_2^2 + \dots + \sin^2 \theta_1 \dots \sin^2 \theta_{d-2} d\theta_{d-1}^2.$$

The precise change of coordinates and the range of the angular variables is given in the Appendix C, eqs. (C.29)–(C.31). Note also that $r dr = \delta_{ab}x^a dx^b$. As far as the metric is concerned, we can still use diffeomorphisms to put the metric in a simpler form. We can indeed redefine the t coordinate adding to it an r -dependent piece in order to eliminate the g_{tr} element of the metric.

We eventually obtain the following metric, which is the most general one that incorporates all the symmetries discussed above:

$$ds^2 = -B^2 dt^2 + C^2 \delta_{ij} dy^i dy^j + F^2 dr^2 + G^2 r^2 d\Omega_{d-1}^2, \quad (3.25)$$

with all the functions depending only on r .

We can still perform reparametrizations of r , which means that the two functions $F(r)$ and $G(r)$ introduce a redundancy. However it depends from case to case whether

one gauge choice (e.g. $F = G$ or $G = 1$) is better than the other, and actually we will discuss a case for which it is better not to fix the ‘ r -gauge’ at all.

It is worth making a remark now that will only be justified in Section 3.3, when going to the discussion of solutions charged under several tensor fields. Even in cases where the $SO(p)$ symmetry of the ‘internal’ space can no longer be invoked, we will still postulate that the metric is diagonal (this being now truly an ansatz on the field configuration), taking a form which is a slight generalization of (3.25):

$$ds^2 = -B^2 dt^2 + \sum_{i=1}^p C_i^2 (dy^i)^2 + F^2 dr^2 + G^2 r^2 d\Omega_{d-1}^2. \quad (3.26)$$

We will actually use this slightly more general metric to compute the Ricci tensor in Appendix C and the thermodynamics of general black p -branes in this section.

We now go on characterizing which components of the n -form field strength are invariant under the Killing vectors and are relevant to the p -brane solutions.

One is guided intuitively by the simplest case, which is the Reissner-Nordström charged black hole in 4 dimensions. In this problem it is quite easy to constrain the form of the 2-form Maxwell field strength. Requiring time translation invariance and $SO(3)$ rotational symmetry, the only surviving components are, in spherical coordinates, $F_{tr} = \mathcal{E}(r)$ and $F_{\theta\varphi} = \mathcal{B}(r)r^2 \sin\theta$. This gives in cartesian coordinates, $F_{ta} = \partial_a E(r)$ and $F_{ab} = \epsilon_{abc} \partial_c \tilde{E}(r)$ respectively, where $E' = \mathcal{E}$ and $\tilde{E}' = \mathcal{B}$.

What we learn in this case is that as far as the physically interesting cases are concerned, either only one index relative to the x^a coordinates appears in the electric case, either all but one of the same indices appear in the magnetic case, the tensor now being proportional to the Levi-Civita d -form ($d = 3$ in the Reissner-Nordström case above).

If we recall that by $SO(p)$ and translation invariance in the internal space either all the y^i indices appear or none, we are thus led to formulate the following two ansätze for a n -form field strength. A $(p+2)$ -form will be said to satisfy an electric ansatz if it is of the form:

$$F_{ti_1 \dots i_p a} = \epsilon_{i_1 \dots i_p} \partial_a E(r). \quad (3.27)$$

It is straightforward to check that with such an ansatz the Bianchi identities (3.20) are trivially satisfied. The expression (3.27) can be rewritten in spherical coordinates as $F_{ti_1 \dots i_p r} = \epsilon_{i_1 \dots i_p} E'(r)$.

A magnetic ansatz is satisfied by a $(d-1)$ -form of the form:

$$F^{a_1 \dots a_{d-1}} = \epsilon^{a_1 \dots a_d} \frac{1}{\sqrt{-g}} e^{-a\phi} \partial_{a_d} \tilde{E}(r). \quad (3.28)$$

Such a $(d-1)$ -form satisfies trivially its equations of motion (3.19), as it should be to correctly represent a magnetic charge. In spherical coordinates, the only non-vanishing component is $F_{\theta_1 \dots \theta_{d-1}} \sim f(r) \sin^{d-2} \theta_1 \dots \sin \theta_{d-2}$.

It can indeed be checked that the n -forms with only non-vanishing components satisfying either the electric (3.27) or the magnetic (3.28) ansatz are invariant under all the

Killing vectors discussed above, and in particular under the ones related to the $SO(d)$ spherical symmetry.

If a n -form is such that $n = D/2$, it can have at the same time non-vanishing electric and magnetic components. Moreover, if the form is (anti-)self-dual, the two ansätze above single out the same component.

The requirement that an electric component of a n -form field strength must be proportional to $\epsilon_{i_1 \dots i_p}$ ceases to hold if there is no longer $SO(p)$ symmetry in the internal space. The absence of the ξ_{ij} Killing vectors will allow us in Section 3.3 to consider both electric and magnetic components depending on an arbitrary number of y^i indices.

It is important to note that once the n -form field is fixed to correspond to the expression (3.27) or (3.28), then the other equation linear in the n -form (the equation of motion in the electric case and the Bianchi identity in the magnetic one) can always be solved in the first place.

Let us end this subsection rewriting the field equations, taking into account the symmetries of the problem. We will do it for an electrically charged p -brane, and thus for a $(p+2)$ -form field strength, but the equations we will obtain will also be valid for a p -brane carrying the magnetic charge of a $(d-1)$ -form, by virtue of the relation (3.21) and the discussion that follows it.

In spherical coordinates, the only non-vanishing components of the $(p+2)$ -form are $F_{ti_1 \dots i_p r} = \epsilon_{i_1 \dots i_p} E'$. Taking into account the metric (3.25), the remaining equation (3.19) for the antisymmetric tensor becomes:

$$\left(e^{a\phi} \frac{(Gr)^{d-1}}{BC^p F} E' \right)' = 0,$$

which gives the solution:

$$F_{ti_1 \dots i_p r} = \epsilon_{i_1 \dots i_p} BC^p F e^{-a\phi} \frac{Q}{(Gr)^{d-1}}, \quad (3.29)$$

where Q is an integration constant, and it is proportional to the electric charge (or alternatively to the magnetic one in what follows).

Using (3.29), we are now able to compute the various terms appearing on the r.h.s. of the equations (3.17) and (3.18). For instance, we have:

$$F_{p+2}^2 = (p+2)! F_{ty_1 \dots y_p r} F^{ty_1 \dots y_p r} = -(p+2)! e^{-2a\phi} \frac{Q^2}{(Gr)^{2(d-1)}}.$$

By virtue of the fact that the components in (3.29) are the only non-vanishing ones, we have the following relation:

$$F^{\mu\lambda_2 \dots \lambda_{p+2}} F_{\nu\lambda_2 \dots \lambda_{p+2}} = \bar{\delta}_\nu^\mu (p+1)! F_{ty_1 \dots y_p r} F^{ty_1 \dots y_p r} = -\bar{\delta}_\nu^\mu (p+1)! e^{-2a\phi} \frac{Q^2}{(Gr)^{2(d-1)}},$$

where $\bar{\delta}_\nu^\mu = \delta_t^\mu \delta_\nu^t + \delta_i^\mu \delta_\nu^i + \delta_r^\mu \delta_\nu^r$.

The last field we have to consider before writing the remaining equations of motion is the dilaton, which depends only on r by spherical symmetry. Note first of all that in

the r.h.s. of (3.17) the $\partial^\mu \phi \partial_\nu \phi$ term contributes only to the $({}^r_r)$ equation by the term $\frac{1}{F^2} \phi'^2$.

The D'Alembertian operator can be written in our set up in the following form:

$$\begin{aligned} \square \phi &= \frac{1}{\sqrt{-g}} \partial_r \left(\sqrt{-g} g^{rr} \partial_r \phi \right) = \frac{1}{BC^p F (Gr)^{d-1}} \left[BC^p F (Gr)^{d-1} \frac{1}{F^2} \phi' \right]' \\ &= \frac{1}{F^2} \left\{ \phi'' + \frac{d-1}{r} \phi' + \phi' [(\ln B)' + p(\ln C)' - (\ln F)' + (d-1)(\ln G)'] \right\}. \end{aligned}$$

We are now ready to collect all the above results and to write the Einstein and the dilaton equations for the metric (3.25) and the $(p+2)$ -form (3.29). The Einstein equations are totally diagonal in these coordinates, and we list hereafter the equations in the following order: $({}^t_t)$, $({}^{y_i}_{y_i})$, $({}^r_r)$, $({}^{\theta_\alpha}_{\theta_\alpha})$ and $\square \phi$.

$$\begin{aligned} \frac{1}{F^2} \left\{ -(\ln B)'' - (\ln B)' \left[(\ln B)' + p(\ln C)' - (\ln F)' + (d-1)(\ln G)' + \frac{d-1}{r} \right] \right\} \\ = -\frac{d-2}{2(D-2)} e^{-a\phi} \frac{Q^2}{(Gr)^{2(d-1)}}, \end{aligned} \quad (3.30)$$

$$\begin{aligned} \frac{1}{F^2} \left\{ -(\ln C)'' - (\ln C)' \left[(\ln B)' + p(\ln C)' - (\ln F)' + (d-1)(\ln G)' + \frac{d-1}{r} \right] \right\} \\ = -\frac{d-2}{2(D-2)} e^{-a\phi} \frac{Q^2}{(Gr)^{2(d-1)}}, \end{aligned} \quad (3.31)$$

$$\begin{aligned} \frac{1}{F^2} \left\{ -(\ln B)'' - p(\ln C)'' - (\ln B)'^2 - p(\ln C)'^2 + (\ln B)'(\ln F)' + p(\ln C)'(\ln F)' \right. \\ \left. - (d-1) \left[(\ln G)'' + (\ln G)'^2 + \frac{2}{r}(\ln G)' - (\ln G)'(\ln F)' - \frac{1}{r}(\ln F)'' \right] \right\} \\ = \frac{1}{2} \frac{1}{F^2} \phi'^2 - \frac{d-2}{2(D-2)} e^{-a\phi} \frac{Q^2}{(Gr)^{2(d-1)}}, \end{aligned} \quad (3.32)$$

$$\begin{aligned} \frac{1}{F^2} \left\{ - \left[(\ln G)' + \frac{1}{r} \right] \left[(\ln B)' + p(\ln C)' - (\ln F)' + (d-1)(\ln G)' + \frac{d-1}{r} \right] \right. \\ \left. - (\ln G)'' + \frac{1}{r^2} + (d-2) \frac{F^2}{r^2 G^2} \right\} = \frac{p+1}{2(D-2)} e^{-a\phi} \frac{Q^2}{(Gr)^{2(d-1)}}, \end{aligned} \quad (3.33)$$

$$\begin{aligned} \frac{1}{F^2} \left\{ \phi'' + \phi' \left[(\ln B)' + p(\ln C)' - (\ln F)' + (d-1)(\ln G)' + \frac{d-1}{r} \right] \right\} \\ = -\frac{a}{2} e^{-a\phi} \frac{Q^2}{(Gr)^{2(d-1)}}. \end{aligned} \quad (3.34)$$

We have used the equations (C.39) of Appendix C, where the computation of the Ricci tensor (i.e. the l.h.s. of the Einstein equations) is performed. We are here in the particular case where all $C_i = C$.

3.2.2 The extremal dilatonic p -brane

We now set up for finding our first solution. It will become clear very shortly that the further restrictions that we will implement on the fields will actually confine the solution to be in a particular class, i.e. to be extremal. This case is however interesting because first of all the equations are easily solved, and also because these extreme p -branes turn out to be physically most interesting, being identified in some cases to fundamental objects.

The first simplification to the equations of motion (3.30)–(3.34) comes actually by choosing a definite gauge for the radial coordinate r . We fix what we call the ‘isotropic gauge’, i.e. we require $F = G$. With this choice, the metric can be written with the function $G^2(r)$ multiplying a flat metric for the d dimensional transverse space. Let us stress that this is not a restriction but a choice of coordinates which allows the equations of motion to take a simpler form for this working case.

The following restrictions, which are now physically sensible, rely on the fact that we focus on extremal solutions. Extremality roughly corresponds to the statement that the mass of the object is equal (or related) to its charge. This means that the object is fully characterized (in the absence of angular momenta) by only one parameter, which we take to be its charge. Moreover, in the context of a supersymmetric theory, extremality is most of the time² related to the fact that the solutions preserves a portion of supersymmetry.

We call this kind of configuration a Bogomol’nyi–Prasad–Sommerfield (BPS) state [30]. It will become clear that this class of BPS states have a complete balance of forces between each other, i.e. two such states can stay statically one next to the other. This in turn implies that all the forces between them either vanish either are proportional and eventually cancel. Since the dilaton is also responsible of one of these forces, the above no-force condition implies that the dilaton cannot introduce a new parameter. The conclusion of this discussion is that as far as an extremal solution is concerned, all the fields should depend on only one parameter. Note that the asymptotical values at infinity are not parameters of this sort because they are fixed from the beginning. Here we will always take e^ϕ and the metric components to be 1 at infinity, and the n -form field strength to vanish there.

The main restriction, or ansatz, that we take for the moment is partly motivated by the structure of the equations of motion (3.30) and (3.31): since their r.h.s. are identical, they are fully compatible with taking $B = C$. This could also be motivated requiring $SO(1, p)$ Lorentz invariance along the whole world-volume of the p -brane.

That extremality implies Lorentz invariance on the world-volume can be heuristically understood as follows: a configuration for which the mass equals the charge, and thus saturates a lower bound, is seen from the point of view of the world-volume of the p -brane as a configuration carrying no energy at all. It seems thus logic to describe this configuration as flat space from the $(p + 1)$ dimensional point of view. Any mass excitation, or in D dimensional language any departure from extremality, breaks the

²There are some extremal solution which are not supersymmetric. However this happens either in theories which do not have maximal supersymmetry, either for more complicated solutions, like the one discussed at the end of Section 3.3. Here we suppose that there are no such complications.

Lorentz invariance in $p + 1$ dimensions and thus should lead to $B \neq C$.

We are left with a set of 4 equations, named respectively (t) , (r) , (α) and $\square\phi$:

$$\begin{aligned} -(\ln B)'' - \frac{d-1}{r}(\ln B)' - (\ln B)' [(p+1)(\ln B)' + (d-2)(\ln G)'] \\ = -\frac{d-2}{2(D-2)}e^{-a\phi}G^{-2(d-2)}\frac{Q^2}{r^{2(d-1)}}, \end{aligned} \quad (3.35)$$

$$\begin{aligned} -(p+1)(\ln B)'' - (d-1)(\ln G)'' - (p+1)(\ln B)'^2 + (p+1)(\ln B)'(\ln G)' \\ - \frac{d-1}{r}(\ln G)' - \frac{1}{2}\phi'^2 = -\frac{d-2}{2(D-2)}e^{-a\phi}G^{-2(d-2)}\frac{Q^2}{r^{2(d-1)}}, \end{aligned} \quad (3.36)$$

$$\begin{aligned} -\left[(\ln G)' + \frac{1}{r}\right] [(p+1)(\ln B)' + (d-2)(\ln G)'] - (\ln G)'' - \frac{d-1}{r}(\ln G)' \\ = \frac{p+1}{2(D-2)}e^{-a\phi}G^{-2(d-2)}\frac{Q^2}{r^{2(d-1)}}, \end{aligned} \quad (3.37)$$

$$\phi'' + \frac{d-1}{r}\phi' + \phi' [(p+1)(\ln B)' + (d-2)(\ln G)'] = -\frac{a}{2}e^{-a\phi}G^{-2(d-2)}\frac{Q^2}{r^{2(d-1)}}. \quad (3.38)$$

As one can be convinced by a first glance at the equations above, and as already noted in Appendix C, the quantity $\varphi = (p+1)\ln B + (d-2)\ln G$ plays a crucial rôle. Noting the structure of the r.h.s. of the equations (3.35) and (3.37), one can actually obtain an equation for φ . Taking the combination of the equations $(p+1)(t) + (d-2)(\alpha)$ one obtains:

$$-\varphi'' - \varphi'^2 - \frac{2d-3}{r}\varphi' = 0. \quad (3.39)$$

The simplest way to satisfy this equation is to take $\varphi = 0$ or, restoring its definition:

$$(p+1)\ln B + (d-2)\ln G = 0, \quad B^{p+1}G^{d-2} = 1. \quad (3.40)$$

This can be justified as follows. The asymptotic value of φ is fixed to zero by the asymptotic values of B and G . The second integration constant coming from the equation (3.39) must not be arbitrary since an extreme solution should depend only on one parameter, and by the equation (3.29) we have already introduced the parameter Q . We thus fix this second constant to be zero.

We have now reduced the problem to the three following equations for G and ϕ :

$$(\ln G)'' + \frac{d-1}{r}(\ln G)' = -\frac{p+1}{2(D-2)}e^{-a\phi}G^{-2(d-2)}\frac{Q^2}{r^{2(d-1)}}, \quad (3.41)$$

$$\phi'' + \frac{d-1}{r}\phi' = -\frac{a}{2}e^{-a\phi}G^{-2(d-2)}\frac{Q^2}{r^{2(d-1)}}, \quad (3.42)$$

$$\frac{(D-2)(d-2)}{p+1}(\ln G)'^2 + \frac{1}{2}\phi'^2 = \frac{1}{2}e^{-a\phi}G^{-2(d-2)}\frac{Q^2}{r^{2(d-1)}}. \quad (3.43)$$

The last equation (3.43) is obtained from (3.36) taking into account (3.40) and (3.41).

Again, we fix the following quantity to vanish in order not to introduce a new parameter in the solution:

$$a \ln G - \frac{p+1}{D-2} \phi = 0.$$

The two equations (3.41) and (3.42) become thus redundant, and the equation (3.43) becomes a first order differential equation for the remaining function, which we take to be ϕ :

$$\frac{\Delta}{a^2(D-2)} \phi'^2 = \frac{1}{2} e^{-\frac{2\Delta}{a(D-2)}\phi} \frac{Q^2}{r^{2(d-1)}}, \quad (3.44)$$

where:

$$\Delta = (p+1)(d-2) + \frac{1}{2} a^2 (D-2). \quad (3.45)$$

Taking the square root of the equation (3.44), we obtain:

$$\left(e^{\frac{\Delta}{a(D-2)}\phi} \right)' = \pm \sqrt{\frac{\Delta}{2(D-2)}} \frac{|Q|}{r^{d-1}}. \quad (3.46)$$

This is straightforwardly solved to give:

$$e^{\frac{\Delta}{a(D-2)}\phi} = 1 + \frac{1}{d-2} \sqrt{\frac{\Delta}{2(D-2)}} \frac{|Q|}{r^{d-2}}. \quad (3.47)$$

We have chosen one of the two signs appearing in (3.46) requiring that the solution is not singular for $r > 0$. The asymptotic value of the dilaton is fixed to zero according to the discussion in Section 3.1. Note that in going from (3.46) to (3.47), we have implicitly assumed that $d > 2$, which is the condition required for having a constant dilaton (and an asymptotically flat geometry) at infinity.

One can show that the equation (3.42), when G is replaced for ϕ , is a direct consequence of (3.46). We have thus solved completely the problem of finding a non-trivial solution depending on only one parameter, Q .

We can characterize completely the solution in terms of the function (which is harmonic in d -space):

$$H = 1 + \frac{1}{d-2} \sqrt{\frac{\Delta}{2(D-2)}} \frac{|Q|}{r^{d-2}} \equiv 1 + \frac{h^{d-2}}{r^{d-2}}. \quad (3.48)$$

The metric components, dilaton and $(p+2)$ -form are thus:

$$B = H^{-\frac{d-2}{\Delta}}, \quad G = H^{\frac{p+1}{\Delta}}, \quad e^\phi = H^{\frac{D-2}{\Delta}}, \quad F_{ty_1 \dots y_p r} = (\pm) \sqrt{\frac{2(D-2)}{\Delta}} (H^{-1})', \quad (3.49)$$

where $(\pm) = Q/|Q|$.

The metric of a spherically symmetric, extremal p -brane is thus:

$$ds^2 = H^{-2\frac{d-2}{\Delta}} (-dt^2 + dy_1^2 + \dots + dy_p^2) + H^{2\frac{p+1}{\Delta}} (dr^2 + r^2 d\Omega_{d-1}^2). \quad (3.50)$$

There are now some interesting remarks to make about the solution we have just found.

The first one arises when we note that the function H which characterizes the solution diverges at $r = 0$. For the region of space-time determined by $r = 0$ to coincide actually with the origin of the spherical coordinates, and thus with the location of the p -brane, a necessary condition is that the radius of the $(d - 1)$ -sphere shrinks to zero size there. Now the radius of the S^{d-1} in the geometry (3.50) is given by Gr . If we consider qualitatively that near $r = 0$ the function H behaves like $H \sim r^{-(d-2)}$, then we have that $Gr \sim r^{-\frac{(p+1)(d-2)}{\Delta}+1} = r^{a^2 \frac{D-2}{2\Delta}}$. This means that the radius is vanishing at $r = 0$ whenever $a \neq 0$, i.e. for any truly dilatonic p -brane. For $a = 0$, the fact that $Gr \rightarrow cst$ is a hint that something different is happening at $r = 0$.

However the radius of the S^{d-1} is not all, one would like to know whether the locus $r = 0$ is a singularity of the geometry or not. This actually can be easily checked using the components of the Riemann tensor in the orthonormal frame, which are given in Appendix C, eqs. (C.20)–(C.24). One can consider the following component:

$$\mathcal{R}^{\hat{t}\hat{i}}_{\hat{t}\hat{i}} = -\frac{1}{G^2}(\ln B)^{\prime 2} = -\frac{(d-2)^2 Q^2}{2\Delta(D-2)} \left[r^{d-2} + \frac{1}{d-2} \sqrt{\frac{\Delta}{2(D-2)}} |Q| \right]^{-2\left(1+\frac{p+1}{\Delta}\right)} r^{-a^2 \frac{D-2}{\Delta}}.$$

It clearly diverges at $r = 0$ when $a \neq 0$. This in turn implies the divergence of the square of the Riemann tensor, i.e. the scalar $K = \mathcal{R}_{\mu\nu\rho\sigma} \mathcal{R}^{\mu\nu\rho\sigma}$.

We can thus conclude that for the dilatonic extreme p -branes, the curvature diverges at the origin, i.e. at the location of the brane itself. This could actually have been guessed from the fact that, for instance, the dilaton field is singular at $r = 0$, and provides a singular source term to the Einstein equations.

Before drawing conclusions on the physical consequences of this singular behaviour, we should consider the non-dilatonic case. Here the locus $r = 0$ yields a finite curvature, and the radius of the sphere does not vanish. It can be shown in general [85], and thus for some particular cases of interest in string/M-theory [82], that the geometry in this case is regular at $r = 0$, which becomes thus an horizon much similar to the one of the extreme Reissner-Nordström black hole.

One can rewrite the $a = 0$ metric in a Schwarzschild-like radial coordinate defined by $R^{d-2} = r^{d-2} + h^{d-2}$:

$$ds^2 = \left(1 - \frac{h^{d-2}}{R^{d-2}}\right)^{\frac{2}{p+1}} \left(-dt^2 + dy_1^2 + \dots + dy_p^2\right) + \left(1 - \frac{h^{d-2}}{R^{d-2}}\right)^{-2} dR^2 + R^2 d\Omega_{d-1}^2. \quad (3.51)$$

This metric is in the $G = 1$ gauge, and now the horizon is at $R = h$. Note that for $p = 0$ we find the $d + 1$ dimensional Reissner-Nordström black hole. We will not discuss here the global structure of the solution above, and we refer to [85, 82].

We will see that when considering black p -branes, i.e. non-extremal ones, also the dilatonic ones possess a regular external horizon. This external horizon coincides with the $r = 0$ singularity in the extremal limit. It is thus tempting to call in what follows the locus $r = 0$ the horizon even for the extremal dilatonic p -branes.

One should not discard these solutions because of their singular behaviour; rather, they should be considered as reproducing the long-distance, low-energy fields of a quantum object. One could think of them in the same way as of the classical electric field produced by, say, an electron, which is singular at the (classical) location of the electron. Note that for the p -branes of 10 dimensional supergravities, one can alternatively consider their 11 dimensional origin [85] and find that they all descend from regular ‘solitonic’ solutions. The complementarity of these two points of view must be ascribed to duality.

We postpone until the derivation of the formula giving the ADM mass the discussion on the relation between the mass and the charge of the above solution (3.49), and we will prove in Section 3.3 that, in the $D = 11$ case, the extreme solutions of the class described above preserve half of the supersymmetries.

3.2.3 The general non-extremal black p -brane

In this subsection we generalize the extremal p -brane solutions to the non-extremal ones, depending now on two parameters. The reason why we took this path and not the opposite one, consisting in deriving the general solution and then specializing to the extremal one, is that the derivation in the general case is more involved. Moreover, our approach to the present problem will be strongly guided by the results of the previous subsection.

We start from the general spherically symmetric metric (3.25) and from the usual electric ansatz for the $(p + 2)$ -form (3.29). The equations are thus simply (3.30)–(3.34).

The form of the above mentioned equations suggests that we introduce a new function:

$$BC^p F^{-1} G^{d-1} = f. \quad (3.52)$$

Note that in the extremal case of the previous subsection, we have that $f = B^{p+1} G^{d-2} = 1$ by virtue of (3.40).

If we also define the following shorthand for the r.h.s.:

$$\mathcal{S}^2 = \frac{1}{2(D-2)} e^{-a\phi} F^2 G^{-2(d-1)} \frac{Q^2}{r^{2(d-1)}},$$

the equations can be rewritten as:

$$(\ln B)'' + \frac{d-1}{r} (\ln B)' + (\ln B)' (\ln f)' = (d-2) \mathcal{S}^2, \quad (3.53)$$

$$(\ln C)'' + \frac{d-1}{r} (\ln C)' + (\ln C)' (\ln f)' = (d-2) \mathcal{S}^2, \quad (3.54)$$

$$\begin{aligned} (\ln f)'' + (\ln F)'' - (\ln F)' (\ln f)' - (\ln F)'^2 + (\ln B)'^2 + p (\ln C)'^2 + (d-1) (\ln G)'^2 \\ + 2 \frac{d-1}{r} (\ln G)' - \frac{d-1}{r} (\ln F)' + \frac{1}{2} \phi'^2 = (d-2) \mathcal{S}^2, \end{aligned} \quad (3.55)$$

$$(\ln G)'' + \frac{d-1}{r} (\ln G)' + (\ln G)' (\ln f)' + \frac{1}{r} (\ln f)' + \frac{d-2}{r^2} \left(1 - \frac{F^2}{G^2} \right)$$

$$= -(p+1)\mathcal{S}^2, \quad (3.56)$$

$$\phi'' + \frac{d-1}{r}\phi' + \phi'(\ln f)' = -a(D-2)\mathcal{S}^2. \quad (3.57)$$

We now redefine once more the variables, taking:

$$\ln B = \ln C + \ln \bar{B}, \quad \ln F = \ln G + \ln \bar{F}.$$

Again, the extremal case corresponds to $\bar{B} = \bar{F} = 1$. Reexpressing (3.53) in these variables, and taking into account (3.54), we obtain the homogeneous equation:

$$(\ln \bar{B})'' + \frac{d-1}{r}(\ln \bar{B})' + (\ln \bar{B})'(\ln f)' = 0. \quad (3.58)$$

Another equation with vanishing r.h.s. can be obtained taking the combination $(p+1)(3.54) + (d-2)(3.56)$:

$$\varphi'' + \frac{d-1}{r}\varphi' + \varphi'(\ln f)' + \frac{(d-2)}{r} \left[(\ln f)' + \frac{(d-2)}{r} (1 - \bar{F}^2) \right] = 0, \quad (3.59)$$

where as usual $\varphi = (p+1)\ln C + (d-2)\ln G$. Note that:

$$\ln f = \varphi + \ln \bar{B} - \ln \bar{F}. \quad (3.60)$$

We are now ready to reduce the variables by taking some combinations of the functions above to vanish. We wish to keep in the end two independent functions, because we expect the non-extremal solution to depend on two parameters, the charge and the mass (or the internal and the external horizons).

An interesting reduction is the following: guided by the previous, extremal, case, we take $\varphi = 0$. This relates C to G . We will in the end also relate ϕ to G . The second independent function is provided by f . Indeed, since $\ln \bar{B}$, $\ln \bar{F}$ and $\ln f$ all vanish in the extremal limit, it is consistent to take $\ln \bar{B}$ and $\ln \bar{F}$ proportional to $\ln f$:

$$\ln \bar{B} = c_B \ln f, \quad \ln \bar{F} = c_F \ln f. \quad (3.61)$$

The relation (3.60) constrains $c_B - c_F = 1$. With these simplifications, one can now solve for f , \bar{B} and \bar{F} . The equation (3.58) becomes a second order differential equation for f , and the equation (3.59), which becomes a first order one, will be used to fix c_F . Note that the solution will contain an arbitrary integration constant, not related to the charge appearing in \mathcal{S}^2 .

The equation (3.58) rewrites:

$$(\ln f)'' + (\ln f)'^2 + \frac{d-1}{r}(\ln f)' = 0,$$

which gives:

$$f = 1 - \frac{2\mu}{r^{d-2}}. \quad (3.62)$$

The second integration constant has been set to one by the asymptotic value of the metric. μ is thus the non-extremality parameter.

To determine c_F , the equation (3.59) gives for $\varphi = 0$:

$$(\ln f)' + \frac{d-2}{r} (1 - f^{2c_F}) = 0.$$

Using the fact that $f' = \frac{d-2}{r}(1-f)$, the equation above holds if and only if $c_F = -\frac{1}{2}$. We thus have the solution:

$$\bar{B} = f^{\frac{1}{2}}, \quad \bar{F} = f^{-\frac{1}{2}}. \quad (3.63)$$

The two equations (3.56) (where we take into account (3.59) for $\varphi = 0$) and (3.57) again suggest, as in the extremal case, that $\ln G$ and ϕ are proportional, actually through the same relation $a \ln G = \frac{p+1}{D-2}\phi$.

This leaves us with the equation for ϕ :

$$\phi'' + \frac{d-1}{r}\phi' + (\ln f)'\phi' = -\frac{a}{2f}e^{-\frac{2\Delta}{a(D-2)}\phi} \frac{Q^2}{r^{2(d-1)}}. \quad (3.64)$$

Δ has been defined in (3.45), and we have used all the above relations to rewrite \mathcal{S}^2 in terms of ϕ and f only.

Performing the same game of substitutions in the equation (3.55), and using (3.56) and the equations satisfied by f , we finally obtain:

$$\frac{\Delta}{a(D-2)}\phi'^2 - \phi'(\ln f)' = \frac{a}{2f}e^{-\frac{2\Delta}{a(D-2)}\phi} \frac{Q^2}{r^{2(d-1)}}. \quad (3.65)$$

The sum of the two above equations is homogeneous and dictates that the function $e^{\frac{\Delta}{a(D-2)}\phi}$ is harmonic. The solution is thus of the same form as in the extreme case:

$$e^{\frac{\Delta}{a(D-2)}\phi} = 1 + \frac{h^{d-2}}{r^{d-2}} \equiv H, \quad (3.66)$$

where h is an arbitrary (positive) parameter. However the parameter h has to be determined from (3.64) or (3.65), which must reduce to algebraic equations for h . Substituting in (3.64) or in (3.65) the expressions (3.66) and (3.62), we indeed obtain the same algebraic equation:

$$(h^{d-2})^2 + 2\mu h^{d-2} = \frac{\Delta}{2(D-2)(d-2)^2}Q^2. \quad (3.67)$$

The solution is thus:

$$h^{d-2} = \sqrt{\frac{\Delta}{2(D-2)(d-2)^2}Q^2 + \mu^2} - \mu, \quad (3.68)$$

which correctly reduces to the extremal value of h which can be extracted from the equation (3.48) when the non-extremality parameter μ vanishes.

The solution is thus fully characterized in terms of the two functions f and H , and thus in terms of the two parameters Q (or h) and μ :

$$B = f^{\frac{1}{2}} H^{-\frac{d-2}{\Delta}}, \quad C = H^{-\frac{d-2}{\Delta}}, \quad F = f^{-\frac{1}{2}} H^{\frac{p+1}{\Delta}}, \quad G = H^{\frac{p+1}{\Delta}},$$

$$e^\phi = H^a \frac{D-2}{\Delta}, \quad F_{ty_1 \dots y_p r} = (\pm) \sqrt{\frac{2(D-2)}{\Delta}} \sqrt{1 + \frac{2\mu}{h^{d-2}}} (H^{-1})'. \quad (3.69)$$

It is straightforward to check that for $\mu = 0$, $f = 1$, the above solution becomes the extremal one (3.49). The metric writes [84]:

$$ds^2 = H^{-2\frac{d-2}{\Delta}} \left(-f dt^2 + dy_1^2 + \dots + dy_p^2 \right) + H^{2\frac{p+1}{\Delta}} \left(f^{-1} dr^2 + r^2 d\Omega_{d-1}^2 \right). \quad (3.70)$$

Note that for vanishing charge, $Q = 0$, we have also $h = 0$ and thus $H = 1$. The metric above thus reduces to the metric of a Schwarzschild black hole in $d + 1$ dimensions, multiplied by a flat R^p space.

The space-time structure of the above geometry can also be analyzed (see for instance [82]), and one finds that the locus $r^{d-2} = 2\mu$ is a regular horizon, while $r = 0$ is a singularity whenever $a \neq 0$. As already stated in the previous subsection, the horizon coincides with the singularity in the extremal limit for a dilatonic brane.

We can perform a change of coordinates defining $R^{d-2} = r^{d-2} + h^{d-2}$. If we define:

$$H_\pm = 1 - \frac{h_\pm^{d-2}}{R^{d-2}}, \quad h_+^{d-2} = h^{d-2} + 2\mu, \quad h_- = h,$$

we have that $H = H_-^{-1}$, $f = H_+ H_-^{-1}$, $r = R H_-^{\frac{1}{d-2}}$ and $\frac{dr}{dR} = H_-^{-\frac{d-3}{d-2}}$. This readily gives the following metric:

$$ds^2 = -H_+ H_-^{2\frac{d-2}{\Delta}-1} dt^2 + H_-^{2\frac{d-2}{\Delta}} \left(dy_1^2 + \dots + dy_p^2 \right)$$

$$+ H_+^{-1} H_-^{-2\frac{p+1}{\Delta}-\frac{d-4}{d-2}} dR^2 + H_-^{a^2 \frac{D-2}{\Delta(d-2)}} R^2 d\Omega_{d-1}^2. \quad (3.71)$$

This metric reduces to the one found in $D = 10$ and in the string frame by Horowitz and Strominger [45]. Note that in this picture μ measures the difference between the external and the internal horizons, namely $2\mu = h_+^{d-2} - h_-^{d-2}$.

For the non-dilatonic branes $a = 0$, we can reproduce the metric discussed in [85]:

$$ds^2 = -H_+ H_-^{-\frac{p-1}{p+1}} dt^2 + H_-^{\frac{2}{p+1}} \left(dy_1^2 + \dots + dy_p^2 \right) + H_+^{-1} H_-^{-1} dR^2 + R^2 d\Omega_{d-1}^2. \quad (3.72)$$

This metric is in the $G = 1$ gauge. We could have actually found this metric solving the equations in this specific gauge. This is indeed possible, however this approach only works in the non-dilatonic case, and this is why we did not consider it here.

3.2.4 Semi-classical thermodynamics and the mass of the black p -branes

The aim of this subsection is to provide us with a formula for the mass of a p -brane or, more generally, of a configuration of p -branes giving rise to a metric like (3.26).

Such a formula for the mass could be simply obtained using the ADM (Arnowitt–Deser–Misner) formalism (see e.g. [86]), and this has already been derived for p -branes in [87]. Here we prefer to rederive the ADM formula in our particular case, and we base our derivation on considerations pertaining to semi-classical thermodynamics of black objects. Almost as a by-product, we will obtain also useful formulas for the Hawking temperature and for the entropy of the same configurations.

Once the formula for the mass will have been found, we will show that the extreme p -branes (3.49) saturate a lower bound for their mass, which is indeed proportional to their charge.

Let us now introduce some aspects of the semi-classical thermodynamics of gravitating systems that we will need to derive the ADM mass formula. The general idea is the following: one wants to compute the partition function of a system including gravitation. If one is only interested in the contribution to the partition function of the (classical) background, then only the leading term of the saddle point approximation is relevant, which means that all fluctuation determinants are neglected. We call this approximation the semi-classical one.

As shown by Gibbons and Hawking [88], the semi-classical contribution to the partition function is non-trivial when the background gravitational field has a horizon, i.e. when it defines a black object. However to correctly compute this contribution, one has to pay much attention to the boundary terms which have to be added to the action (3.9) in order to implement the boundary conditions on the fields. The analysis of which boundary terms are needed in general, and how to regularize them, can be carried out (see for instance [89, 90]) but it involves the introduction of quite a lot of formalism.

Here we will use only some useful aspects of the formalism, and we will short-cut the rest. Inspired by [91], the short-cut exploits some simplifications introduced by the Hamiltonian formalism.

In order to compute a partition function in field theory, one has to define a path integral over field configurations which are periodic in the imaginary time $\tau = it$, with period β . In theories including gravitation, this step means that we are really considering a Euclidean section of a complexified metric [88]. Here comes the first surprise: when the background geometry is the one of a non-extreme black object, the Euclidean section stops at the locus where the Lorentzian section defines a horizon, and it is smooth provided the period β has a definite value. Moreover, the value of β , which is the inverse temperature, is in perfect agreement with the Hawking temperature T_H . The conclusion is that the Euclidean section of a black object is a smooth manifold with only one boundary at infinity, and that this procedure singles out a definite temperature T_H .

If the integrand in the path integral is e^{-I_E} , the Euclidean action I_E is simply minus the Lorentzian one (3.9), where however the metric is replaced by one of Euclidean

signature. The gravitational part thus reads:

$$I_E = -\frac{1}{16\pi G_D} \int d^D x \sqrt{g} R + \text{boundary terms at } \infty. \quad (3.73)$$

We will no longer include the matter terms as in (3.9) since they are not relevant to the following discussion, centered on the derivation of the ADM mass formula. These terms are however important for a general treatment of the thermodynamics of black objects, and in particular for the distinction between the canonical and the grand canonical ensembles, distinction which can be subtle when dealing with electrically and/or magnetically charged objects [90].

The action I_E can be used to define the partition function in the canonical ensemble in the following way [88]:

$$-\ln Z(\beta) = \beta F = \beta M - S \approx I_E, \quad (3.74)$$

where the symbol \approx denotes that the action is evaluated on the equations of motion (and also that the path integral is reduced to its saddle point value). M is the ADM mass, S is the entropy and β is the inverse temperature, $\beta = T^{-1}$. Using the first law of (black hole) thermodynamics, $\delta M = T\delta S$, it is straightforward to show that we must have:

$$\delta I_E \approx M\delta\beta, \quad (3.75)$$

which is consistent with the fact that the canonical partition function is a function of the temperature, and that the equations of motion are derived from I_E taking the variation at fixed β .

Consider now the Hamiltonian action. For a static space-time like (3.26) and in the Euclidean section it can be written as:

$$I_h = \int d^D x N \mathcal{H}, \quad (3.76)$$

where $N \equiv B$ is the so-called lapse function, which does not appear in the Hamiltonian density \mathcal{H} and thus acts as a Lagrange multiplier. Accordingly, the action I_h vanishes on-shell.

Now, following [92], this action cannot give the true Hamiltonian precisely because it always vanishes on-shell. Moreover, its variation yields a non-vanishing boundary term. One thus has to add a boundary term to I_h which all at once cancels, when varied, the non-trivial variation of I_h , and when evaluated yields the energy of the configuration. The variation at infinity of I_h is thus [92]:

$$\delta I_h|_{\infty} \approx -\beta\delta M. \quad (3.77)$$

We have thus now the procedure to compute the ADM mass. First we compute I_h simply taking (3.73) and expurgating from it all terms containing derivatives of B . This will be done pushing away these terms into boundary integrals that will all be discarded. Then we compute the variation of I_h , knowing that we will find, besides the equations of motion, the boundary term at infinity (3.77). The boundary term of the variation

at the horizon on the other hand will give us an interesting relation from which we will extract the expressions for both T and S .

Let us thus start by computing the action I_E . We use the results (C.39) of the Appendix C. Then if we take the by now familiar function $\varphi = \ln B + \sum_i \ln C_i - \ln F + (d-1)\ln(Gr)$, and consider that $R = R_t^i + \sum_i R_r^i + R_r^{\hat{r}} + (d-1)R_{\hat{\alpha}}^{\hat{\alpha}}$, we obtain:

$$I_E = \frac{1}{16\pi G_D} \int d^D x \omega_{d-1} e^\varphi \left\{ 2\varphi'' + \varphi'^2 + 2(\ln F)'' + (\ln B)'^2 + \sum_i (\ln C_i)'^2 - (\ln F)'^2 \right. \\ \left. + (d-1)(\ln Gr)'^2 - (d-1)(d-2)\frac{F^2}{G^2 r^2} \right\}, \quad (3.78)$$

where $\omega_{d-1} = \sin^{d-2} \theta_1 \dots \sin \theta_{d-2}$. We now wish to single out all dependence on B . Accordingly, we introduce the new auxiliary function γ such that $\varphi = \ln B + \gamma$. The action I_E can thus be rewritten in two pieces:

$$I_E = I_B + I_h.$$

The first piece contains all the terms with derivatives of B , and is actually a total derivative:

$$I_B = \frac{1}{8\pi G_D} \int d^D x \omega_{d-1} B e^\gamma \left\{ (\ln B)'' + (\ln B)'^2 + (\ln B)'\gamma' \right\} \\ = \frac{1}{8\pi G_D} \int d^D x \omega_{d-1} (e^\gamma B)'$$

The second piece is thus the sought for Hamiltonian action:

$$I_h = \frac{1}{16\pi G_D} \int d^D x \omega_{d-1} B e^\gamma \left\{ 2\gamma'' + \gamma'^2 + 2(\ln F)'' + \sum_i (\ln C_i)'^2 - (\ln F)'^2 \right. \\ \left. + (d-1)(\ln Gr)'^2 - (d-1)(d-2)\frac{F^2}{G^2 r^2} \right\}. \quad (3.79)$$

Note that the action above should be completed with other bulk terms arising from the matter terms in the action (3.9), which would correctly yield $I_h \approx 0$. However these terms are not relevant for the present discussion, since we are not interested in recovering the equations of motion.

We now take the variation of the action (3.79) and we care only about the boundary terms which are produced integrating by parts. We obtain:

$$\delta I_h \approx \frac{1}{8\pi G_D} \int d^D x \omega_{d-1} \left\{ B e^\gamma \left[\delta\gamma' - (\ln B)'\delta\gamma + \delta(\ln F)' - (\ln B)'\delta(\ln F) \right. \right. \\ \left. \left. - \gamma'\delta(\ln F) + \sum_i (\ln C_i)'\delta(\ln C_i) - (\ln F)'\delta(\ln F) + (d-1)(\ln Gr)'\delta(\ln Gr) \right] \right\}'.$$

We can now perform the integral and we obtain two contributions, one from infinity and one from the horizon $r = r_h$. If we also substitute γ for its definition, we have:

$$\begin{aligned} \delta I_h \approx & \frac{\beta L^p \Omega_{d-1}}{8\pi G_D} B C_1 \dots C_p \frac{1}{F} (Gr)^{d-1} \left[\sum_i \frac{1}{C_i} \delta C'_i + (d-1) \frac{1}{G} \delta G' \right. \\ & + (d-1) \frac{1}{Gr} \delta G - \sum_i \frac{B'}{BC_i} \delta C_i - (d-1) \frac{B'}{BG} \delta G \\ & \left. - \sum_i \frac{C'_i}{C_i F} \delta F - (d-1) \frac{G'}{FG} \delta F - (d-1) \frac{1}{Fr} \delta F \right] \Bigg|_{r=r_h}^{r \rightarrow \infty}. \end{aligned} \quad (3.80)$$

Here β is the period of the Euclidean time, L^p is the volume of the ‘longitudinal’ space spanned by the y ’s and Ω_{d-1} is the volume of the $(d-1)$ -sphere S^{d-1} .

Let us focus first on the boundary term at infinity. We have to consider the behaviour of the functions appearing in (3.80), and to implement the requirement of asymptotic flatness (recall that now we have crucially $d \geq 3$). By inspection of the metrics (3.70) and (3.71), we see that all the black p -brane solutions are such that:

$$B, C_i, F, G \equiv \mathcal{F} = 1 + \mathcal{O}\left(\frac{1}{r^{d-2}}\right). \quad (3.81)$$

The leading term of the functions appearing in (3.80) is thus respectively the following:

$$\mathcal{F} = 1, \quad \delta \mathcal{F} = \mathcal{O}\left(\frac{1}{r^{d-2}}\right), \quad \mathcal{F}', \delta \mathcal{F}' = \mathcal{O}\left(\frac{1}{r^{d-1}}\right).$$

When taking into account this behaviour, the only terms in (3.80) which do not die off at infinity are the following:

$$\delta I_h|_{\infty} \approx \frac{\beta L^p \Omega_{d-1}}{8\pi G_D} r^{d-1} \left[\sum_i \delta C'_i + (d-1) \delta G' + (d-1) \frac{1}{r} \delta G - (d-1) \frac{1}{r} \delta F \right] \Bigg|_{r \rightarrow \infty}.$$

Now, recalling the equation (3.77) relating the variation of the action I_h at infinity to the variation of the mass, we derive the expression for the ADM mass:

$$M = \frac{L^p \Omega_{d-1}}{8\pi G_D} r^{d-1} \left[\frac{d-1}{r} (F - G) - \sum_i C'_i - (d-1) G' \right] \Bigg|_{r \rightarrow \infty}. \quad (3.82)$$

Note that we do not have to subtract any constant part from (3.82) since the expression vanishes for flat space, i.e. when all functions are equal to 1.

The formula (3.82) can also be shown to be invariant under reparametrizations of r which respect the behaviour at infinity (3.81), i.e. reparametrizations such that:

$$\bar{r} = r \left[1 + \mathcal{O}\left(\frac{1}{r^{d-2}}\right) \right].$$

Before evaluating the expression of the ADM mass for the black p -branes of the previous subsection, we take a quick look at the second term in (3.80), the variation at the horizon $\delta I_h|_{r_h}$.

There is however a subtlety when considering the variation of the action at the horizon. One in fact needs a definition of the location of the horizon which is independent from the parameters of the metric that one is varying. Without delving into a truly diffeomorphism-invariant formulation of the boundary term at the horizon, we simply use the possibility to operate a reparametrization of the radial coordinate to always bring back the location of the horizon to, say, $\xi = 0$. The new radial coordinate should also be such that the functions defining the metric have a controllable behaviour at the horizon.

Since the Euclidean topology of a black hole is $R^2 \times S^{d-1}$ (see for instance [91]), and thus that of a black p -brane is $R^2 \times T^p \times S^{d-1}$, where the Euclidean time τ and the coordinate ξ are respectively the angular variable and the radial coordinate on the R^2 factor, all the functions in the metric can be taken of order 1, $C_i, F, (Gr) = \mathcal{O}(1)$, except B which has to be $B = \mathcal{O}(\xi)$. We have implicitly used the fact that now $(Gr) \equiv R(\xi)$ denotes a function of ξ , r being the original, non-gauge fixed radial variable.

In order to avoid conical singularity at the origin of the R^2 plane, which coincides with the location of the horizon, the angular variable τ has to have a definite periodicity. If the metric near the horizon is:

$$ds^2 \simeq B_h'^2 \xi^2 d\tau^2 + F_h^2 d\xi^2 + \dots, \quad (3.83)$$

where the subscript h denotes the value of the function at $\xi = 0$, we can define $\bar{\xi} = F_h \xi$ such that:

$$ds^2 \simeq \left(\frac{B_h'}{F_h} \right)^2 \bar{\xi}^2 d\tau^2 + d\bar{\xi}^2 + \dots$$

We define the surface gravity at the horizon by:

$$\kappa = \left(\frac{B'}{F} \right) \Big|_{r=r_h}. \quad (3.84)$$

Then there is no conical singularity at the horizon if $\tau \sim \tau + \beta_H$, with $\beta_H = \frac{2\pi}{\kappa}$. The Hawking temperature is thus the inverse of β_H :

$$T_H = \frac{\kappa}{2\pi} = \frac{1}{2\pi} \left(\frac{B'}{F} \right) \Big|_{r=r_h}. \quad (3.85)$$

Note that although we have derived this expression in the particular ‘gauge’ (3.83), the above expression for T_H is invariant under reparametrizations of the radial coordinate, and it is thus a good expression for a general r .

We can now reconsider the horizon term of the variation (3.80). The only terms that survive are the ones containing B' , all the others vanish due to $B|_{r_h} = 0$:

$$\begin{aligned} \delta I_h|_{r_h} &\approx \frac{\beta L^p \Omega_{d-1}}{8\pi G_D} C_1 \dots C_p \frac{1}{F} (Gr)^{d-1} B' \left[\sum_i \frac{1}{C_i} \delta C_i + (d-1) \frac{1}{Gr} \delta(Gr) \right] \Big|_{r=r_h} \\ &\approx \frac{\beta L^p \Omega_{d-1}}{8\pi G_D} \frac{B'}{F} \delta \left(C_1 \dots C_p (Gr)^{d-1} \right) \Big|_{r=r_h}. \end{aligned} \quad (3.86)$$

As noted in [91], this term is of the form:

$$\delta I_h|_{r_h} \approx \frac{\beta}{\beta_H} \frac{\delta A_h}{4G_D}, \quad (3.87)$$

where A_h is the volume of the horizon (the $(D - 2)$ -dimensional manifold defined by $r = r_h$). Such a term in the variation would break covariance, since it singles out a definite location on the (Euclidean) manifold, i.e. the horizon. However, we still have to implement that the periodicity has to be $\beta = \beta_H$ in order to have a truly smooth manifold. That can be done simply adding a term to I_h such that:

$$I'_h = I_h - \frac{A_h}{4G_D}. \quad (3.88)$$

The variation at the horizon of the new action is thus:

$$\delta I'_h|_{r_h} \approx \left(\frac{\beta}{\beta_H} - 1 \right) \frac{\delta A_h}{4G_D},$$

which enforces $\beta = \beta_H$ if A_h , which is a function of the metric, is to vary freely. If we also add the boundary term at infinity βM to I'_h , we obtain once again the action I_E . Indeed, one can check that the variation of $I'_h + \beta M$ yields the equations of motion when β is kept fixed. Taking into account (3.74), (3.88) and the fact that $I_h \approx 0$, we can equate the on-shell values of I_E and of $I'_h + \beta M$, thus obtaining the familiar formula for the entropy:

$$S = \frac{A_h}{4G_D}. \quad (3.89)$$

More explicitly, we have the following expression for it:

$$S = \frac{L^p \Omega_{d-1}}{4G_D} C_1 \dots C_p (Gr)^{d-1} \Big|_{r=r_h}. \quad (3.90)$$

It is also an r -reparametrization invariant expression.

Let us now finally apply the formulas (3.82), (3.85) and (3.90) to the black p -brane solution (3.69).

For the ADM mass, we have:

$$M_p = \frac{L^p \Omega_{d-1}}{8\pi G_D} \left[(d-1)\mu + (d-2) \frac{D-2}{\Delta} h^{d-2} \right]. \quad (3.91)$$

Using the expression (3.68) for h^{d-2} , (3.91) can be rearranged as:

$$M_p = \frac{L^p \Omega_{d-1}}{8\pi G_D} \left\{ \frac{1}{\Delta} \left[\frac{1}{2} a^2 (D-2)(d-1) + p(d-2)^2 \right] \mu \right. \\ \left. + (d-2) \frac{D-2}{\Delta} \sqrt{\frac{\Delta}{2(D-2)(d-2)^2 Q^2 + \mu^2}} \right\}.$$

For fixed Q , the bound $\mu \geq 0$ yields for the ADM mass:

$$M_p \geq \frac{L^p \Omega_{d-1}}{8\pi G_D} \sqrt{\frac{D-2}{2\Delta}} |Q|, \quad (3.92)$$

which is naturally saturated by the $\mu = 0$, extreme p -brane solutions (3.49).

In order to compare the ADM mass with the physical charge, we compute the charge density in the canonical way:

$$\mathcal{Q} = \frac{1}{16\pi G_D} \int_{S^{d-1}} *F_{p+2}, \quad (3.93)$$

where $*F_{p+2} \equiv \tilde{F}_{d-1}$ as defined in (3.21), and the charge density has the right dimensions of energy per unit p -volume.

If we compute the total amount of charge carried by the solution (3.69), we find:

$$|q| = L^p |\mathcal{Q}| = \frac{L^p \Omega_{d-1}}{16\pi G_D} |Q|. \quad (3.94)$$

The final form of the bound on the ADM mass, or Bogomol'nyi bound, is thus:

$$M_p \geq \sqrt{\frac{2(D-2)}{\Delta}} |q|. \quad (3.95)$$

We now compute the Hawking temperature T_H for the black p -branes of the solution (3.69). The behaviour of T_H when the extremal limit $\mu \rightarrow 0$ is taken will deserve some comments.

In terms of the functions f and H , see (3.62) and (3.66), we have the following surface gravity (3.84):

$$\kappa = \frac{1}{2} f' H^{-\frac{D-2}{\Delta}} \Big|_{r=r_h}.$$

Note that the horizon is at $r_h = (2\mu)^{\frac{1}{d-2}}$, i.e. at $f = 0$. When we substitute for f and H , we find for the Hawking temperature:

$$\begin{aligned} T_H &= \frac{d-2}{4\pi} \frac{1}{(2\mu)^{\frac{1}{d-2}}} \left(1 + \frac{h^{d-2}}{2\mu} \right)^{-\frac{D-2}{\Delta}} \\ &= \frac{d-2}{4\pi} (2\mu)^{-\frac{1}{d-2} + \frac{D-2}{\Delta}} (2\mu + h^{d-2})^{-\frac{D-2}{\Delta}}. \end{aligned} \quad (3.96)$$

In the limit $\mu \rightarrow 0$, h is finite as seen from its definition (3.68), thus the second term in the above expression is finite. The first factor on the other hand gives a vanishing, finite or infinite Hawking temperature at extremality depending whether the combination $-\frac{1}{d-2} + \frac{D-2}{\Delta}$ is respectively positive, zero or negative.

Let us now see which branes of type II/M-theory give rise to which behaviour at extremality. In 10 or 11 dimensional maximal supergravities, we will see in the next subsection that Δ as given in (3.45) is always such that $\Delta = 2(D-2)$. Then the

dependence of the temperature T_H on μ is governed by the combination $-\frac{1}{d-2} + \frac{1}{2} = \frac{d-4}{2(d-2)}$. This gives the following behaviours (see also [93]): all the branes of type II/M-theory have a vanishing extremal Hawking temperature except the 10 dimensional 5-branes, which have a finite temperature in the extremal limit, and the D6-brane of type IIA theory which has an infinite extremal temperature. We are of course only considering branes which allow for an asymptotically flat geometry, $d \geq 3$.

The fact that the D6-brane is singled out, and that the 5-branes of type II theories are the limiting cases, is a pattern that will again show up when we will consider in Chapter 5 the decoupling of the physics on the world-volume of the brane from the bulk supergravity dynamics. It is tempting to think that these two features (the decoupling from the bulk and the vanishing of the Hawking temperature at extremality) are closely related. Indeed the presence of the Hawking temperature signals, from the brane world-volume point of view, that quanta are being exchanged with the surrounding space, i.e. the bulk. However a more detailed analysis is required, and we will get back to this problem when we will consider the world-volume effective actions of the branes.

The last quantity we wish to compute for the solutions (3.70) is the entropy (3.90). We have that:

$$S = \frac{L^p \Omega_{d-1}}{4G_D} H^{\frac{D-2}{\Delta}} r^{d-1} \Big|_{r=r_h} .$$

The expression is independent of f . Plugging in (3.66) and $r_h = (2\mu)^{\frac{1}{d-2}}$, we obtain:

$$\begin{aligned} S &= \frac{L^p \Omega_{d-1}}{4G_D} \left(1 + \frac{h^{d-2}}{2\mu} \right)^{\frac{D-2}{\Delta}} (2\mu)^{\frac{d-1}{d-2}} \\ &= \frac{L^p \Omega_{d-1}}{4G_D} (2\mu)^{\frac{d-1}{d-2} - \frac{D-2}{\Delta}} \left(2\mu + h^{d-2} \right)^{\frac{D-2}{\Delta}} . \end{aligned} \quad (3.97)$$

Since we have the relation $\frac{d-1}{d-2} - \frac{D-2}{\Delta} = \frac{a^2(D-2)(d-1)}{2\Delta(d-2)} + \frac{p(d-2)}{\Delta}$, the entropy always vanishes in the extremal limit, except for the non-dilatonic black hole, i.e. the 0-brane with $a = 0$. It is amusing to see that the Reissner-Nordström black hole falls into this last, peculiar, class.

3.2.5 The branes of M-theory

As a physical conclusion to this rather mathematical section, we list and discuss hereafter the extremal branes that are found in M-theory, or more precisely in 11 dimensional supergravity and in type II string theories. However, this is not merely a zoological review of some weird specimens. Indeed, as we will show in the next section, these extremal branes will turn out to be the building blocks of a lot of new, more complicated configurations, which carry more than one charge. Moreover, some of these new configurations will show new features like a non-vanishing entropy in the extremal limit. The focus on extremal branes is also motivated by the conjectured identification of some, and possibly all, of these solitons with (BPS saturated) fundamental objects.

Let us rewrite the solution for clarity (we everywhere replace $d = D - p - 1$):

$$ds^2 = H^{-2\frac{D-p-3}{\Delta}} \left(-dt^2 + \sum_i dy_i^2 \right) + H^{2\frac{p+1}{\Delta}} \delta_{ab} dx^a dx^b, \quad (3.98)$$

$$e^\phi = H^{a\frac{D-2}{\Delta}}, \quad F_{ty_1\dots y_{p-1}} = E', \quad E \equiv A_{ty_1\dots y_p} = \sqrt{\frac{2(D-2)}{\Delta}} (H^{-1} - 1), \quad (3.99)$$

with:

$$H = 1 + \frac{1}{D-p-3} \sqrt{\frac{\Delta}{2(D-2)}} \frac{Q}{r^{D-p-3}}, \quad \Delta = (p+1)(D-p-3) + \frac{1}{2} a^2 (D-2). \quad (3.100)$$

All the discussion in the previous subsections was based on the above electric solution. However we now exploit the symmetry of the equations of motion displayed in (3.21) to obtain a magnetic solution. This is straightforward: we simply take the above solution and we replace a by $-a$ and F by \tilde{F} . Using (3.21) we can then compute the magnetic $(D-p-2)$ -form field strength:

$$F^{\theta_1\dots\theta_{D-p-2}} = \frac{1}{\sqrt{-g}} e^{-(a)\phi} \tilde{F}_{ty_1\dots y_{p-1}} = \frac{1}{\sqrt{-g}} e^{a\phi} E',$$

which thus gives instead of the line (3.99):

$$e^\phi = H^{-a\frac{D-2}{\Delta}}, \quad F_{\theta_1\dots\theta_{D-p-2}} = Q\omega_{D-p-2} \quad (\omega_{D-p-2} = \sin^{D-p-3} \theta_1 \dots \sin \theta_{D-p-3}). \quad (3.101)$$

The first quantity one has to compute to write down a specific solution is Δ . Now, since for an electric-magnetic pair of branes deriving from the same field strength we have $p_{el} + p_{mag} = D - 4$, we see that Δ is the same for both and thus characterizes the n -form. We will now see that for all the cases of interest below, Δ will always satisfy $\Delta = 2(D-2)$.

Let us start with 11 dimensional supergravity. Here we have only a 4-form field strength and since there is no dilaton, we put $a = 0$. The 4-form gives rise to 2- and 5-branes, called respectively M2 and M5. As anticipated, we have $\Delta = 3 \cdot 6 = 18$.

The two solutions are thus:

$$ds_{M2}^2 = H^{-\frac{2}{3}} \left(-dt^2 + dy_1^2 + dy_2^2 \right) + H^{\frac{1}{3}} \left(dx_1^2 + \dots + dx_8^2 \right), \quad F_{ty_1y_2r} = (H^{-1})', \quad (3.102)$$

and:

$$ds_{M5}^2 = H^{-\frac{1}{3}} \left(-dt^2 + dy_1^2 + \dots + dy_5^2 \right) + H^{\frac{2}{3}} \left(dx_1^2 + \dots + dx_5^2 \right), \quad F_{\theta_1\dots\theta_4} = Q\omega_4. \quad (3.103)$$

Since there is no dilaton, the Einstein metric coincides with the original metric as in (3.3). These solutions were found respectively in [94] and [95].

Going to type II theories in 10 dimensions, we can separate the fields in two classes, those coming from the NSNS sector and those coming from the RR one.

Beginning with the NSNS 3-form field strength, we have $a = -1$ and it yields an electric 1-brane which we call F1 and a magnetic 5-brane referred to as NS5. The F1 is identified to the fundamental string since it couples to the same 2-form potential which couples to the world-sheet of the type II string. The NS5 has been called alternatively the solitonic 5-brane, and also the neutral 5-brane (when it is embedded in heterotic string theory).

The value of Δ is straightforwardly $\Delta = 2 \cdot 6 + \frac{1}{2} \cdot 8 = 16$. We then have the following metric and dilaton for the F1 and NS5:

$$ds_{F1}^2 = H^{-\frac{3}{4}} \left(-dt^2 + dy_1^2 \right) + H^{\frac{1}{4}} \left(dx_1^2 + \dots + dx_8^2 \right), \quad e^\phi = H^{-\frac{1}{2}}, \quad (3.104)$$

$$ds_{NS5}^2 = H^{-\frac{1}{4}} \left(-dt^2 + dy_1^2 + \dots + dy_5^2 \right) + H^{\frac{3}{4}} \left(dx_1^2 + \dots + dx_4^2 \right), \quad e^\phi = H^{\frac{1}{2}}. \quad (3.105)$$

The 3-form can be easily deduced from (3.99) and (3.101).

The two metrics above are again written in the Einstein frame. However in string theory, we can define another natural metric, which is the string metric and is obtained by a rescaling of the Einstein one. We rewrite the equation (3.5):

$$\hat{g}_{\mu\nu} = e^{\frac{\phi}{2}} g_{\mu\nu},$$

the metric in the string frame carrying (temporarily) a hat.

In the string frame, the metrics above become:

$$d\hat{s}_{F1}^2 = H^{-1} \left(-dt^2 + dy_1^2 \right) + dx_1^2 + \dots + dx_8^2, \quad (3.106)$$

$$d\hat{s}_{NS5}^2 = -dt^2 + dy_1^2 + \dots + dy_5^2 + H \left(dx_1^2 + \dots + dx_4^2 \right). \quad (3.107)$$

These metrics are solutions of the actions (3.4) and (3.7). The above solutions were discussed respectively in e.g. [96] and [57].

The RR n -forms all have a coupling to the dilaton such that $a_n = \frac{5-n}{2}$. This gives for an electric $(p+2)$ -form $a_{p+2} = \frac{3-p}{2}$ and for a magnetic $(8-p)$ -form $a_{8-p} = -\frac{3-p}{2}$. We can thus write in a unified way:

$$\varepsilon a = \frac{3-p}{2},$$

where ε is $+(-)$ if the p -brane is electric (magnetic). Again it can be shown that $\Delta = (p+1)(7-p) + (3-p)^2 = 16$.

The metric of the p -branes charged under the RR forms, which we call D_p in view of their relation with the D-branes, is:

$$ds_{D_p}^2 = H^{-\frac{7-p}{8}} \left(-dt^2 + dy_1^2 + \dots + dy_p^2 \right) + H^{\frac{p+1}{8}} \left(dx_1^2 + \dots + dx_{9-p}^2 \right), \quad e^\phi = H^{\frac{3-p}{4}}. \quad (3.108)$$

In the string frame, it gives [45]:

$$d\hat{s}_{D_p}^2 = H^{-\frac{1}{2}} \left(-dt^2 + dy_1^2 + \dots + dy_p^2 \right) + H^{\frac{1}{2}} \left(dx_1^2 + \dots + dx_{9-p}^2 \right). \quad (3.109)$$

For all the solutions discussed above, the mass saturates the bound (3.95) and moreover the mass is strictly equal to the total charge:

$$M = |q|, \quad q = \frac{L^p \Omega_{D-p-2}}{16\pi G_D} Q.$$

With the branes described above we have exhausted all the objects carrying a charge in 11 or 10 dimensions. When one considers a toroidal compactification of M-theory, all these charges are related by U-duality. However, in order to have a complete U-duality multiplet of charges, one has to include also the Kaluza-Klein (KK) charges generated by the dimensional reduction procedure.

Properly speaking, the KK charged objects are branes living in the compactified space-time and carrying a charge (electric or magnetic) with respect to the 2-form field strength generated by dimensional reduction, see Appendix A. If after compactification on one direction the space-time dimension is D , we then have an electric KK 0-brane and a magnetic KK $(D-4)$ -brane. These two objects correspond, in the $D+1$ dimensional uncompactified space-time to configurations where the only non-trivial field is the metric, and which are identified respectively to a KK wave and a KK monopole. Note that the metric in the uncompactified space necessarily has non-trivial off-diagonal terms, and thus is not of the class (3.25) considered in this section.

Nevertheless, we derive hereafter the metric of the KK wave and KK monopole simply by KK ‘oxidation’, i.e. by reversing the KK procedure to derive a $D+1$ dimensional metric from a D dimensional field configuration. Since the only relevant field in $D+1$ dimensions is the metric, we can work for general D .

Let us start with the KK wave. In D dimensions, the KK 2-form has a coupling to the (KK) dilaton $a = \sqrt{\frac{2(D-1)}{D-2}}$, see Appendix A and more precisely the discussion leading to the equation (A.15). For the electric KK 0-brane we thus have $\Delta = (D-3) + \frac{1}{2} \frac{2(D-1)}{D-2} (D-2) = 2(D-2)$.

The solution in D dimensions is thus:

$$ds_D^2 = -H^{-\frac{D-3}{D-2}} dt^2 + H^{\frac{1}{D-2}} (dx_1^2 + \dots + dx_{D-1}^2), \quad (3.110)$$

$$e^\phi = H^{\frac{a}{2}}, \quad A_t = H^{-1} - 1, \quad H = 1 + \frac{1}{D-3} \frac{Q}{r^{D-3}}. \quad (3.111)$$

The first step before going to $D+1$ dimensions, is to rescale the metric (3.110) to put it in the $D+1$ dimensional Einstein frame. Using the formula (A.11), we have that:

$$\hat{g}_{\mu\nu} = H^{-\frac{1}{D-2}} g_{\mu\nu},$$

where we have already used that $h_{yy} \equiv e^{2\sigma} = e^{\frac{2}{a}\phi} = H$. y is the compact direction.

Collecting the above results, the $D+1$ dimensional metric can be written as:

$$ds_{D+1}^2 = -H^{-1} dt^2 + dx_1^2 + \dots + dx_{D-1}^2 + H \left[dy + (H^{-1} - 1) dt \right]^2.$$

Note that all dependence on D has disappeared from the exponents. If we define the new harmonic function K such that $H = 1 + K$, we arrive at the final form for the metric of a KK wave:

$$ds_W^2 = -dt^2 + dy^2 - K(dt - dy)^2 + dx_1^2 + \dots + dx_{D-1}^2. \quad (3.112)$$

Consider now our last object, the KK monopole. Since in the reduced D dimensional picture it has to be magnetically charged under a 2-form field strength, it is a $(D - 4)$ -brane. We can thus simply consider it in its most simple manifestation, when it is a 0-brane in 4 dimensions. Going to higher dimensions is then simply a matter of adding to the $D + 1$ dimensional metric p flat compact directions with trivial metric (e.g. a trivial T^p factor).

In 4 dimensions we have a dilaton coupling of $a = \sqrt{3}$ and $\Delta = 4$. The magnetic solution is the following:

$$ds_4^2 = -H^{-\frac{1}{2}}dt^2 + H^{\frac{1}{2}}(dr^2 + r^2d\Omega_2^2), \quad (3.113)$$

$$e^\phi = H^{-\frac{\sqrt{3}}{2}}, \quad F_{\theta\varphi} = Q \sin \theta, \quad H = 1 + \frac{Q}{r}. \quad (3.114)$$

From (3.114) we can write a magnetic potential:

$$A_\varphi = Q(1 - \cos \theta). \quad (3.115)$$

Again, we find that $h_{yy} = H^{-1}$ and that the rescaling to the 5 dimensional Einstein metric leads to $\hat{g}_{\mu\nu} = H^{\frac{1}{2}}g_{\mu\nu}$. The 5 dimensional metric is thus:

$$ds_5^2 = -dt^2 + H^{-1}[dy + Q(1 - \cos\theta)d\varphi]^2 + H(dr^2 + r^2d\Omega_2^2).$$

In D dimensions, we finally have for the metric of a KK monopole:

$$ds_{KKM}^2 = -dt^2 + dy_1^2 + \dots + dy_{D-5}^2 + H^{-1}(dz + A_i dx^i)^2 + H dx_i dx^i, \quad i = 1 \dots 3,$$

$$\partial_i A_j - \partial_j A_i = -\epsilon_{ijk} \partial_k H. \quad (3.116)$$

We recognize that the above metric describes a geometry which is $R_{(t)} \times T^{D-5} \times TN$, where TN is the Euclidean Taub-NUT space [97]. The direction z (which was formerly y) is called the NUT direction, and has to be periodically identified with period $4\pi Q$ in order to avoid a (physical) conical singularity at the origin $r = 0$. Note that the NUT direction is generally not considered as a direction longitudinal to the KK monopole (even if its compactness is necessary for the existence of the monopole), and thus the KK monopole is formally a $(D - 5)$ -brane (see e.g. [53]). This means that in 11 dimensional supergravity we have a KK6, and in type II theories we have a KK5.

This completes the review of all the branes which populate the menagerie of string/M-theory.

3.3 Intersections of branes

In this section we search for new, more complex, solutions to the equations of motion derived from the general action (3.9), which includes several antisymmetric tensor fields. The new solutions we are now focusing on will be different with respect to the ones of Section 3.2 in that they carry several charges. Another new feature will be that the harmonic function characterizing the extremal solutions will be allowed to be a sum of an arbitrary number of contributions, centered at different points. The price to pay will be the loss of spherical symmetry.

The aim of finding new solutions, which are more general, is not merely mathematical. The solutions carrying several charges physically represent elementary p -branes, the building blocks discussed in subsection 3.2.5, which intersect each other with some definite rules. From the way they intersect we will find hints for what are their quantum interactions. Moreover, we will be able to give a prescription for building supersymmetric black holes with non-vanishing semi-classical entropy.

As it is already clear from above, we will focus essentially on extremal configurations. This is motivated by supersymmetry. Indeed, results obtained at the low-energy, classical level can be better extrapolated to the quantum theory if they are somehow protected by some supersymmetry properties, which highly constrain the quantum corrections that could spoil the classical result. Nevertheless, we will present at the end of this section a solution which, even in its extremal limit, breaks all supersymmetries. Its relevance in M-theory is still unclear.

Historically, the prototypes of intersecting brane solutions were the solutions found by Guven [95] in 11 dimensional supergravity. More recently, Papadopoulos and Townsend [98] correctly reinterpreted these solutions as intersecting membranes, and furthermore they used the same principle to build new solutions with intersecting fivebranes. At the same time, general stringy concepts such as the various dualities [20, 17, 40, 38] and the recently discovered D-branes [48] led to several conjectures about the intersections of the branes of M-theory, mainly based on supersymmetry arguments [26] and on the extrapolation, using dualities, of D-brane physics [99, 74, 100]. This work culminated in the proposal by Tseytlin of the ‘harmonic superposition rule’ [101], which generalized further the solutions of [98] to include an independent harmonic function for each set of non-parallel branes. Tseytlin also applied all sorts of dualities to the solutions he presented, thus obtaining a classification of intersecting brane solutions. Other papers that appeared soon after Tseytlin’s proposal are [102], where the harmonic superposition rule was found independently, and [103] where it was used in order to discuss four dimensional black holes. A thorough classification of all intersections down to 2 effective space-time dimensions was given in [104]. The common approach of all these papers was to deduce the expressions for the fields from the harmonic superposition rule and then prove that they are consistent solutions, this task being facilitated either by dualities or by supersymmetry. A somewhat different approach was proposed again by Tseytlin [105], using brane probes in brane backgrounds and asking that the probe feels no attraction nor repulsion. In this way, the intersection rules for the marginal bound states of the branes of M-theory were derived.

The approach followed here, and based on [1], is different from the ones described above in that it aims at deriving the harmonic superposition rule from the equations of motion of a general low-energy effective theory and only in the end it is specialized to the cases of interest in M-theory. In this purely classical (and bosonic) context, we find solutions and general intersection rules which are fully consistent with the ones previously found (or guessed) by duality arguments or by other means. The main advantage of our approach is that all the branes are treated manifestly on the same footing. Note that contemporarily a similar approach was undertaken in [106].

We now set up for solving the equations of motion (3.13)–(3.16) (including the Bianchi identities for the $\mathcal{M} n_I$ -form field strengths). The first step, as at the beginning of Section 3.2, is to write the metric in a simplified form. However we now want to find solutions not necessarily constrained by the large amount of symmetries discussed in subsection 3.2.1. In particular, we still want our solutions to be static and translational invariant in the longitudinal space, but we want to drop the $SO(p)$ ‘internal’ symmetry and the $SO(d)$ ‘external’ symmetry. This is to allow respectively for anisotropic combinations of different branes, and for several branes not necessarily located at the same point in the transverse space.

An important remark is now in order, for the notation not to be confusing. We will call p the number of directions which are translational invariant, also called the longitudinal directions. This however does no longer mean that we necessarily have a p -brane in the solution, i.e. that the solution is charged under an electric $(p+2)$ -form field strength or its magnetic dual. It suffices that we have a collection of \mathcal{N} q_A -branes, all with $q_A \leq p$, which fill out a total internal space of dimension p , when all their intersections are taken into account. All these branes will have a common d dimensional transverse space (also called the overall transverse space). The internal directions which are longitudinal to some branes and transverse to some others, are in general called relative transverse directions. Note that the branes are often generally referred to as ‘ p -branes’, and in this case the ‘ p ’ only reflects that we are talking of objects with arbitrary spatial extension.

The general metric we will take is the following:

$$ds^2 = -B^2(x^a)dt^2 + \sum_{i=1}^p C_i^2(x^a)dy_i^2 + G^2(x^a)\delta_{ab}dx^a dx^b, \quad a, b = 1 \dots d. \quad (3.117)$$

This is a generalization with respect to (3.26) in that the spherical symmetry is no longer required, and with respect to (3.25) also the internal $SO(p)$ symmetry is broken. Let us point out that unlike the metric (3.25) which was found in this form imposing its symmetries and using some of the remaining diffeomorphisms, the metric (3.117) is really an ansatz inspired by the solutions found in the previous section. In particular, imposing isotropicity in the overall transverse space is a very strong ansatz when there are no symmetries there (at least when $d \geq 3$). Also, we impose the metric to be diagonal, while we could have allowed for off-diagonal elements. We thus exclude some of the solutions described in the subsection 3.2.5, namely the waves and the KK monopoles. This is mainly a concession to treatability of the equations of motion. Also, the procedure of KK oxidation, i.e. of elevating to higher dimensions a solution, will not be possible because

we included only one scalar field, the dilaton. Solutions with several scalars which can thus correctly take into account all the moduli arising by dimensional reduction can be found in the literature, see e.g. [82, 107]. It should be noted however that all these solutions with off-diagonal terms in the metric can generically be related to diagonal configurations by U-dualities, and more precisely T-dualities. Now it is well known [38] how T-duality acts on the fields of the 10 dimensional supergravities, and this provides a way of indirectly deducing field configurations which include waves and/or KK monopoles. The dual configurations with and without KK waves and monopoles can be guessed using the dualization properties of the branes listed in Appendix B.

A last remark on the metric (3.117) is that all the functions are taken to depend only on the overall transverse space. This means that for a generic brane involved in the configuration, the fields it produces will not depend on the relative transverse directions. We will see, once the solutions will have been found, how it is possible to link these solutions with reduced dependence on the transverse space with those discussed in the preceding section.

3.3.1 Finding the classical solutions and the intersection rules

As in Section 3.2, the first step towards solving the equations of motion will be to focus on the antisymmetric tensor fields. However now we will not solve completely the set of equations of motion and Bianchi identities (3.15) and (3.16), rather we make a suitable ansatz to render one of the two equations for each n -form completely trivial. The other equation is then a non-trivial equation for the n -form.

In the following, we call an electric ansatz for a n -form an ansatz which trivially solves its Bianchi identities (3.16), while we call a magnetic ansatz an ansatz which solves trivially its equations of motion (3.15). The ansätze will be more general than the ones for a singly charged brane, (3.27) and (3.28). This is because there is no more $SO(p)$ internal symmetry, and thus the constraints on the presence of all or none of the internal indices are relaxed.

Let us start with the electric ansatz. A straightforward generalization of (3.27) is the following:

$$F_{ti_1 \dots i_{q_A} a} = \epsilon_{i_1 \dots i_{q_A}} \partial_a E_A(x). \quad (3.118)$$

This is thus an ansatz for a $(q_A + 2)$ -form which gives an electric charge to a q_A -brane, and the index A runs over all the \mathcal{N} branes. Note that the number of branes can be (and generally is) larger than the number \mathcal{M} of antisymmetric tensor fields. Several branes charged under the same field strength will be distinguished by the directions in which they lie, i.e. the directions labelled by $\{i_1 \dots i_{q_A}\}$ in the case above. The ansatz above is inspired by the fact that the $(q_A + 1)$ -form potential must couple to the world-volume of the brane, and should thus contain all the indices of the directions tangential to it.

A magnetic ansatz generalizing (3.28) is readily found considering the Hodge duality (3.21). The $(D - q_A - 2)$ -form field strength must contain $d - 1$ indices of the overall transverse space and all the indices of the relative transverse (internal) space:

$$F^{i_{q_A+1} \dots i_p a_1 \dots a_{d-1}} = \epsilon^{i_{q_A+1} \dots i_p} \epsilon^{a_1 \dots a_d} \frac{1}{\sqrt{-g}} e^{-a_A \phi} \partial_{a_d} E_A(x). \quad (3.119)$$

With respect to (3.28) we have dropped the tilde on the function E_A , because it will soon become clear that it will play exactly the same rôle as the function appearing in (3.118). The coupling a_A refers to the field strength, but for simplicity it carries the label of the brane. The above ansätze can be trivially seen to solve respectively the Bianchi identities (3.16) and the equations of motion (3.15).

In order to rewrite the equations of motion derived from the action (3.9) taking into account the ansätze (3.117) and (3.118) and/or (3.119), we first introduce some notation. Let us define the two quantities:

$$V_A = B \prod_{i \parallel A} C_i, \quad V_A^\perp = \prod_{i \perp A} C_i, \quad (3.120)$$

where the first product runs over all the internal directions parallel to the brane labelled by A and the second one over the internal directions transverse to it. Let us also define the following quantities:

$$\varepsilon_A = \begin{cases} + & \text{if } q_A\text{-brane is electric} \\ - & \text{if } q_A\text{-brane is magnetic} \end{cases} \quad (3.121)$$

and:

$$\delta_A^i = \begin{cases} D - q_A - 3 & \text{for } i \parallel A \\ -(q_A + 1) & \text{for } i \perp A \end{cases} \quad (3.122)$$

The set of equations (3.13)–(3.16) then becomes:

$$R_t^t = -\frac{1}{2} \sum_A \frac{D - q_A - 3}{D - 2} (V_A G)^{-2} e^{\varepsilon_A a_A \phi} (\partial E_A)^2, \quad (3.123)$$

$$R_i^i = -\frac{1}{2} \sum_A \frac{\delta_A^i}{D - 2} (V_A G)^{-2} e^{\varepsilon_A a_A \phi} (\partial E_A)^2, \quad (3.124)$$

$$R_b^a = \frac{1}{2} G^{-2} \partial_a \phi \partial_b \phi + \frac{1}{2} \sum_A (V_A G)^{-2} e^{\varepsilon_A a_A \phi} \left[-\partial_a E_A \partial_b E_A + \frac{q_A + 1}{D - 2} \delta_{ab} (\partial E_A)^2 \right], \quad (3.125)$$

$$\square \phi = -\frac{1}{2} \sum_A \varepsilon_A a_A (V_A G)^{-2} e^{\varepsilon_A a_A \phi} (\partial E_A)^2, \quad (3.126)$$

$$\partial_a \left(\frac{V_A^\perp}{V_A} G^{d-2} e^{\varepsilon_A a_A \phi} \partial_a E_A \right) = 0. \quad (3.127)$$

We use the components of the Ricci tensor as given in the equations (C.25)–(C.27) of Appendix C, we have defined $(\partial E_A)^2 = \partial_a E_A \partial_a E_A$ and $\square \phi$ is given by:

$$\square \phi = \frac{1}{G^2} \left\{ \partial_a \partial_a \phi + \partial_a \phi \partial_a \left[\ln B + \sum_i \ln C_i + (d - 2) \ln G \right] \right\}.$$

The last equation (3.127) derives either from the equations of motion (3.15) or from the Bianchi identities (3.16), depending on the electric or magnetic nature of the brane labelled by A .

One remark is in order before we continue. In deriving (3.124), we have implicitly assumed that there are no two branes in the configuration which are both charged under the same (electric or magnetic) field strength and which intersect over all but one of their world-volume directions. Such a configuration would give off-diagonal elements to the $\binom{i}{j}$ Einstein equations. However our metric ansatz (3.117) excludes non-trivial solutions of this kind, since the Ricci tensor components R_j^i are diagonal.

We are now ready to perform some further simplifications, or further ansätze, in order to actually find a class of solutions to the equations above. The first simplification is motivated by the structure of the Ricci tensor as given in (C.25)–(C.27), and is indeed a generalization of the relation (3.40) in the case of a singly charged brane. We impose on the metric components the following condition:

$$BC_1 \dots C_p G^{d-2} = 1. \quad (3.128)$$

This condition greatly simplifies the expression for the Ricci tensor, and moreover it is consistent with the equations of motion simply by the fact that we will indeed find some solutions. It is slightly more complicated to see that this condition singles out (at least from the known solutions) the brane configurations which are extremal. To see this, however, it is better to restore spherical symmetry and to rewrite the condition in a r -reparametrization covariant way, suitable for a general metric like (3.26). One finds that the condition (3.128) is equivalent to imposing that:

$$\frac{B'}{B} + \sum_i \frac{C'_i}{C_i} + (d-2) \left(\frac{G'}{G} + \frac{1}{r} - \frac{F}{Gr} \right) = 0.$$

It is easy to check that the above condition holds for the metric (3.70) or (3.71) if and only if the non-extremality parameter μ vanishes. We will thus often refer to (3.128) as the ‘extremality condition’.

The equations (3.123)–(3.127) now become:

$$\partial_a \partial_a \ln B = \frac{1}{2} \sum_A \frac{D - q_A - 3}{D - 2} S_A^2 (\partial E_A)^2, \quad (3.129)$$

$$\partial_a \partial_a \ln C_i = \frac{1}{2} \sum_A \frac{\delta_A^i}{D - 2} S_A^2 (\partial E_A)^2, \quad (3.130)$$

$$\begin{aligned} \partial_a \ln B \partial_b \ln B + \sum_i \partial_a \ln C_i \partial_b \ln C_i + (d-2) \partial_a \ln G \partial_b \ln G + \frac{1}{2} \partial_a \phi \partial_b \phi \\ + \delta_{ab} \partial_c \partial_c \ln G = \frac{1}{2} \sum_A S_A^2 \left[\partial_a E_A \partial_b E_A - \frac{q_A + 1}{D - 2} \delta_{ab} (\partial E_A)^2 \right], \end{aligned} \quad (3.131)$$

$$\partial_a \partial_a \phi = -\frac{1}{2} \sum_A \varepsilon_A a_A S_A^2 (\partial E_A)^2, \quad (3.132)$$

$$\partial_a (S_A^2 \partial_a E_A) = 0, \quad (3.133)$$

where we have defined:

$$S_A^2 = \frac{1}{V_A^2} e^{\varepsilon_A a_A \phi}. \quad (3.134)$$

It is worth noting that the condition (3.128) can be written for each brane $V_A V_A^\perp G^{d-2} = 1$, and it is indeed by virtue of this relation that the same combination S_A appears both in the Einstein equations and in the Maxwell-like equation (3.133).

The second, crucial, simplification is to reduce the number of independent functions to \mathcal{N} . This is motivated by the requirement that all the branes constituting the configuration are extremal, and that moreover there is no binding energy between them [1, 3]. They thus define a BPS marginal bound state obeying the no-force condition [105]. This means that each brane can be pulled apart from the others at zero cost in energy, until the fields around it are a good approximation of the fields describing a singly charged brane solution. Now such a solution is fully characterized by a single (harmonic) function. It is thus expected that the solution with \mathcal{N} intersecting branes is characterized by \mathcal{N} independent (harmonic) functions.

Since we have already \mathcal{N} functions E_A , it is natural to take them for the moment as the only independent functions in the problem. For the Maxwell-like equations (3.133) not to couple these functions in a non-trivial way, we have to assume that the combinations S_A only depend on the E_A carrying the same label:

$$S_A \sim E_A^\gamma.$$

Then (3.133) directly implies that $E_A^{2\gamma+1}$ is a harmonic function (for $\gamma \neq -\frac{1}{2}$). Now, according to the discussion above, if all the other functions have to be expressed in terms of the E_A 's, and should reproduce the single brane solutions in the appropriate limit, then we have to take for B :

$$\ln B = \sum_A \tilde{\alpha}_A^{(B)} \ln E_A,$$

and similarly for the other functions C_i , G and ϕ . The equation (3.129) reduces to an algebraic equation for the coefficients $\tilde{\alpha}_A^{(B)}$ if and only if $\gamma = -1$, i.e. if $S_A^2 (\partial E_A)^2 \sim (\partial \ln E_A)^2$. If we had $\gamma \neq -1$, the E_A 's would be either constant or not truly independent.

We are thus forced to take the following ansatz, which introduces the new functions H_A :

$$E_A = l_A H_A^{-1}, \quad S_A = H_A. \quad (3.135)$$

The constant l_A is to be determined shortly, by the equations of motion.

The equations (3.133) simply state now that the H_A 's are harmonic:

$$\partial_a \partial_a H_A = 0. \quad (3.136)$$

The most general solution to the above equation is the following:

$$H_A = 1 + \sum_k c_A \frac{Q_{A,k}}{|x^a - x_k^a|^{d-2}}, \quad (3.137)$$

where, requiring asymptotic flatness, we fix the overall factor to 1; the c_A 's are constants to be determined when one computes the charge of the branes, and $\{x_k\}$ are the locations

in transverse space of several branes of the same type, i.e. parallel. The harmonic functions above thus represent a multi-center solution.

Note that the solution (3.137) assumes that $d \geq 3$. We will comment afterwards on the cases $d = 2$ and $d = 1$.

We now express the functions B , C_i and ϕ in terms of the H_A 's:

$$\begin{aligned}\ln B &= \sum_A \alpha_A^{(B)} \ln H_A, \\ \ln C_i &= \sum_A \alpha_A^{(i)} \ln H_A, \\ \phi &= \sum_A \alpha_A^{(\phi)} \ln H_A.\end{aligned}\tag{3.138}$$

The function G will be determined later, using (3.128). The equations (3.129), (3.130) and (3.132) reduce to equations for the coefficients appearing in (3.138), when (3.136) are taken into account. We find that:

$$\begin{aligned}\alpha_A^{(B)} &= -\frac{D - q_A - 3}{D - 2} \alpha_A, \\ \alpha_A^{(i)} &= -\frac{\delta_A^i}{D - 2} \alpha_A, \\ \alpha_A^{(\phi)} &= \varepsilon_A a_A \alpha_A,\end{aligned}\tag{3.139}$$

where:

$$\alpha_A = \frac{1}{2} l_A^2.\tag{3.140}$$

For $d \neq 2$, (3.128) and (3.138) imply that:

$$\ln G = \sum_A \frac{q_A + 1}{D - 2} \alpha_A \ln H_A.\tag{3.141}$$

This trivially solves the part in the equations (3.131) which is proportional to δ_{ab} . The remaining piece of the same equations (3.131) gives us the value of α_A , and surprisingly gives as a set of consistency conditions the pairwise intersection rules for the branes involved in the configuration.

The set of equations (3.131) becomes:

$$\begin{aligned}\sum_{A,B} \alpha_A \alpha_B \partial_a \ln H_A \partial_b \ln H_B \left[\frac{(D - q_A - 3)(D - q_B - 3)}{(D - 2)^2} + \sum_{i=1}^p \frac{\delta_A^i \delta_B^i}{(D - 2)^2} \right. \\ \left. + (d - 2) \frac{(q_A + 1)(q_B + 1)}{(D - 2)^2} + \frac{1}{2} \varepsilon_A a_A \varepsilon_B a_B \right] = \sum_A \alpha_A \partial_a \ln H_A \partial_b \ln H_A.\end{aligned}\tag{3.142}$$

It can be rewritten as:

$$\sum_{A,B} (M_{AB} \alpha_A - \delta_{AB}) \alpha_B \partial_a \ln H_A \partial_b \ln H_B = 0.$$

For the H_A 's to be truly independent, we must have:

$$M_{AB}\alpha_A = \delta_{AB},$$

which implies two sets of conditions, $M_{AA}\alpha_A = 1$ and $M_{AB} = 0$ for $A \neq B$. The first set of equations assigns a value to α_A :

$$\alpha_A = (M_{AA})^{-1} = \frac{D-2}{\Delta_A}, \quad (3.143)$$

where we have defined as before:

$$\Delta_A = (q_A + 1)(D - q_A - 3) + \frac{1}{2}a_A^2(D - 2). \quad (3.144)$$

We have now an expression for all the parameters in the solution, and we can thus use (3.143), (3.140) and (3.139) together with (3.135), (3.138) and (3.141) to characterize completely the fields of the solution:

$$B = \prod_A H_A^{-\frac{D-q_A-3}{\Delta_A}}, \quad C_i = \prod_A H_A^{-\frac{\delta_A^i}{\Delta_A}}, \quad G = \prod_A H_A^{\frac{q_A+1}{\Delta_A}}, \quad (3.145)$$

$$e^\phi = \prod_A H_A^{\varepsilon_A a_A \frac{D-2}{\Delta_A}}, \quad E_A = \sqrt{\frac{2(D-2)}{\Delta_A}} H_A^{-1}. \quad (3.146)$$

The E_A 's coincide with the potentials, up to a constant if the latter are to vanish at infinity. We have fixed the overall sign to a positive value. Note that there is only one sign to fix for each set of branes. Accordingly, all the branes of each type have to have only positive (or only negative) charges.

The above derivation thus proves the 'harmonic superposition rule' as formulated by Tseytlin in [101]. Note that the Δ_A are strictly the same as in the singly charged solution (3.49), thus the solution above is really the superposition of single brane solutions.

We are however left now with an additional set of consistency conditions, i.e. the conditions $M_{AB} = 0$. Since there are no more parameters to adjust, these conditions will really tell us whether the above solution is a solution of the equations of motion or not. Happily, it is easy to see that we can give a simple expression for M_{AB} when $A \neq B$:

$$\begin{aligned} M_{AB} &= \frac{1}{(D-2)^2} \left[(D - q_A - 3)(D - q_B - 3) + \bar{q} (D - q_A - 3)(D - q_B - 3) \right. \\ &\quad - (q_A - \bar{q})(D - q_A - 3)(q_B + 1) - (q_B - \bar{q})(q_A + 1)(D - q_B - 3) \\ &\quad \left. + (p - q_A - q_B + \bar{q})(q_A + 1)(q_B + 1) + (D - p - 3)(q_A + 1)(q_B + 1) \right] \\ &\quad + \frac{1}{2} \varepsilon_A a_A \varepsilon_B a_B \\ &= \bar{q} + 1 - \frac{(q_A + 1)(q_B + 1)}{D - 2} + \frac{1}{2} \varepsilon_A a_A \varepsilon_B a_B, \end{aligned} \quad (3.147)$$

where we have defined \bar{q} the dimension of the intersection between the brane labelled by A and the one labelled by B ($\bar{q} \leq q_A, q_B$), and we have rewritten $d = D - p - 1$.

The consistency condition is thus a restriction on the dimension of the intersection for each pair of branes in the compound:

$$\bar{q} + 1 = \frac{(q_A + 1)(q_B + 1)}{D - 2} - \frac{1}{2} \varepsilon_A a_A \varepsilon_B a_B. \quad (3.148)$$

This relation can be interpreted as an equation for \bar{q} . It thus gives the dimension of the intersection of two branes, given only the dimension of the branes, their coupling to the dilaton and the requirement that they should form a bound state with vanishing binding energy. Note that the intersection rules above are perfectly pairwise, in the sense that they are not influenced by the presence of other branes in the configuration. For instance, we see that in (3.147), even the value of p , the total number of ‘internal’ directions, disappears from the final expression.

Note that asking that S_A as given in (3.134) verifies $S_A = H_A$, and making use of the relations (3.138) and (3.139), we hopefully obtain the same equations (3.143) and (3.148).

The $d = 2$ and $d = 1$ cases Let us now comment the cases that we have temporarily excluded from the derivation above, i.e. the cases for which $d = 2$ or $d = 1$. These cases are relevant for the physics in an effective three and two dimensional space-time respectively. However, despite the somewhat pathological physics in these lower dimensions, the results will be very similar to the $d \geq 3$ case. Most of all, the intersection rules (3.148) are insensitive to this issue.

The derivation of the intersection rules in the case $d = 2$ is a little bit more involved, mainly because of the several $(d - 2)$ factors that appear in the equations. In particular, the function G does not appear any more in the ‘extremality’ ansatz (3.128). This has as a consequence that the relation (3.141) is no longer implied by the other relations (3.138) and (3.139), but has to be derived directly from (3.131), noting that if the off-diagonal part of this set of equations is to vanish generically, then its part proportional to δ_{ab} also has to vanish independently. Once this trick is understood, the solution is then easily shown to be exactly the same as before, with the exception that now the harmonic solutions of (3.136) have a logarithmic behaviour:

$$H_A^{(d=2)} = h_A + \sum_k c_A Q_{A,k} \ln |x^a - x_k^a|.$$

The solutions can be expressed formally as in (3.145), however one has to keep in mind that the harmonic functions H_A diverge at infinity. This points towards a more careful treatment of this problem, possibly along the lines of [108, 109]. For the extrapolation of the intersection rules from supergravity to M-theory, it will be nevertheless useful to know that these rules still hold in this case.

The $d = 1$ case is technically simpler than the preceding one. Here the derivation is exactly the same as in the $d \geq 3$ case, except that the partial derivatives are now simple

derivatives. The intersection rules and the solution (3.145) are (formally) unchanged, provided we now write for H_A :

$$H_A^{(d=1)} = h_A + \sum_k c_A Q_{A,k}(x - x_k).$$

Note that branes with only one transverse direction are actually domain walls, and these in general require the supergravity of which they are solutions to be massive. If however there is no such $(D - 2)$ -brane in the configuration, no cosmological constant seems to be required.

Euclidean branes As a last remark on the derivation of the solution (3.145)–(3.146) and of the intersection rules, we note that it extends also to Euclidean configurations of branes. The time is no longer singled out, and it can be considered as one of the y_i coordinates. There is thus no longer a timelike direction that has to be common to all the branes. The equations one obtains are exactly the same as (3.129)–(3.133), except that there is always an additional factor of $-\varepsilon_A$ appearing inside the sum in each r.h.s. This is because, since now $g_{tt} > 0$, the stress-energy tensor of an electric field has the opposite sign with respect to the Lorentzian case.

The derivation goes along the same way as for the Lorentzian case, and the only result which differs is (3.140). It becomes:

$$\alpha_A = -\varepsilon_A \frac{1}{2} l_A^2.$$

As a consequence, the fields E_A are given by:

$$E_A = \sqrt{-\varepsilon_A \frac{2(D-2)}{\Delta_A}} H_A^{-1}.$$

This means that the electric fields are purely imaginary. This is however a well-known trick to make sense of electrically charged black holes when one goes to Euclidean gravity (see e.g. [90] or [108]).

Note that the condition on the intersection (3.148) has now potentially more solutions, since \bar{q} can now take also the value $\bar{q} = -1$ (thus defining an intersection on a point in Euclidean space). Also, a single brane solution which only exists in the Euclidean formulation is the (-1) -brane, or instanton.

3.3.2 Further general considerations on intersecting brane solutions

In this subsection we try to collect some of the direct physical consequences of the intersecting brane solutions found previously. In the next subsection we discuss the intersection rules in the context of string/M-theory.

We have first to elaborate a little bit on the remark that the solution (3.145)–(3.146) is a superposition of single extremal brane solution as the one treated in subsection 3.2.2.

There is a difference between the single extremal brane solution (3.49) and the solution (3.145) in which all but one of the branes have been put to zero charge: in the second solution the dimension of the transverse space d can be smaller than $D - q - 1$, where q is the dimension of the brane. Are both solutions physically meaningful and how are they related to each other?

The answer to this question is well known and the explanation involves taking the transverse space of the q -brane to have some compact directions. We present hereafter the argument for one additional compact direction, the generalization is straightforward by induction.

Suppose we have a q -brane and we have defined $\bar{d} = D - q - 1$. However only $d = \bar{d} - 1$ of the space-like directions are non-compact, while the extra direction is periodically identified $\bar{x} \sim \bar{x} + L$. The topology of the transverse space is now $S^1 \times R^d$. A q -brane sitting at one point of this transverse space is seen from the point of view of the covering space of the S^1 factor as an equally spaced array of q -branes in the \bar{x} direction, at the same point in the R^d space.

This is correctly described by a multi-center solution obtained from (3.145)–(3.146) keeping only one non-trivial charge. The harmonic function characterizing the solution reads:

$$H = 1 + \sum_k \frac{Q}{|x - x_k|^{\bar{d}-2}} = 1 + \sum_k \frac{Q}{[r^2 + (\bar{x} - kL)^2]^{\frac{\bar{d}-2}{2}}}, \quad r^2 = x_1^2 + \dots + x_d^2.$$

One can approximate the sum assuming that the distance in non-compact space is much larger than the size of the compact direction, $\frac{L}{r} \ll 1$. Indeed we have:

$$\begin{aligned} \sum_k \frac{1}{[r^2 + (\bar{x} - kL)^2]^{\frac{n}{2}}} &= \frac{1}{r^n} \sum_k \frac{1}{\left[1 + \left(\frac{\bar{x}}{r} - k\frac{L}{r}\right)^2\right]^{\frac{n}{2}}} \\ &\simeq \frac{1}{r^n} \frac{r}{L} \int_{-\infty}^{\infty} du \frac{1}{(1 + u^2)^{\frac{n}{2}}} = \frac{1}{Lr^{n-1}} \int_0^\pi d\theta \sin^{n-2} \theta, \end{aligned}$$

where we have made the change of variables $u = k\frac{L}{r} - \frac{\bar{x}}{r}$ and then $u = \cot \theta$. The first change of variables, due to the smallness of $\frac{L}{r}$, changes the sum in a continuous integral. The integral:

$$I_m = \int_0^\pi d\theta \sin^m \theta$$

verifies the following identity:

$$I_m = \frac{m-1}{m} I_{m-2}, \quad m \geq 2.$$

The two basic values are $I_1 = 2$ and $I_0 = \pi$. We can also express the volume of a $(d-1)$ -sphere in terms of it:

$$\Omega_{d-1} = 2I_{d-2}I_{d-3} \dots I_1I_0 = \frac{2\pi^{\frac{d}{2}}}{\Gamma\left(\frac{d}{2}\right)}.$$

The harmonic function then becomes:

$$H = 1 + \frac{I_{d-3}}{L} \frac{Q}{r^{d-2}}.$$

We thus see that ‘compactifying’ one of the transverse directions, and assuming that we are not too close to the location of the brane, we recover a harmonic function in the non-compact space.

We can also show that the charge density, computed as in (3.93), is the same for the uncompactified and the compactified spaces if the harmonic functions are related as above. Indeed,

$$\mathcal{Q} = \frac{1}{16\pi G_D} \int_{S^{\bar{d}-1}} *F_{q+2} = (d-1)\Omega_d \tilde{\mathcal{Q}}, \quad \tilde{\mathcal{Q}} = \frac{1}{16\pi G_D} \sqrt{\frac{2(D-2)}{\Delta}} Q,$$

for the brane in non compact $R^{\bar{d}}$ space. In the compact space on the other hand:

$$\mathcal{Q}' = \frac{1}{16\pi G_D} \int_{S^1 \times S^{d-1}} *F_{q+2} = (d-2)L\Omega_{d-1} \frac{I_{d-3}}{L} \tilde{\mathcal{Q}}.$$

The two charge densities \mathcal{Q} and \mathcal{Q}' are indeed equal:

$$\mathcal{Q}' = \mathcal{Q} \frac{(d-1)\Omega_d}{(d-2)\Omega_{d-1}I_{d-3}} = \mathcal{Q} \frac{(d-1)I_{d-1}}{(d-2)I_{d-3}} = \mathcal{Q}.$$

The procedure that we have shown above can be straightforwardly continued to other compact transverse directions. Note that the solutions found in the previous subsection are exact solutions of the classical equations of motion. In this case, it is often said that the solutions are ‘smeared’ over the transverse compact directions, in order to eliminate the dependence of the fields in these directions. The procedure discussed above shows that the ‘smeared’ solutions are classical approximations to the (quantum) solutions, in the sense that an average is taken on the location in the compact transverse directions. Properly speaking, the above solutions are good ‘long distance’ approximations only up to distances of the order of the compactification length. Alternatively, one can distribute the branes along the compact directions, keeping the total charge fixed. In this way, the charge will look like a uniform distribution at a closer distance. See [85] where the procedure above, considered however in the reverse order, leads to the higher dimensional resolution of otherwise singular geometries.

After this rather lengthy but necessary parenthesis, we are now totally able to discuss the physics of the intersecting brane configurations.

Let us recall the metric:

$$ds^2 = \prod_A H_A^{-2\frac{D-q_A-3}{\Delta_A}} dt^2 + \sum_i \prod_A H_A^{-2\frac{\delta_A^i}{\Delta_A}} dy_i^2 + \prod_A H_A^{2\frac{q_A+1}{\Delta_A}} \delta_{ab} dx^a dx^b, \quad (3.149)$$

and the field strengths:

$$F_{ti_1 \dots i_{q_A} a} = \epsilon_{i_1 \dots i_{q_A}} \sqrt{\frac{2(D-2)}{\Delta_A}} \partial_a H_A^{-1}, \quad \text{if } \varepsilon_A = (+), \quad (3.150)$$

$$F_{i_{q_A+1} \dots i_p a_1 \dots a_{d-1}} = -\epsilon_{i_{q_A+1} \dots i_p} \epsilon_{a_1 \dots a_{d-1}} \sqrt{\frac{2(D-2)}{\Delta_A}} \partial_{a_d} H_A, \quad \text{if } \varepsilon_A = (-), \quad (3.151)$$

where the harmonic functions are as in (3.137). The charge densities are defined as in the preceding section by:

$$\mathcal{Q}_A = \frac{1}{16\pi G_D} \int_{T^{p-q_A} \times S^{d-1}} *F_{q_A+2}, \quad \text{if } \varepsilon_A = (+), \quad (3.152)$$

$$\mathcal{Q}_A = \frac{1}{16\pi G_D} \int_{T^{p-q_A} \times S^{d-1}} F_{D-q_A-2}, \quad \text{if } \varepsilon_A = (-). \quad (3.153)$$

Both of the above definitions give, when taking into account respectively (3.150) and (3.151), the following value for the charge density:

$$\mathcal{Q}_A = \frac{L^{p-q_A} \Omega_{d-1}}{16\pi G_D} (d-2) \sqrt{\frac{2(D-2)}{\Delta_A}} c_A \sum_k Q_{A,k}, \quad (3.154)$$

where L is the (common) size of all the internal directions. As noted before, all the charges $Q_{A,k}$ have to be of the same sign³, which here is taken to be positive.

It is then straightforward to compute the ADM mass as in (3.82) (the multi-center solutions are approximately spherically symmetric when the branes are confined to a certain region of transverse space) and to express it in terms of the charge densities (3.154):

$$M = \sum_A \sqrt{\frac{2(D-2)}{\Delta_A}} L^{q_A} \mathcal{Q}_A. \quad (3.155)$$

This formula actually proves the assumption that our solutions represent bound states of branes with vanishing binding energy. Indeed, each term in the sum above is nothing else than the mass of each constituent extremal q_A -brane, as in (3.95).

This is an important result: the two mathematical conditions of imposing first the relation (3.128) and then the formulation in terms of \mathcal{N} independent function were correctly interpreted as enforcing extremality and the no-force condition.

If we consider a configuration in which all the branes are located at the same point in the overall transverse space (this is truly a bound state), we have a spherically symmetric configuration and we can moreover compute the entropy using the formula (3.90), with the horizon coinciding with $r = 0$. We find:

$$S = \frac{L^p \Omega_{d-1}}{4G_D} \prod_A H_A^{\frac{D-2}{\Delta_A}} r^{d-1} \Big|_{r \rightarrow 0} = \frac{L^p \Omega_{d-1}}{4G_D} \prod_A (c_A Q_A)^{\frac{D-2}{\Delta_A}} r^{d-1 - \sum_A (d-2) \frac{D-2}{\Delta_A}} \Big|_{r \rightarrow 0}. \quad (3.156)$$

The actual value of the power of r depends on the number and (presumably) on the nature of the branes involved in the bound state, and it determines whether the solution has a finite entropy or not. We will see shortly how to achieve a finite entropy in configurations of extreme branes in M-theory.

³ This can also be motivated by supersymmetry (see subsection 3.3.4), and from D-brane physics [26, 110].

3.3.3 Applications to M-theory

It is now time to specialize to the intersecting brane solutions that one finds in 11 dimensional supergravity and in 10 dimensional string theories. We will first consider extensively the theories with maximal supersymmetry ($D = 11$ and type II theories), and in the end comment on the theories with less supersymmetry.

Let us start with the intersections of the branes of 11 dimensional supergravity, also called M-branes. Taking into account that we only have one 4-form field strength and no dilaton, the formula for the intersection (3.148) rewrites simply:

$$\bar{q} + 1 = \frac{(q_A + 1)(q_B + 1)}{9}. \quad (3.157)$$

Adopting the obvious notation $q_A \cap q_B = \bar{q}$, we have the following intersections involving the two M-branes, M2 and M5:

$$M2 \cap M2 = 0, \quad M5 \cap M5 = 3, \quad M2 \cap M5 = 1. \quad (3.158)$$

The supergravity solutions representing special cases of the first two intersections above were derived in [98], using the solutions previously found in [95]. The third intersection was postulated in [99, 74] from D-brane and duality arguments, and the classical solution relative to it was shown in [101, 102].

The most intriguing intersection is the third one. According to the discussion in the previous subsection, the supergravity solution is ‘smeared’ over all the longitudinal 6-dimensional space. The M2-branes are not localized on the M5-brane and vice-versa. If however we extrapolate the solution to a configuration in which both branes are infinite and well-localized, then we arrive at the picture of an M5-brane cutting an M2-brane in two pieces, since the intersection has the dimension of the boundary of the M2-brane. A second, more speculative step is to consider that the two parts of the M2-brane can break up, leaving two open M2-branes with their boundaries attached to the M5-brane. This is a familiar picture in D-brane physics, the novelty being that also non-perturbative objects like the M-branes behave like the fundamental strings. In the next chapter this picture will be made more precise, but additional input will be needed.

Note that in [102] a configuration giving $M5 \cap M5 = 1$ is discussed. However the harmonic functions have a different dependence than in the solutions discussed in this section. We will comment on this different case in Section 3.4.

Getting one dimension lower, we find ourselves in the low-energy effective theories of type II strings. Consider first of all the intersections of the RR-charged branes. Since the coupling to the dilaton is $\varepsilon a = \frac{3-q}{2}$, the intersection rules can be rewritten as:

$$\bar{q} = \frac{1}{2}(q_A + q_B - 4). \quad (3.159)$$

This rule gives the following intersections (where the D is conventional):

$$Dq \cap Dq = q - 2, \quad (3.160)$$

$$D(q - 2) \cap Dq = q - 3, \quad (3.161)$$

$$D(q - 4) \cap Dq = q - 4. \quad (3.162)$$

The intersection (3.161) was also conjectured in [99], and the same discussion as for the $M2 \cap M5 = 1$ case applies. The last intersection (3.162) actually represents a $D(q-4)$ -brane within a Dq -brane, as suggested in [73].

There are some interesting remarks that can be made about the above intersections of branes. Note first that the intersection $D0 \cap D6$ makes no sense at all, since it would give $\bar{q} = 1!$ This means that the $D0$ -brane and the $D6$ -brane cannot form a marginal bound state, and this applies also to all the configurations which are dual to this one. The second remark is that there are several intersections giving $\bar{q} = -1$, e.g. $D1 \cap D1 = -1$. Though meaningless in ordinary space-time, these configurations are nevertheless relevant in Euclidean space, and can for instance represent instantons.

The last set of branes we consider are the NSNS ones, for which $a = -1$. We have the following intersections involving the fundamental string $F1$ and the solitonic fivebrane $NS5$:

$$F1 \cap F1 = -1, \quad F1 \cap NS5 = 1, \quad NS5 \cap NS5 = 3, \quad (3.163)$$

$$F1 \cap Dq = 0, \quad (3.164)$$

$$Dq \cap NS5 = q - 1, \quad q \leq 6. \quad (3.165)$$

The three intersections (3.163) are perfectly consistent through type IIB S-duality with the intersections involving the $D1$ and the $D5$.

The relation (3.164) seems really a classical ‘hint’ that fundamental strings can end on RR-charged branes, which should then consequently be called D(irichlet)-branes. Of course historically the discovery was done the other way round, defining the D-branes in (perturbative) string theory and then showing that they are the carriers of RR charge [48]. However here the last relation (3.165) seems to imply further that all the D-branes themselves can have their boundaries on the $NS5$. Again, this will be elaborated in the next chapter.

The concluding remark about all the intersections that we have discussed above is that they were all derived in a strikingly similar way, simply applying the formula (3.148) to the pair of branes at hand. All the intersections are on an equal footing within this classical approach, while they have a clearly different origin in any one of the theories in which they can be embedded. Note the other approach discussed in [105], which however only applies to the string/M-theory framework.

We only have presented above pairwise intersections. As pointed out in the derivation of (3.148), one can build configurations with more than two branes simply imposing the intersection rules on each pair of branes in the compound.

It is straightforward but lengthy to prove that the Chern-Simons terms present in the original actions (3.3), (3.6) and (3.8) modify the equations of motion with terms which vanish identically for the solutions corresponding to the intersections above. We can thus confidently consider the intersecting brane solutions as configurations in M-theory. The Chern-Simons terms will however turn out to play a non-trivial rôle when we will discuss more accurately the possibility of having open branes.

We now turn to the problem of finding bound states of branes with a non-vanishing entropy (3.156).

As far as branes in M-theory are concerned, we had already noted in subsection 3.2.5 that they all verify $\Delta_A = 2(D-2)$. Note that this is insensitive to the possible existence of additional compact directions transverse to the particular brane labelled by A . This fact simplifies the expression for the entropy which becomes, when \mathcal{N} branes participate to the bound state:

$$S \sim \prod_A Q_A^{\frac{1}{2}} r^{d-1-\frac{1}{2}\mathcal{N}(d-2)} \Big|_{r \rightarrow 0}.$$

If we define $\bar{D} = D - p = d + 1$, the condition for a finite entropy is:

$$\mathcal{N} = 2 \frac{\bar{D} - 2}{\bar{D} - 3}. \quad (3.166)$$

For $4 \leq \bar{D} \leq 11$ there are only two cases which give an integer \mathcal{N} , and these are:

$$\bar{D} = 5, \quad \mathcal{N} = 3 \quad \text{and} \quad \bar{D} = 4, \quad \mathcal{N} = 4. \quad (3.167)$$

Restoring the right coefficients, the entropy in these two cases is:

$$S = \frac{L^p \Omega_{d-1}}{4G_D} \prod_A (c_A Q_A)^{\frac{1}{2}}. \quad (3.168)$$

As a straightforward exercise, we can write the metrics of a $\bar{D} = 5$ and a $\bar{D} = 4$ black hole. For the $\bar{D} = 5$ configuration, we take the one in 11 dimensions consisting of three different types of M2-branes, all intersecting each other in one point. The metric reads:

$$\begin{aligned} ds_5^2 = & -(H_1 H_2 H_3)^{-\frac{2}{3}} dt^2 + H_1^{-\frac{2}{3}} (H_2 H_3)^{\frac{1}{3}} (dy_1^2 + dy_2^2) + H_2^{-\frac{2}{3}} (H_1 H_3)^{\frac{1}{3}} (dy_3^2 + dy_4^2) \\ & + H_3^{-\frac{2}{3}} (H_1 H_2)^{\frac{1}{3}} (dy_5^2 + dy_6^2) + (H_1 H_2 H_3)^{\frac{1}{3}} (dr^2 + r^2 d\Omega_3^2). \end{aligned} \quad (3.169)$$

For the $\bar{D} = 4$ black hole, we take its realization in type IIA theory as a bound state of three D4-branes, intersecting each other over 2 dimensions, and one D0-brane:

$$\begin{aligned} ds_4^2 = & -H_{D0}^{-\frac{7}{8}} (H_{D41} H_{D42} H_{D43})^{-\frac{3}{8}} dt^2 + H_{D0}^{\frac{1}{8}} (H_{D41} H_{D42})^{-\frac{3}{8}} H_{D43}^{\frac{5}{8}} (dy_1^2 + dy_2^2) \\ & + H_{D0}^{\frac{1}{8}} (H_{D41} H_{D43})^{-\frac{3}{8}} H_{D42}^{\frac{5}{8}} (dy_3^2 + dy_4^2) + H_{D0}^{\frac{1}{8}} (H_{D42} H_{D43})^{-\frac{3}{8}} H_{D41}^{\frac{5}{8}} (dy_5^2 + dy_6^2) \\ & + H_{D0}^{\frac{1}{8}} (H_{D41} H_{D42} H_{D43})^{\frac{5}{8}} (dr^2 + r^2 d\Omega_2^2). \end{aligned} \quad (3.170)$$

Starting from these two solutions, and using the U-dualities, one can obtain all the configurations with respectively 3 charges in 5 dimensions and 4 charges in 4 dimensions. Note two interesting facts: first, the equation (3.166) implies that there are no ‘stringy’ extreme black holes with non-vanishing entropy in $\bar{D} \geq 6$; secondly, the intersection rules are such that there are no allowed configurations for which the entropy diverges, i.e. it is impossible to achieve $\mathcal{N} > 2 \frac{\bar{D}-2}{\bar{D}-3}$.

Let us now digress on a non-orthodox case. We did not consider up to now the branes of the non-maximally supersymmetric string theories, namely type I string theory and the two heterotic string theories. All of these theories include a Super Yang-Mills sector

with a gauge group of rank 16. We consider here only the fields relative to the abelian part of the gauge group.

Consider one of the heterotic theories. The coupling of the 2-form field strengths to the dilaton is $a_2 = -\frac{1}{2}$. There are electrically charged 0-branes and magnetically charged 6-branes, which we call respectively H0 and H6. We now encounter the first difference with respect to the maximally supersymmetric theories: if we compute the Δ for these branes, according to (3.45), we find $\Delta = 1 \cdot 7 + \frac{1}{2} \cdot \frac{1}{4} \cdot 8 = 8$. Remember that previously we always had $\Delta = 16$ in $D = 10$, and that moreover this last value holds for the F1 and NS5 which are also present in the heterotic theories.

The intersections that one finds are the following:

$$H0 \cap H0 = -1, \quad H0 \cap H6 = 0, \quad H6 \cap H6 = 5, \quad (3.171)$$

$$H0 \cap F1 = -1, \quad H0 \cap NS5 = 0, \quad (3.172)$$

$$H6 \cap F1 = 1, \quad H6 \cap NS5 = 4. \quad (3.173)$$

Going to type I theory, we have that the coupling of the 2-forms is given by $a_2 = \frac{1}{2}$. We have thus exactly the same pattern of intersections, where however the NSNS branes are replaced by the D1 and the D5. This is consistent with Heterotic/type I duality.

These 0- and 6-branes of the 10 dimensional $N = 1$ theories are seemingly different from the ones encountered in the maximal supergravities. Indeed, a quick look at the $N = 1$ superalgebra in $D = 10$ [42] directly tells us that there is no central charge corresponding to their (electric or magnetic) charge. This implies that the above configurations cannot be supersymmetric. Their treatability and their potential rôle in the wider context of the underlying string theories is thus uncertain, and probably limited. Note that the H0 and H6 branes, and their intersections, provide examples of extremal branes which are nevertheless non-supersymmetric.

Our last comment is about what we did not do. The derivation of the intersecting brane configurations is straightforwardly generalizable to the case where there are several scalars instead of only one, the dilaton. The general solution and the intersection rules would have come out in quite the same form, with now the coupling to the scalars being a vector instead of a single value. The interest of doing this, besides finding solutions in a more general theory, would have been that now some solutions could have been extrapolated to 10 or 11 dimensional solutions which included waves and/or KK monopoles. Also some intersection rules for these two objects would have been derived. However, the relation between the 10 dimensional couplings and the ones in the necessarily reduced space-time grow more and more complicated as one reduces further dimensions. At a certain moment, it becomes more convenient to apply the T-duality rules as described in [38]. The approach using several scalars has already been followed in [107, 82, 111]. Here, we have preferred to give the derivation in a simpler, even if incomplete, case.

3.3.4 Supersymmetry and the branes

We now consider in more detail one of the most interesting properties of the branes and of their intersections, that is supersymmetry. We have already elaborated on the

power of supersymmetry in extrapolating low-energy, classical results to the quantum underlying theory. We had also anticipated that all the extremal branes presented in subsection 3.2.5 were supersymmetric, i.e. that they preserve one half of the space-time supersymmetries of the theory in which they live.

An interesting problem is the following. The intersecting brane solutions presented in this section combine several of these half-supersymmetric branes. It is thus natural to ask whether the intersecting configuration is supersymmetric and how many supersymmetries it preserves. Here also, we had anticipated that the solutions are indeed supersymmetric.

In this subsection, we will consider the supersymmetric properties of the branes of 11 dimensional supergravity. We will first prove that the basic branes, the M2 and the M5-brane, preserve half of the supersymmetries, and then we will consider the most general configuration of intersecting branes. The intersection rules (3.148) will, as expected, play a crucial rôle in establishing the existence and the amount of preserved supersymmetries.

We focus on 11 dimensional supergravity for two reasons. One is that it is a simpler theory with respect to 10 dimensional type II theories (it has only one antisymmetric tensor field and no dilaton). The second reason is that all the branes, and the intersecting configurations, are related by dualities. Since the dualities preserve the supersymmetry, proving the supersymmetry for a configuration of M-branes proves it also for the other 10 dimensional configurations. Of course this is the drawback of supersymmetry, one can no longer work in a general setting as with (3.9).

What we will actually do is to start with a fairly general configuration of fields, and then impose that some supersymmetries are preserved. We will get in the end the same solution as we found solving the equations of motion. However in the present approach the ansätze will come out of the equations rather than being imposed. If we were only interested in solutions in 11 dimensional supergravity, we could have indeed chosen this approach from the beginning. The advantage of the ‘bosonic’ approach is nevertheless that we have a much wider range of application of our results, and a very concise formula for the intersection rules.

Let us begin writing the supersymmetric variation of the gravitino ψ_μ in 11 dimensional supergravity [43] (in a vanishing gravitino background):

$$\delta\psi_\mu = D_\mu\eta + \frac{1}{288}(\Gamma_{\mu\nu\rho\sigma\lambda} - 8g_{\mu\nu}\Gamma_{\rho\sigma\lambda})F^{\nu\rho\sigma\lambda}\eta, \quad (3.174)$$

$$D_\mu\eta = \partial_\mu\eta + \frac{1}{2}\omega_\mu^{\hat{\rho}\hat{\sigma}}\Sigma_{\hat{\rho}\hat{\sigma}}\eta. \quad (3.175)$$

The 32-component Majorana spinor η is the supersymmetry parameter. The number of its independent arbitrary components measures the portion of preserved supersymmetry. As in Appendix C, the hatted indices are (flat) Lorentz indices while the unhatted ones are covariant indices. A Γ matrix with a covariant index is defined by:

$$\Gamma_\mu = e_\mu^{\hat{\sigma}}\Gamma_{\hat{\sigma}}, \quad \text{where} \quad \Gamma_{\hat{\mu}}\Gamma_{\hat{\nu}} + \Gamma_{\hat{\nu}}\Gamma_{\hat{\mu}} = 2\eta_{\hat{\mu}\hat{\nu}}. \quad (3.176)$$

Γ matrices with several indices are antisymmetric products of simple Γ matrices:

$$\Gamma_{\mu_1\dots\mu_n} = \Gamma_{[\mu_1}\dots\Gamma_{\mu_n]}.$$

The spin connection $\omega_{\mu}^{\hat{\rho}\hat{\sigma}}$ is discussed also in Appendix C, and the matrices:

$$\Sigma_{\hat{\mu}\hat{\nu}} = \frac{1}{4}(\Gamma_{\hat{\mu}}\Gamma_{\hat{\nu}} - \Gamma_{\hat{\nu}}\Gamma_{\hat{\mu}}) = \frac{1}{2}\Gamma_{\hat{\mu}\hat{\nu}}$$

are the $SO(1, 10)$ Lorentz generators acting on the spinor representation.

We do not need to write the supersymmetric variations of the bosonic fields because, since the fermionic fields (the gravitino) have vanishing expectation value in the classical background we are considering, they are trivially zero. Asking that $\delta\psi_{\mu} = 0$ will be sufficient to prove that the solution is unchanged with respect to that particular supersymmetry transformation.

Consider first the M2-brane. The metric and 4-form field strength we start with are:

$$ds^2 = -B^2 dt^2 + \sum_{i=1}^2 C^2 dy_i^2 + G^2 \delta_{ab} dx^a dx^b, \quad a, b = 1 \dots 8,$$

$$F_{tija} = \epsilon_{ij} \partial_a E.$$

The spin connection relative to this geometry is given in Appendix C, eqs. (C.17)–(C.19).

Supposing that the supersymmetry parameter also depends only on the x 's, we obtain after some algebra (the index 0 stands for t):

$$\delta\psi_t = \frac{1}{6C^2G} \Gamma_{\hat{0}} (3C^2 \partial_a B \Gamma_{\hat{a}} - \partial_a E \Gamma_{\hat{a}} \Gamma_{\hat{0}\hat{1}\hat{2}}) \eta, \quad (3.177)$$

$$\delta\psi_i = \frac{1}{6BCG} \Gamma_{\hat{i}} (3BC \partial_a C \Gamma_{\hat{a}} - \partial_a E \Gamma_{\hat{a}} \Gamma_{\hat{0}\hat{1}\hat{2}}) \eta, \quad (3.178)$$

$$\delta\psi_a = \partial_a \eta + \frac{1}{2G} \partial_b G \Gamma_{\hat{a}\hat{b}} \eta + \frac{1}{12BC^2} \partial_b E \Gamma_{\hat{a}\hat{b}} \Gamma_{\hat{0}\hat{1}\hat{2}} \eta - \frac{1}{6BC^2} \partial_b E \Gamma_{\hat{0}\hat{1}\hat{2}} \eta. \quad (3.179)$$

Begin with imposing $\delta\psi_t = 0$. It implies:

$$3C^2 \partial_a B = \pm \partial_a E, \quad (1 \mp \Gamma_{\hat{0}\hat{1}\hat{2}}) \eta = 0. \quad (3.180)$$

This can be seen as follows. If $M = \not{u} + \not{v}\tilde{\Gamma}$, where $\not{u} = u_a \Gamma_a$ (we drop hats) and $\tilde{\Gamma}$ is such that $\tilde{\Gamma}^2 = 1$, $tr \tilde{\Gamma} = 0$, $\tilde{\Gamma} \Gamma_a = -\Gamma_a \tilde{\Gamma}$ and $\tilde{\Gamma}^T = \tilde{\Gamma}$, as $\Gamma_{\hat{0}\hat{1}\hat{2}}$ (for definiteness, we are in a basis in which all the Γ 's are real), then we have:

$$\begin{aligned} \det M &= (\det M^T \det M)^{\frac{1}{2}} = (\det M^T M)^{\frac{1}{2}} = \left[\det(u^2 + v^2 + 2u \cdot v \tilde{\Gamma}) \right]^{\frac{1}{2}} \\ &= (u^2 + v^2 + 2u \cdot v)^8 (u^2 + v^2 - 2u \cdot v)^8 = |u + v|^{16} |u - v|^{16}. \end{aligned}$$

The condition $M\eta = 0$ has non-trivial solutions if and only if $\det M = 0$, and this in turn implies $u_a = \pm v_a$. The spinor η is then solution of $(1 \pm \tilde{\Gamma})\eta = 0$, which projects it onto a subspace of 16 (real) dimensions over 32.

The equation $\delta\psi_i = 0$ then becomes, making use of (3.180):

$$\frac{C}{2G} \Gamma_{\hat{i}} (\partial_a \ln C - \partial_a \ln B) \Gamma_{\hat{a}} \eta = 0,$$

which thus gives:

$$B = C \quad \Rightarrow \quad E = \pm B^3. \quad (3.181)$$

We thus see that $SO(1, 2)$ Lorentz invariance on the world-volume of the M2-brane is imposed on us by supersymmetry. Note that if we wanted an asymptotically vanishing potential E , we should write, instead of the second relation above, $E = \pm(B^3 - 1)$; the conclusions are obviously unchanged.

Asking that $\delta\psi_a = 0$, and using (3.180) and (3.181), we arrive at the expression:

$$\partial_a \eta - \frac{1}{2} \partial_a \ln B \eta + \frac{1}{2} \left(\partial_b \ln G + \frac{1}{2} \partial_b \ln B \right) \Gamma_{\hat{a}\hat{b}} \eta = 0.$$

This equation is solved⁴ for:

$$BG^2 = 1, \quad \eta = B^{\frac{1}{2}} \eta_{(0)}, \quad (3.182)$$

where $\eta_{(0)}$ is a constant spinor. Note that the first relation above can be rewritten as $B^3 G^6 = 1$, i.e. exactly in the same form as for the ‘extremality’ condition (3.40).

Using the equations of motion for the 4-form field strength one can now with a minimum effort solve completely the problem, finding that $\partial_a \partial_a E^{-1} = 0$. If one writes the solution in terms of the harmonic function $H = E^{-1}$, then the solution (3.102) is found once again.

We have thus proved that the solution (3.102) is preserved by the supersymmetries with parameter η such that:

$$\eta = H^{-\frac{1}{6}} \eta_{(0)}, \quad \eta_{(0)} = \frac{1 + \Gamma_{\hat{0}\hat{1}\hat{2}}}{2} \eta_{(0)}. \quad (3.183)$$

We thus say that the M2-brane preserves half of the supersymmetries.

Let us now consider the M5-brane. Its metric and field strength are taken to be generally:

$$ds^2 = -B^2 dt^2 + \sum_{i=1}^5 C^2 dy_i^2 + G^2 \delta_{ab} dx^a dx^b, \quad a, b = 1 \dots 5,$$

$$F^{abcd} = \frac{1}{BC^5 G^5} \epsilon^{abcde} \partial_e \tilde{E}.$$

The variation of the gravitino is then the following:

$$\delta\psi_t = \frac{1}{12C^5 G} \Gamma_{\hat{0}} (6C^5 \partial_a B \Gamma_{\hat{a}} + \partial_a \tilde{E} \Gamma_{\hat{a}} \Gamma_{(5)}) \eta, \quad (3.184)$$

$$\delta\psi_i = \frac{1}{12BC^4 G} \Gamma_{\hat{i}} (6BC^4 \partial_a C \Gamma_{\hat{a}} + \partial_a \tilde{E} \Gamma_{\hat{a}} \Gamma_{(5)}) \eta, \quad (3.185)$$

$$\delta\psi_a = \partial_a \eta + \frac{1}{2G} \partial_b G \Gamma_{\hat{a}\hat{b}} \eta + \frac{1}{12BC^5} \partial_a \tilde{E} \Gamma_{(5)} \eta - \frac{1}{6BC^5} \partial_b \tilde{E} \Gamma_{\hat{a}\hat{b}} \Gamma_{(5)} \eta. \quad (3.186)$$

⁴That the solution (3.182) is the only one can be proved [94] using arguments relative to the integrability of the supersymmetric condition $\delta\psi_\mu = 0$, see for instance [112].

We have used the following properties of the Γ matrices (all indices hereafter are flat): suppose $\Gamma_{(d)} = \frac{1}{d!} \epsilon^{a_1 \dots a_d} \Gamma_{a_1 \dots a_d} \equiv \Gamma_1 \dots \Gamma_d$; we then have:

$$\epsilon^{ab_1 \dots b_{d-1}} \Gamma_{b_1 \dots b_{d-1}} = (d-1)! \Gamma^a \Gamma_{(d)}, \quad \epsilon^{abc_1 \dots c_{d-2}} \Gamma_{c_1 \dots c_{d-2}} = -(d-2)! \Gamma^{ab} \Gamma_{(d)}. \quad (3.187)$$

The condition $\delta\psi_t = 0$ implies in a way exactly similar as for the M2-brane (except that we now have $\Gamma_{(5)} \Gamma_a = \Gamma_a \Gamma_{(5)}$):

$$6C^5 \partial_a B = \pm \partial_a \tilde{E}, \quad (1 \pm \Gamma_{(5)}) \eta = 0. \quad (3.188)$$

Then $\delta\psi_i = 0$ in turn gives:

$$B = C \quad \Rightarrow \quad \tilde{E} = \pm B^6. \quad (3.189)$$

The last condition is $\delta\psi_a = 0$, and it is solved by:

$$B^2 G = 1, \quad \eta = B^{\frac{1}{2}} \eta_{(0)}. \quad (3.190)$$

Again, the first relation above can be seen to be equivalent to the extremality condition (3.40) for a 5-brane, $B^6 G^3 = 1$.

The Bianchi identities for the 4-form impose straightforwardly that $\tilde{E}^{-1} \equiv H$ is harmonic. We have thus rediscovered the M5-brane solution as in (3.103). The M5-brane is therefore invariant under a supersymmetry variation of parameter:

$$\eta = H^{-\frac{1}{12}} \eta_{(0)}, \quad \eta_{(0)} = \frac{1 - \Gamma_{(5)}}{2} \eta_{(0)}. \quad (3.191)$$

Note that the sign in front of $\Gamma_{(5)}$ is a consequence of the conventions. Using the properties of the Γ matrices in 11 dimensions, most notably the fact that $\Gamma_0 \dots \Gamma_{10} = 1$, the condition on the constant spinor can be rewritten in the more familiar form:

$$\Gamma_{\hat{0}\hat{1}\hat{2}\hat{3}\hat{4}\hat{5}} \eta_{(0)} = \pm \eta_{(0)},$$

where all the indices now lie along the world-volume of the M5-brane.

We are now ready to go to the general case, where we have an arbitrary number of (intersecting) M2- and M5-branes. Accordingly, we start with a general metric as (3.117), and with a set of non-trivial electric and magnetic components of the 4-form field strength:

$$F_{ti_1 i_2 a} = \epsilon_{i_1 i_2} \partial_a E_N, \quad N = 1 \dots \mathcal{N}_{el}, \quad (3.192)$$

$$F^{i_6 \dots i_p a_1 \dots a_{d-1}} = \frac{1}{BC_1 \dots C_p G^d} \epsilon^{i_6 \dots i_p} \epsilon^{a_1 \dots a_d} \partial_{a_d} \tilde{E}_M, \quad M = 1 \dots \mathcal{N}_{mag}. \quad (3.193)$$

Note that using (3.120), we can rewrite $BC_1 \dots C_p = V_M V_M^\perp$ and $F^{ti_1 i_2 a} = -\frac{1}{V_N^2 G^2} \epsilon_{i_1 i_2} \partial_a E_N$.

It is useful to keep in mind the previous computations for a single M2-brane and a single M5-brane to compute the gravitino variation in this general case. We obtain:

$$\delta\psi_t = \frac{1}{2G} \partial_a B \Gamma_0 \Gamma_{\hat{a}} \eta - \frac{1}{6} \sum_N \frac{B}{V_N G} \partial_a E_N \Gamma_0 \Gamma_{\hat{a}} \Gamma_N \eta + \frac{1}{12} \sum_M \frac{B}{V_M G} \partial_a \tilde{E}_M \Gamma_0 \Gamma_{\hat{a}} \Gamma_M^\perp \eta,$$

$$\begin{aligned}
\delta\psi_i &= \frac{1}{2G}\partial_a C_i \Gamma_i \Gamma_{\hat{a}} \eta - \frac{1}{6} \sum_{N \parallel i} \frac{C_i}{V_N G} \partial_a E_N \Gamma_i \Gamma_{\hat{a}} \Gamma_N \eta + \frac{1}{12} \sum_{N \perp i} \frac{C_i}{V_N G} \partial_a E_N \Gamma_i \Gamma_{\hat{a}} \Gamma_N \eta \\
&\quad + \frac{1}{12} \sum_{M \parallel i} \frac{C_i}{V_M G} \partial_a \tilde{E}_M \Gamma_i \Gamma_{\hat{a}} \Gamma_M^\perp \eta - \frac{1}{6} \sum_{M \perp i} \frac{C_i}{V_M G} \partial_a \tilde{E}_M \Gamma_i \Gamma_{\hat{a}} \Gamma_M^\perp \eta, \\
\delta\psi_a &= \partial_a \eta + \frac{1}{2G} \partial_b G \Gamma_{\hat{a} \hat{b}} \eta + \frac{1}{12} \sum_N \frac{1}{V_N} \partial_b E_N \Gamma_{\hat{a} \hat{b}} \Gamma_N \eta - \frac{1}{6} \sum_N \frac{1}{V_N} \partial_a E_N \Gamma_N \eta \\
&\quad + \frac{1}{12} \sum_M \frac{1}{V_M} \partial_a \tilde{E}_M \Gamma_M^\perp \eta - \frac{1}{6} \sum_M \frac{1}{V_M} \partial_b \tilde{E}_M \Gamma_{\hat{a} \hat{b}} \Gamma_M^\perp \eta. \tag{3.194}
\end{aligned}$$

We have written Γ_N for the product of Γ matrices longitudinal to the M2-brane labelled by N , and Γ_M^\perp for the product of Γ matrices transverse to the M5-brane labelled by M . The notation $A \parallel i$ ($A \perp i$) denotes that the brane is longitudinal (transverse) to the direction y_i .

The set of conditions $\delta\psi_t = 0$ and $\delta\psi_i = 0$ can be combined in the $\mathcal{N}_{el} + \mathcal{N}_{mag}$ following equations:

$$\begin{aligned}
\partial_a \ln V_N \Gamma_{\hat{a}} \eta - \frac{1}{V_N} \partial_a E_N \Gamma_{\hat{a}} \Gamma_N \eta - \frac{1}{2} \sum_{N' \neq N} \bar{q}_{NN'} \frac{1}{V_{N'}} \partial_a E_{N'} \Gamma_{\hat{a}} \Gamma_{N'} \eta \\
\quad + \frac{1}{2} \sum_M (\bar{q}_{NM} - 1) \frac{1}{V_M} \partial_a \tilde{E}_M \Gamma_{\hat{a}} \Gamma_M^\perp \eta = 0, \\
\partial_a \ln V_M \Gamma_{\hat{a}} \eta - \frac{1}{V_M} \partial_a \tilde{E}_M \Gamma_{\hat{a}} \Gamma_M^\perp \eta - \frac{1}{2} \sum_N (\bar{q}_{NM} - 1) \frac{1}{V_N} \partial_a E_N \Gamma_{\hat{a}} \Gamma_N \eta \\
\quad + \frac{1}{2} \sum_{M' \neq M} (\bar{q}_{MM'} - 3) \frac{1}{V_{M'}} \partial_a \tilde{E}_{M'} \Gamma_{\hat{a}} \Gamma_{M'}^\perp \eta = 0.
\end{aligned}$$

As before, \bar{q}_{AB} is the dimension of the intersection of the two branes in its label. The equations above are easily solved as follows. We have to take $\bar{q}_{NN'} = 0$, $\bar{q}_{NM} = 1$ and $\bar{q}_{MM'} = 3$, i.e. exactly the intersection rules (3.158). Note however that we do not recover the general formula (3.157). The equations are then completely solved by:

$$\begin{aligned}
V_N &= \pm E_N, & (1 \mp \Gamma_N) \eta &= 0, \\
V_M &= \pm \tilde{E}_M, & (1 \pm \Gamma_M^\perp) \eta &= 0. \tag{3.195}
\end{aligned}$$

We will comment shortly on the compatibility of the several projections on the supersymmetry parameter η .

The relations above however do not completely solve the conditions $\delta\psi_t = 0$ and $\delta\psi_i = 0$, since typically we have $p+1 \geq \mathcal{N}_{el} + \mathcal{N}_{mag}$ (taking into account the intersection rules). Using (3.195), they further impose:

$$\frac{B}{2G} \Gamma_{\hat{0}} \left[\partial_a \ln B - \frac{1}{3} \sum_N \partial_a \ln E_N - \frac{1}{6} \sum_M \partial_a \ln \tilde{E}_M \right] \Gamma_{\hat{a}} \eta = 0$$

$$\begin{aligned} \frac{C_i}{2G} \Gamma_{\hat{i}} \left[\partial_a \ln C_i - \frac{1}{3} \sum_{N \parallel i} \partial_a \ln E_N + \frac{1}{6} \sum_{N \perp i} \partial_a \ln E_N \right. \\ \left. - \frac{1}{6} \sum_{M \parallel i} \partial_a \ln \tilde{E}_M + \frac{1}{3} \sum_{M \perp i} \partial_a \ln \tilde{E}_M \right] \Gamma_{\hat{a}} \eta = 0 \end{aligned}$$

The solution to this equations gives the expression of B and C_i in terms of the E_N and \tilde{E}_M :

$$B = \prod_N E_N^{\frac{1}{3}} \prod_M \tilde{E}_M^{\frac{1}{6}}, \quad C_i = \prod_{N \parallel i} E_N^{\frac{1}{3}} \prod_{N \perp i} E_N^{-\frac{1}{6}} \prod_{M \parallel i} \tilde{E}_M^{\frac{1}{6}} \prod_{M \perp i} \tilde{E}_M^{-\frac{1}{3}}. \quad (3.196)$$

In order to determine the expression for G and for the preserved spinor η , we solve the last set of equations $\delta\psi_a = 0$, which gives:

$$\begin{aligned} \partial_a \eta - \frac{1}{6} \sum_N \partial_a \ln E_N \eta - \frac{1}{12} \sum_M \partial_a \ln \tilde{E}_M \eta \\ + \frac{1}{2} \left(\partial_b \ln G + \frac{1}{6} \sum_N \partial_b \ln E_N + \frac{1}{3} \sum_M \partial_b \ln \tilde{E}_M \right) \Gamma_{\hat{a}\hat{b}} \eta = 0. \end{aligned}$$

The solution is thus:

$$G = \prod_N E_N^{-\frac{1}{6}} \prod_M \tilde{E}_M^{-\frac{1}{3}}, \quad \eta = \prod_N E_N^{\frac{1}{6}} \prod_M \tilde{E}_M^{\frac{1}{12}} \eta_{(0)}, \quad (3.197)$$

where $\eta_{(0)}$ is the usual constant spinor. It can be trivially checked that the functions above verify the extremality ansatz $BC_1 \dots C_p G^{d-2} = 1$. The equations of motion and Bianchi identities for the 4-form field strength are now simple to solve and they yield that $H_N = E_N^{-1}$ and $H_M = \tilde{E}_M^{-1}$ are harmonic functions.

In terms of these harmonic functions, we have the same solution as in (3.145)–(3.146) applied to the 11 dimensional supergravity. We have thus shown above that this solution preserves the supersymmetries with parameter constrained by:

$$\eta = \prod_N H_N^{-\frac{1}{6}} \prod_M H_M^{-\frac{1}{12}} \eta_{(0)}, \quad \Gamma_N \eta_{(0)} = \eta_{(0)}, \quad \Gamma_M \eta_{(0)} = -\eta_{(0)}. \quad (3.198)$$

We might be worried that the projections on the constant spinor $\eta_{(0)}$ eventually project it onto a zero-dimensional subspace. This would be the case if there were pairs of Γ_A matrices ($A = N$ or M) which anticommuted. It would have the consequence that:

$$\eta_{(0)} = \Gamma_A \eta_{(0)} = \Gamma_A \Gamma_B \eta_{(0)} = -\Gamma_B \Gamma_A \eta_{(0)} = -\eta_{(0)} \equiv 0.$$

Happily, and owing to the intersection rules (3.158), all such pairs of matrices commute. If there are \mathcal{N} branes in the configuration, then $\eta_{(0)}$ is projected by:

$$\mathcal{P}_{\mathcal{N}} \eta_{(0)} = \frac{1 + \Gamma_{[1]}}{2} \frac{1 + \Gamma_{[2]}}{2} \dots \frac{1 + \Gamma_{[\mathcal{N}]}}{2} \eta_{(0)} = \eta_{(0)}. \quad (3.199)$$

The dimension of the subspace onto which $\eta_{(0)}$ is projected can be computed performing the trace of the projector, which is always:

$$\text{tr}\mathcal{P}_{\mathcal{N}} \geq \frac{1}{2^{\mathcal{N}}}(32). \quad (3.200)$$

The equality is obtained whenever there are no products of the Γ_A matrices appearing in (3.199) that give the unity. If on the other hand we have, for instance, $\Gamma_{[1]}\Gamma_{[2]}\dots\Gamma_{[\mathcal{N}-1]} = \Gamma_{[\mathcal{N}]}$, then it is easy to show that we have $\mathcal{P}_{\mathcal{N}} = \mathcal{P}_{\mathcal{N}-1}$.

We end this subsection giving two examples, one which saturates the bound (3.200) and the other which does not. The first one is the configuration already displayed in (3.169), consisting of three sets of M2-branes lying respectively in the directions $\hat{1}\hat{2}$, $\hat{3}\hat{4}$ and $\hat{5}\hat{6}$. This configuration can be trivially seen to preserve $\frac{1}{8}$ of the supersymmetries.

The second configuration is the following: take two sets of M5-branes lying in the directions $\hat{1}\hat{2}\hat{3}\hat{4}\hat{5}$ and $\hat{1}\hat{2}\hat{3}\hat{6}\hat{7}$, and two sets of M2-branes along $\hat{4}\hat{6}$ and $\hat{5}\hat{7}$. Note that the intersection rules are respected. We can write the following table:

	0	1	2	3	4	5	6	7	8	9	10
M5	×	×	×	×	×	×	—	—	—	—	—
M5	×	×	×	×	—	—	×	×	—	—	—
M2	×	—	—	—	×	—	×	—	—	—	—
M2	×	—	—	—	—	×	—	×	—	—	—

It is now easy to see that, if we define $\bar{\Gamma} = \Gamma_8\Gamma_9\Gamma_{10}$, we have:

$$(\Gamma_6\Gamma_7\bar{\Gamma})(\Gamma_4\Gamma_5\bar{\Gamma})(\Gamma_0\Gamma_4\Gamma_6) = -\Gamma_0\Gamma_5\Gamma_7.$$

Note that the sign of the fourth projector is however determined by the signs of the three other ones. This means that if we had only the first three sets of branes, we could add the fourth one without breaking additional supersymmetry, but one has to be careful not to add anti-branes in the last step otherwise the supersymmetry is completely broken (note that choosing which are the branes and which the anti-branes is a matter of convention). The conclusion is thus that the above configuration, which is dual to the IIA configuration shown in (3.170), preserves also $\frac{1}{8}$ supersymmetries.

This, and the dualities, imply that the 5 dimensional extremal black holes with 3 charges, and the 4 dimensional ones with 4 charges, the same that have a non-vanishing entropy, all are supersymmetric and preserve $\frac{1}{8}$ of the 32 supersymmetries. This is crucial in establishing the relation between the semi-classical entropy and the microscopic counting of states in the D-brane picture.

3.3.5 A non-supersymmetric, extremal solution

We conclude this section on intersecting branes with an odd case. The solution that we present in this subsection shows that the penalty for violating the intersection rules is the loss of supersymmetry.

Keeping the things simple we only consider a solution carrying two charges, thus a bound state of two (sets of) branes. We also still focus on extremal solutions, though this case could be easily extended to a non-extremal version.

A naive way to look for a solution is to try to formulate the problem in the most simple way. Most notably, what complicates the solutions from the point of view of the reduced $(D-p)$ -dimensional space-time, is the presence of non-trivial moduli fields, such as the dilaton and the metric in the compact space longitudinal to the branes. We can thus search for solutions with trivial moduli, i.e. for which ϕ and C_i are constant and thus keep the value that is imposed at infinity throughout all the space.

In order to do this, we can take advantage of the structure of the r.h.s. of the equations (3.123)–(3.127), in particular the following ones:

$$\begin{aligned} R_i^i &= -\frac{1}{2} \frac{\delta_1^i}{D-2} (V_1 G)^{-2} e^{\varepsilon_1 a_1 \phi} (\partial E_1)^2 - \frac{1}{2} \frac{\delta_2^i}{D-2} (V_2 G)^{-2} e^{\varepsilon_2 a_2 \phi} (\partial E_2)^2, \\ \square \phi &= -\frac{1}{2} \varepsilon_1 a_1 (V_1 G)^{-2} e^{\varepsilon_1 a_1 \phi} (\partial E_1)^2 - \frac{1}{2} \varepsilon_2 a_2 (V_2 G)^{-2} e^{\varepsilon_2 a_2 \phi} (\partial E_2)^2. \end{aligned}$$

The labels 1 and 2 refer to the two branes present in the problem. Both of the r.h.s. above are sums of terms with coefficients which can have an arbitrary sign. If we now implement in the equations that the moduli are constant, i.e. we take $C_i = 1$ and $\phi = 0$, the l.h.s. become trivially vanishing. The equations reduce to:

$$\frac{1}{2(D-2)(BG)^2} [-\delta_1^i (\partial E_1)^2 - \delta_2^i (\partial E_2)^2] = 0, \quad (3.201)$$

$$\frac{1}{2(BG)^2} [-\varepsilon_1 a_1 (\partial E_1)^2 - \varepsilon_2 a_2 (\partial E_2)^2] = 0. \quad (3.202)$$

At a first glance, we see that for the equations (3.201) to be satisfied, δ_1^i and δ_2^i must be of opposite sign for each direction y_i . This means, if we recall the definition (3.122), that each direction has to be longitudinal to one brane and transverse to the other. In other words, the two branes intersect in a completely orthogonal way, $q_1 \cap q_2 = 0$. If for the moment we assume that both branes are extended objects ($q > 0$), the equations (3.201) imply two relations:

$$(D - q_1 - 3)(\partial E_1)^2 - (q_2 + 1)(\partial E_2)^2 = 0, \quad (3.203)$$

$$(q_1 + 1)(\partial E_1)^2 - (D - q_2 - 3)(\partial E_2)^2 = 0. \quad (3.204)$$

The sum of these two equations gives the following highly constraining equation:

$$(D - 2)[(\partial E_1)^2 - (\partial E_2)^2] = 0.$$

We thus take both functions to be equal:

$$E_1 = E_2 \equiv E. \quad (3.205)$$

This physically means that both branes will have the same total amount of charge. Note that since the functions coincide, in particular the centers will coincide also.

Plugging (3.205) in (3.203) or (3.204), we have the following constraint on the extension of the branes:

$$q_1 + q_2 = D - 4. \quad (3.206)$$

They have thus the right dimensions to be an electric-magnetic dual pair of branes. This is confirmed reinspecting the equation (3.202), which forces to have $\varepsilon_1 a_1 = -\varepsilon_2 a_2$. Since conventionally (3.206) implies that ε_1 and ε_2 will be of opposite sign, we must have $a_1 = a_2$, i.e. the two branes are really an electric-magnetic dual pair.

This conclusion is unchanged if one of the branes is a 0-brane (say $q_1 = 0$), since then its dual is a $(D - 4)$ -brane, and the equation (3.204), which is the only one, is still satisfied.

Note that in the case $q_1 = 0$, we could have more general cases, not necessarily leading to $a_1 = a_2$. However none of these cases applies directly to string/M-theory.

We can now solve completely the problem, writing the remaining equations:

$$\begin{aligned} R_t^t &= -\frac{1}{2}(BG)^{-2}(\partial E)^2, \\ R_b^a &= (BG)^{-2} \left\{ -\partial_a E \partial_b E + \frac{1}{2} \delta_b^a (\partial E)^2 \right\}, \\ \partial_a \left(\frac{G}{B} \partial_a E \right) &= 0. \end{aligned}$$

Note that necessarily, (3.206) implies that we have $d = 3$. Accordingly, the extremality condition (3.128) becomes $BG = 1$. The equations above become:

$$\begin{aligned} \partial_a \partial_a \ln B &= \frac{1}{2B^2} (\partial E)^2, \\ 2\partial_a \ln B \partial_b \ln B &= \frac{1}{B^2} \partial_a E \partial_b E, \\ \partial_a \left(\frac{1}{B^2} \partial_a E \right) &= 0. \end{aligned} \quad (3.207)$$

The solution is easily found as for (3.145)–(3.146), and it yields:

$$E = \sqrt{2}H^{-1}, \quad B = H^{-1}, \quad G = H, \quad (3.208)$$

where H is the harmonic function:

$$H = 1 + \sum_k \frac{Q_k}{|x - x_k|}. \quad (3.209)$$

As already noted, the fact that we obtain a multi-centered solution should not be misunderstood. Our construction forbids us to separate the q_1 -branes from the q_2 -branes in each bound state. However there is a vanishing force between two (or more) such bound states.

The metric thus writes simply:

$$ds^2 = -H^{-2} dt^2 + \sum_{i=1}^{q_1+q_2} dy_i^2 + H^2 dx_a dx_a. \quad (3.210)$$

In the simplest case in which we have only one center, $H = 1 + \frac{Q}{r}$, we can perform the change of coordinates $R = r + Q$, moving the horizon to the value $R = Q$. The metric becomes:

$$ds^2 = - \left(1 - \frac{Q}{R}\right)^2 dt^2 + \sum_{i=1}^{q_1+q_2} dy_i^2 + \left(1 - \frac{Q}{R}\right)^{-2} dR^2 + R^2 d\Omega_2^2. \quad (3.211)$$

This is exactly the metric of a 4 dimensional extreme Reissner-Nordström black hole, multiplied by a flat p -dimensional internal space. The generalization to the non-extreme Reissner-Nordström black hole is straightforward, but the extreme version is enough for the present discussion.

Let us now compute the mass-charge relation of the above solution. To fix the notations, we set ourselves in 10 dimensional type IIA theory, and we consider a bound state of D0- and D6-branes. Remember that these two branes were unable to form a bound state with vanishing binding energy.

The RR 2-form has the following two non-trivial components:

$$F_{ta} = \sqrt{2}\partial_a H^{-1}, \quad F_{ab} = -\sqrt{2}\epsilon_{abc}\partial_c H.$$

Note that with respect to the solution (3.146) applied to the D0- and D6-brane, we have an extra factor of $\sqrt{2}$.

The charges carried by the solution (3.208) are thus:

$$\mathcal{Q}_0 = \frac{1}{16\pi G_{10}} \int_{T^6 \times S^2} *F_2 = \sqrt{2} \frac{L^6}{4G_{10}} Q, \quad q_0 = \mathcal{Q}_0, \quad (3.212)$$

$$\mathcal{Q}_6 = \frac{1}{16\pi G_{10}} \int_{S^2} F_2 = \sqrt{2} \frac{1}{4G_{10}} Q, \quad q_6 = L^6 \mathcal{Q}_6. \quad (3.213)$$

As expected, the solution carries the same amount of charges, $q_0 = q_6$. This does not necessarily mean that, if we consider the elementary branes with quantized charge, we should have an equal number of both kinds. Rather, the ratio will be fixed by the volume of the T^6 in the relevant units.

The ADM mass is easily obtained computing the formula (3.82) for the metric (3.210):

$$M = \frac{L^6}{G_{10}} |Q|.$$

We have now to recall that in section 3.2.5 we established that for all the branes of M-theory, we had that $M_p = |q_p|$. The expression for the ADM mass can thus be rewritten as:

$$M = 2\sqrt{2}|q_0| = \sqrt{2}(|q_0| + |q_6|) \equiv \sqrt{2}(M_0 + M_6). \quad (3.214)$$

This result is interesting. The energy of the bound state is bigger than the energy of the two constituent objects. This should be ascribed to the fact that from the perturbative string theory point of view, D0- and D6-branes exert a repulsive force on each other [113]. One has then to supply extra energy to bind them. We thus do not seem to be in presence of a stable bound state. At best, it could be a metastable one. Some

of the properties of this bound state were analyzed in [114] (see also [115], and [116] where the D-brane description of the bound state is given, along with evidence for its metastability).

Another feature of this solution is that, as one can easily guess from (3.211), it has a non-vanishing entropy. This also violates the rule for the supersymmetric black holes, which in four dimensions need 4 charges to have a non-vanishing entropy.

The two remarks above are strong hints against the preservation of supersymmetry by the solution (3.208). It is actually very easy to prove it directly.

In order to be able to use the equations of the preceding subsection, we have first to perform some dualities to reformulate the problem in 11 dimensional supergravity. The D0+D6 bound state can be T-dualized to a D1∩D5=0 intersection in IIB theory, then S-dualized to a F1∩NS5=0 configuration, and eventually elevated (after another T-duality) to a M2∩M5=0 bound state in 11 dimensions. Note that the latter configuration is also a solution of the kind described above. The chain of dualities ensures that the two solutions have the same supersymmetry properties.

Let us suppose that the M2-brane is stretched in the $\hat{1}\hat{2}$ directions, while the M5 lies in $\hat{3}\hat{4}\hat{5}\hat{6}\hat{7}$. Then supersymmetry imposes, for instance, $\delta\psi_{y_1} = 0$. Plugging the solution (3.208) in the second of the expressions (3.194), we get the following expression:

$$\begin{aligned}\delta\psi_{y_1} &= -\frac{1}{6BG}\partial_a E \Gamma_{\hat{1}}\Gamma_{\hat{a}}\Gamma_{M_2}\eta - \frac{1}{6BG}\partial_a E \Gamma_{\hat{1}}\Gamma_{\hat{a}}\Gamma_{M_5}^\perp\eta \\ &= -\frac{\sqrt{2}}{6}\partial_a H^{-1}\Gamma_{\hat{1}}\Gamma_{\hat{a}}(\Gamma_{M_2} + \Gamma_{M_5}^\perp)\eta.\end{aligned}$$

Here we have defined $\Gamma_{M_2} = \Gamma_0\Gamma_1\Gamma_2$ and $\Gamma_{M_5}^\perp = \Gamma_1\Gamma_2\bar{\Gamma}$, with $\bar{\Gamma} = \Gamma_8\Gamma_9\Gamma_{10}$ (we have dropped hats).

The condition for preserved supersymmetry would thus read:

$$\tilde{\Gamma}\eta = \eta, \quad \tilde{\Gamma} = -\Gamma_{M_2}\Gamma_{M_5}^\perp. \quad (3.215)$$

However it is easy to see that $\tilde{\Gamma}^2 = -1$. This implies that the condition (3.215) has no non-trivial solutions. Therefore, the bound state (3.208) breaks all supersymmetries.

We have thus presented in this subsection an example of a solution of a different kind of the ones presented in the previous subsections.

3.4 Conclusions and related work

In this last section we recollect the results obtained in this chapter, and we (rapidly) review the recent work on classical p -brane solutions other than the ones considered here.

The main objective of this chapter was to study the basic p -branes that one encounters in M-theory, and to treat them in a unified way. The need to unify the treatment is inspired by U-duality [20, 17, 42], which states that from the effective lower dimensional space-time point of view, all the charges carried by the different branes are on the same footing. While string theory ‘breaks’ this U-duality symmetry, choosing the

NSNS string to be the fundamental object of the perturbative theory, the supergravity low-energy effective theories realize the U-duality at the classical level.

Here we have pushed the ‘unification’ further choosing a general action (3.9) which can be reduced to the bosonic sector of any one of the supergravities of interest in 10 or 11 dimensions. The evidence that it was indeed worthwhile considering this unified picture, is that we could very easily recover all the branes of M-theory in section 3.2.5, and moreover we were able to present the very general intersection rule (3.148). This is an interesting fact since it relates the properties of the various branes directly in 11 and 10 dimensions, while U-duality strictly speaking needs that we carry out compactification.

Supersymmetry is an odd ingredient in the present chapter. Its main disadvantage is that it is impossible to formulate a supersymmetric version of the general theory that we have considered. On the other hand, the supersymmetric properties of the branes are crucial in relating the classical solutions to the quantum objects. Here we have proved, in the context of 11 dimensional supergravity, that for the ‘elementary’ branes, extremality coincides with supersymmetry. Also, the proof that the extremal intersecting brane bound states leading to a non-vanishing semi-classical entropy are indeed supersymmetric is very important in view of their relevance in black hole physics.

One could wonder whether the ansätze made in order to solve the (bosonic) equations of motion (as (3.128) and the reduction to \mathcal{N} independent functions) were actually motivated by supersymmetry. It should be more precise to say that they were motivated by extremality. Indeed, the no-force condition between the branes participating to the configuration is a consequence of extremality, as the same is true for BPS monopoles [117]. As they are, the BPS conditions for monopoles and, here, for branes do not need supersymmetry, but coincide with the preservation of some supersymmetries in a (sufficiently) supersymmetric theory. Supersymmetry thus generically prevents the BPS conditions to be corrected by quantum effects when one goes beyond the classical solutions. A good example for this issue is to consider brane configurations in a theory with less than maximal supersymmetry. There are for instance configurations in type I string theory which are closely related to configurations in type IIB theory. Namely, the D1-brane, the D5-brane, the wave and the KK-monopole are exactly the same in both theories. It has been shown in [118] that a configuration which preserves 1/8 of type II supersymmetry can be transposed to a non-supersymmetric, but still extremal configuration in type I, satisfying the no-force condition. Leaving aside the considerations on the relevance of this bound state in the quantum theory, we see that imposing the no-force condition and extremality does not single out only supersymmetric solutions.

It should also not be underestimated that the derivation of the intersecting solutions presented in this chapter is a thorough consistency check of all the dualities acting on, and between, the supergravity theories. It is straightforward to check that, starting from one definite configuration, all its dual configurations are also found between the solutions presented here (with the exception of the solutions involving waves and KK monopoles). In this line of thoughts, we presented a recipe for building five and four dimensional extreme supersymmetric black holes. Some of these black holes were used in the literature to perform a microscopic counting of their entropy, as in [61, 62] for the 5-dimensional ones. Actually, the only (5 dimensional) black holes in the U-duality

‘orbit’ that were counted were the ones containing only D-branes and KK momentum. It is still an open problem to directly count the microscopic states of the same black hole but in a different M-theoretic formulation.

Some of the intersection rules (3.148) point towards an M-theory interpretation in terms of open branes ending on other branes. This idea will be elaborated and made firmer in the next chapter. It suffices to say here that this interpretation is consistent with dualities if we postulate that the ‘open character’ of a fundamental string ending on a D-brane is invariant under dualities. S-duality directly implies, for instance, that D-strings can end on NS5-branes [99]. Then T-dualities imply that all the D-branes can end on the NS5-brane. In particular, the fact that the D2-brane can end on the NS5-brane should imply that the M5-brane is a D-brane for the M2-branes [99, 74, 100] (this could also be extrapolated from the fact that a F1-string ends on a D4-brane). In the next chapter we will see how these ideas are further supported by the presence of the Chern-Simons terms in the supergravities, and by the structure of the world-volume effective actions of the branes.

We end this chapter mentioning some of the generalizations that have been done concerning intersecting branes, and some related solutions that were also studied.

The most logical generalization of the present work is to include internal waves (or KK momenta) and KK monopoles to the brane configurations. This problem was already mentioned before in this chapter. There are several ways to do this. The first one, which does not really rely on solving the equations of motion, consists in performing dualities on the ‘diagonal’ configurations like the ones presented here. One can then classify all the BPS configurations including all the branes of subsection 3.2.5, and this has indeed been done in [119]. Another way to include the branes with KK origin, is to generalize the metric ansatz to include at least some off-diagonal elements. However this seems to be a major complication to the equations discussed here. A short-cut to this complication is to generalize instead to a model comprising several scalars [107, 82, 111]. These extra scalars can now be taken to be some of the moduli fields arising from the internal metric, while the KK 2-form field strengths are easily inserted between the other antisymmetric tensor fields, their coupling to the scalars being a trace of their origin. The other tensor fields have also to be ‘multiplied’ according to the KK reduction procedure, as shown in the Appendix A. This leads to a proliferation of ‘matter’ fields, and thus of branes, which complicates the intersection rules that one can indeed find. Reading off the particular intersection rules of a single, 10- or 11-dimensional object becomes thus rather cumbersome. This would be however the price to pay to really unify all the branes of M-theory.

In this chapter we presented only extremal configurations of intersecting branes. The natural further step to take would be to consider also non-extremal configurations of intersecting branes. There is however a subtlety: there could be a difference between intersections of non-extremal branes, and non-extremal intersections of otherwise extremal branes. If we focus on bound states (and thus not on configurations of well separated branes), it appears that a non-extremal configuration would be characterized

for instance by \mathcal{N} charges and by its mass. There is only one additional parameter with respect to the extremal configurations. Physically, we could have hardly expected to have, say, as many non-extremality parameters as the number of branes in the bound state. Indeed, non-extremality can be roughly associated to the branes being in an excited state, and it would have thus been very unlikely that the excitations did not mix between the various branes in the bound state. Non-extremal intersecting brane solutions were found first in [120], and were derived from the equations of motion following a similar approach as here in [121, 122].

It is at first puzzling that the solutions found in the papers cited above follow exactly the same intersection rules as the ones for extremal bound states, (3.148). One could have hoped to recover now a much larger class of intersection, unconstrained by supersymmetry for instance. However what these ‘black intersections’ describe are compounds of branes which exert no force on each other, which have been gathered to the same location in transverse space and have now an excess of mass (with respect to the sum of their extremal energies). For intersections violating (3.148) there is generically a non-vanishing attractive or repulsive force. This prevents the participating branes to form simple bound states as the ones described in [120, 121, 122]. The solution presented in subsection 3.3.5, and its generalizations [123, 124], show that the bound states of branes that actually repel each other are indeed very different. Bound states of branes which attract each other also have different features, and will be discussed below. Note that the attractive, repulsive or vanishing nature of the force between two branes can be quite accurately accounted for in the case where the two branes are D-branes [26, 113].

The intersections presented in this chapter are actually orthogonal in the sense that each brane lies along some definite directions. It is however possible to consider branes lying at an angle with respect to some of the directions. Take for instance a string, or 1-brane, lying in the $\hat{1}$ - $\hat{2}$ plane, at an angle θ from the $\hat{1}$ direction. If the two directions are both compact (and of the same size), then the string will have a non-trivial winding number in both directions. Taking the mass of the configuration to be proportional to the total length of the string (which is quantized), the two charges will be respectively $q_1 \sim M \cos \theta$ and $q_2 \sim M \sin \theta$. We thus see that an ‘angled’ brane should not be confused with a marginal bound state of two strings, since the mass will go like $M \sim \sqrt{q_1^2 + q_2^2}$ instead of $\sim |q_1| + |q_2|$.

From D-brane considerations, it is possible to show [125] that some configurations of several D-branes intersecting at angles are still supersymmetric. This requires that the rotations applied to, say, the second brane with respect to the first one, are not general but of a restricted type. For example, if we consider two D2-branes, and we start with a configuration in which they are parallel to each other, the most general rotation belongs to $SO(4)$. It turns out that the only rotations consistent with supersymmetry are the ones belonging to a particular $SU(2)$ subgroup. For instance, one can go continuously from the $D2 \parallel D2$ configuration to the $D2 \cap D2 = 0$ one, but the $D2 \cap D2 = 1$ is never allowed.

Supergravity solutions corresponding to D-branes at angles were found in [126, 128, 127]. The resulting solutions contain as expected off-diagonal elements in the internal metric, and the derivation from the equations of motion as in [128] is accordingly rather intricate.

We still have to discuss the bound states of the configurations of branes which exert an attractive force between each other. These are truly bound states, in the sense that there is a non-vanishing binding energy (i.e. the mass of the bound state is lower than the sum of the masses of the constituents). They are also called non-marginal bound states, in opposition with the marginal ones treated in this chapter. The archetype of these bound states is the dyon of a Yang-Mills theory, which has a mass going like $M \sim \sqrt{q_{el}^2 + q_{mag}^2}$. If this dyon is embedded in a supersymmetric theory (SYM), it turns out that it preserves 1/2 of the supersymmetries. This is a higher fraction than for the marginal bound states, which preserve 1/4 in a generic configuration with two objects.

Such dyonic branes were presented in [129], where it is further shown that the configuration $M2 \cap M5 = 2$, or in another notation $M2 \subset M5$, is half-supersymmetric, and its mass is given by $M \sim \sqrt{q_2^2 + q_5^2}$. A particular feature of this solution is that, in order to prove that it solves the equations of motion, the Chern-Simons term in (3.3) plays a crucial rôle. Moreover, if the M5 lies in the directions from y_1 to y_5 , and the M2 in the y_1 and y_2 directions, then the 4-form field strength has also a non-trivial $F_{y_3 y_4 y_5 a}$ component. Intersections of branes where the constituents are these kind of $M2 \subset M5$ dyonic branes were considered in [130] for the extremal case, and in [131] in the non-extremal generalization. The intersection rules have to be satisfied for both the components of each dyonic brane.

Compactifying and applying dualities to the $M2 \subset M5$ bound state one can classify all such non-marginal bound states. These comprise the configurations $Dp \subset D(p+2)$, $F1 \subset Dp$ ($p \geq 1$), $Dp \subset NS5$ ($p \leq 5$) and the KK momentum orthogonal to any one of the branes (except the KK monopole). These configurations all preserve the same amount of supersymmetry as, for instance, the bigger brane alone. From the point of view of the world-volume of the latter, the smaller brane can be interpreted as a flux of a field strength of the world-volume effective action (see [132] for the D-branes). Note that other configurations with intersections which seem to violate (3.148) are actually non-marginal bound states of this kind. Take the $D2 \subset D4$ bound state and T-dualize in one direction of the compact space transverse to the D2 and longitudinal to the D4. We obtain a configuration such that $D3 \cap D3 = 2$, with an off-diagonal element in the internal metric directly related by dualities to the $F_{y_3 y_4 y_5 a}$ component in the $M2 \subset M5$ configuration. In fact, this clearly represents a D3-brane wrapped at an angle on the T^4 torus. The configurations of ‘angled’ branes and of dyonic branes thus belong to the same duality orbit, as the formulas for their mass might have suggested.

Other half-supersymmetric bound states of this class are the $SL(2, Z)$ multiplets of 1- and 5-branes in type IIB theory [40, 55], or more precisely the configurations $F1 \parallel D1$ and $NS5 \parallel D5$, also called (p, q) 1- and 5-branes, where p is the NSNS charge and q the RR charge of the compound. The classical solutions corresponding to this latter case were actually found more simply performing an $SL(2, Z)$ transformation on the F1 or NS5 solutions.

The main unsatisfactory point in the intersecting brane solutions presented in this chapter is that they are not localized in the relative transverse space. This is disappointing if one wanted, for instance, to have a supergravity description of the brane configurations used to derive field theory results, as in [64, 65].

In [102] (inspired by [133]) a solution is presented which corresponds to a $M5 \cap M5 = 1$ configuration, which follows the harmonic superposition rule, provided however that the harmonic functions depend on the respective relative transverse space (i.e. they are functions of two different spaces). The problem now is that the harmonic functions do not depend on the overall transverse space (which is 1-dimensional in the case above), the configuration thus not being localized there. A method actually inspired by the one presented here to derive the intersecting brane solutions, has been applied in [134] to the intersections of this second kind. Imposing that the functions depend on the relative transverse space(s) (with factorized dependence) and not on the overall one, the authors of [134] arrive at a formula for the intersections very similar to (3.148), with $\bar{q} + 3$ on the l.h.s. This rule correctly reproduces the $M5 \cap M5 = 1$ configuration, and moreover also all the configurations of two D-branes with 8 Neumann-Dirichlet directions, which preserve 1/4 supersymmetries but were excluded from the intersecting solutions derived in this chapter (only the configurations with 4 ND directions were found as solutions). One such configuration is e.g. $D0 \subset D8$.

Very recently, some approximate solutions were found in [135] which represented marginal bound states like, for instance, $D2 \subset D6$ and $NS5 \subset D6$, not only localized in transverse space, but where the smaller brane is additionally localized within the D6-brane. An analogous configuration representing $F1 \subset NS5$ was further shown to satisfy the ‘modified’ Laplace equation which was introduced in [136] to implicitly characterize the solutions with localized intersections. Unfortunately, the solutions discussed in [135] are valid only near the location of the bound state, and for branes within branes.

The main goal when considering intersecting brane solutions is certainly to find a general solution in which the intersection is localized in both relative transverse directions, and in the overall transverse space. Finding solutions representing branes with fully localized intersections is most interesting with respect to several issues as the world-volume dynamics of the branes involved, and also to analyze the possible corrections and the decoupling limit when field theory phenomena are derived from brane configurations as in e.g. [64, 65].

Chapter 4

The dynamics of open branes

In this chapter we will take a step beyond the classical description of p -branes. We aim at considering the dynamics of the quantum objects which are described at low energies by the classical p -brane solutions discussed in the previous chapter. It is now crucial that we specialize to the branes of M-theory, since it is the structure of the latter theory that we eventually want to uncover, and also because its particular features will turn out to be necessary ingredients in the discussion presented in this chapter.

The particular aspect of brane dynamics that we will focus on is the possibility that the branes of M-theory have to open, provided their boundary is restricted to lie along the world-volume of another brane. The original example, and by far the most thoroughly understood, is the type II fundamental string that can end on the D-branes, which are the (non-perturbative) objects carrying Ramond-Ramond (RR) charge. The D-branes were introduced in Section 2.5. We will show that the D-branes themselves can in some cases be open and have their boundaries attached to some other brane. It thus follows that there are no two distinct classes of branes, one with branes that can open and the other with branes which collect the boundaries of the open branes. Rather, most of the branes display both aspects.

Here we will adopt a different strategy than the one which led to the definition of D-branes, and which required a microscopic description of the open brane (the fundamental string in that case). We will start from the classical solutions which could lead to an interpretation in terms of branes ending on other branes, and check that when the equations of motions are completed with the correct source and Chern-Simons like terms, then the opening of the branes is always consistent with charge conservation. In the course of this derivation, we will see that many ingredients like supersymmetry, gauge invariance and world-volume effective actions are necessary for the subtle consistency of the whole picture.

The chapter is organized as follows. In Section 4.1 we suggest the transition between configurations with intersecting branes and configurations with open branes. In Section 4.2 we provide evidence that charge conservation allows for such an interpretation. The proof is a combination of arguments relying on the Chern-Simons terms of supergravity origin and on restoration of gauge invariance. Both arguments lead to the determination of the world-volume fields for the ‘host’ brane on which the open brane ends, as reviewed

in Section 4.3 for type II and M theories. Section 4.4 contains the discussion of two limiting cases. The last section contains brief speculations.

This chapter is essentially based on [2] (and also on part of [3]).

4.1 A different light on some intersections

General intersection rules between extremal p -branes can be derived for type II string theories and for M-theory, the latter being the (hypothetical) 11 dimensional theory the low energy effective action of which is 11 dimensional supergravity. The complete set of rules was presented in [98, 101, 102, 105, 1], and we gave a thorough account in the previous chapter on how these rules can be altogether derived from the low-energy supergravity equations of motion. The intersections we are concerned with are orthogonal (in the sense that each brane of the configuration is either parallel or perpendicular to each direction) and the bound states they describe have vanishing binding energy, i.e. they are marginal bound states. This latter property makes them supersymmetric, and thus leaves us the hope that some of the features of the quantum bound states are related to the classical supergravity solutions discussed in the previous chapter.

Let us recall the condition that guarantees that a multiple brane configuration has zero binding energy (see [1] and Section 3.3). It is sufficient that each intersecting pair satisfies the following rule: the intersection of a p_a -brane with a p_b -brane must have dimension

$$\bar{q} = \frac{(p_a + 1)(p_b + 1)}{D - 2} - 1 - \frac{1}{2}\varepsilon_a a_a \varepsilon_b a_b, \quad (4.1)$$

where D is the space-time dimension, a_a and a_b are the couplings of the field strengths to the dilaton in the Einstein frame and ε_a and ε_b are $+1$ or -1 if the corresponding brane is electrically or magnetically charged. In eleven dimensional supergravity, the a 's are equal to zero while in ten dimensions $\varepsilon a = \frac{1}{2}(3 - p)$ for Ramond-Ramond (RR) fields and $a = -1$ for the Neveu-Schwarz-Neveu-Schwarz (NSNS) three form field strength. Note that we introduced a slight change of notations with respect to (3.148).

The cases in which we will be primarily concerned here are the ones in which \bar{q} has the same dimension as the boundary of one of the two constituent branes, i.e. $p_a - 1$ or $p_b - 1$. We can single out such cases from the complete list discussed in Section 3.3. In M-theory, we have only one case, namely:

$$M2 \cap M5 = 1. \quad (4.2)$$

The prefix to M2 and M5 stands for their 11 dimensional (M-theoretic) origin. In ten dimensional type II theories, we have three different series of cases:

$$F1 \cap Dp = 0, \quad p = 0, \dots, 8, \quad (4.3)$$

$$Dp \cap D(p+2) = p - 1, \quad p = 1, \dots, 6, \quad (4.4)$$

$$Dp \cap NS5 = p - 1, \quad p = 1, \dots, 6. \quad (4.5)$$

Above, F1 stands for the fundamental string, Dp stands for a p -brane carrying RR charge and NS5 stands for the solitonic NSNS 5-brane. Of all the cases above, the $p = 0$ case

of (4.3) and the $p = 6$ case of (4.5) are particular in that the ‘host’ brane (the brane on which the open brane should end) has the dimension of the boundary itself. We will have to treat these two cases separately. Also, the cases in which a D8-brane participates to the configuration should be treated in the framework of massive type IIA supergravity [137].

All the cases above, (4.2)–(4.5), will turn out to effectively correspond to a well-defined possibility for the first brane to be open and to end on the second one. It is certainly surprising that such a good guess is essentially based on the knowledge of the classical solutions, which are actually ‘smeared’ over all the relative transverse directions, and thus represent at best delocalized intersections. In the following, we will represent a configuration in which a p_a -brane (the open brane) ends on a p_b -brane (the host brane) by the symbol $p_a \mapsto p_b$. Accordingly, the intersections (4.2)–(4.5) correspond to the following open brane configurations: $M2 \mapsto M5$, $F1 \mapsto Dp$, $Dp \mapsto D(p+2)$ and $Dp \mapsto NS5$. Note that the $p = 1$ case of (4.3) and the $p = 5$ case of (4.5) both seem to give rise to a second possibility, respectively $D1 \mapsto F1$ and $NS5 \mapsto D5$.

That the fundamental type II strings can be open, with their ends tied to the world-volume of a D-brane, is the first appearance of ‘open branes’ of the kind we are considering here. Indeed, the fundamental type II strings are generically closed, and they carry a charge (the electric charge of the NSNS 3-form field strength) that would not be conserved if they were to open freely at any point of space-time. Charge conservation however allows the strings to open provided their ends lie along the world-volume of a D-brane [48]. The presence of the RR-charged D-branes is required by the conjectured U-duality non-perturbative symmetry of the theory. The first case which was considered, and the only one to have a microscopic description, is thus $F1 \mapsto Dp$.

The repeated use of dualities allows then to predict all the other configurations with open branes, along the lines of [99, 74]. The configurations with $Dp \mapsto D(p+2)$ and $Dp \mapsto NS5$ are obtained by S- and T-dualities, while the case $M2 \mapsto M5$ is obtained considering, for instance, the strong coupling, 11 dimensional description of the type IIA configuration $F1 \mapsto D4$.

From the classical point of view, the main obstacle towards the opening of branes is charge conservation. Generically and in D -dimensional flat space-time, the charge of a p -brane is measured performing an integral of the relevant field strength on a $(D-p-2)$ -dimensional sphere S^{D-p-2} surrounding the brane in its transverse space:

$$Q_p \sim \int_{S^{D-p-2}} *F_{p+2}.$$

If the brane is open, then we can slide the S^{D-p-2} off the loose end and shrink it to zero size. This implies that the charge must vanish. This conclusion is avoided if in the process above, the S^{D-p-2} necessarily goes through a region in which the equation

$$d * F_{p+2} = 0 \tag{4.6}$$

no longer holds. Only when the r.h.s. of this equation is strictly vanishing we are allowed to deform the S^{D-p-2} while keeping the charge unchanged (and thus conserved). Thus, any higher dimensional source in the r.h.s. of (4.6), other than the source of the p -brane

itself, is a potential topological obstruction to sliding off the S^{D-p-2} . This is the core of the description of how extended objects carrying a conserved charge can be open, provided they end on some other higher dimensional object.

The matter of concretely providing the source term in (4.6) which ensures charge conservation for the open branes can be addressed in different ways. The approach of Strominger [99] was based on the knowledge of the world-volume effective actions, and more precisely of the couplings between world-volume fields and (pull-backs to the world-volume of) space-time fields. The boundaries of the open branes are then identified with sources for the world-volume fields. An alternative approach was given by Townsend [138] who showed that the presence of Chern-Simons type terms in supergravity allows for charge conservation for well defined pairings of open and ‘host’ branes.

In the rest of this chapter, we will show that in every case where the zero binding energy condition is satisfied and where the intersection has dimension $p_a - 1$, it is possible to open the p_a -brane along its intersection with the p_b -brane. This means that charge conservation is always preserved. The general case will be $p_a \leq p_b$, while the case $p_a = p_b + 1$ will necessitate a special treatment. The boundary of each open brane configuration carries a charge living in the closed brane on which it terminates. To each such configuration corresponds a second one. Their respective boundary charges are electric-magnetic dual of each other in the host brane and are coupled to field strengths in that brane. The fields pertaining to all branes are all related by their origin, and can be seen as Goldstone bosons of broken supersymmetry.

More precisely, we shall first review and apply the analysis of charge conservation in terms of the Chern-Simons type terms [138] present in supergravity to the zero binding energy configurations. We shall then complete this analysis with gauge invariance and supersymmetry considerations and obtain all world-volume field strengths and their coupling to boundary charges.

Note that configurations with branes ending on other branes were crucially used in [64] and in all the subsequent literature, where some issues related to charge conservation in the opening mechanism were also considered.

4.2 Charge conservation for the open branes

In order for a p_a -brane to open along its intersection with a p_b -brane, its boundary must behave, by charge conservation, as an induced charge living in the p_b -brane. Considering the general case for which $p_a \leq p_b$, this charge will act as a source for a field strength on the p_b -brane world-volume. As shown below this field strength is related to (and thus determined by) the Chern-Simons type term present in the supergravity action.

Since the interplay of the different ingredients is rather subtle, we present in the first place a specific example, and afterwards we proceed to the general case.

4.2.1 A D2-brane ending on a D4-brane

We work out in this subsection the example of an open D2-brane ending on a D4-brane in the framework of type IIA string theory

The equation of motion for the 4-form field strength has to be supplemented by the Chern-Simons term present in the full IIA supergravity action (3.6), and by the source due to the presence of the D2-brane. The equation thus reads, neglecting the dilaton and all numerical factors:

$$d * F_4 = F_4 \wedge H_3 + \mu_2 \delta_7. \quad (4.7)$$

The tension of the (elementary) D2-brane is thus μ_2 , and the δ_7 localizes the D2-brane in the transverse space. Since there is also a D4-brane in the configuration, the Bianchi identity for F_4 is also modified by a source term:

$$dF_4 = \mu_4 \delta_5. \quad (4.8)$$

On the other hand, due to the absence of NS5-branes, H_3 can be globally defined as $H_3 = dB_2$. The equation (4.7) can be rewritten as:

$$d(*F_4 - F_4 \wedge B_2) = \mu_2 \delta_7 - \mu_4 \delta_5 \wedge B_2. \quad (4.9)$$

We can now integrate both sides of this equation over a 7-sphere S^7 which intersects the D2-brane only once (this is possible only if the D2-brane is open). The result is:

$$0 = \mu_2 - \mu_4 \int_{S^2} \hat{B}_2, \quad (4.10)$$

where the hat denotes the pull-back to the world-volume of the D4-brane of a space-time field.

The expression (4.10) is actually not exact (as we discuss below), but it is relevant because it allows us to see that the Chern-Simons term indicates the presence on the world volume of the D4-brane of a 2-form field strength, for which the (string-like) boundaries of the D2-branes act as magnetic charges. As we will discuss shortly, the gauge invariant combination is:

$$\mathcal{F}_2 = dV_1 - \hat{B}_2. \quad (4.11)$$

Considering now the world-volume action of the D2-brane, we know that there is a minimal coupling to the RR 3-form potential:

$$I_{D2} = \mu_2 \int_{W_3} \hat{A}_3 + \dots$$

When the D2-brane is open, the gauge transformation $\delta A_3 = d\Lambda_2$ becomes anomalous:

$$\delta I_{D2} = \mu_2 \int_{(\partial W)_2} \hat{\Lambda}_2.$$

The standard way to cancel this anomaly is by constraining the boundary $(\partial W)_2$ to lie on the D4-brane world-volume where a 2-form gauge potential V_2 , transforming as

$\delta V_2 = \hat{\Lambda}_2$, couples to it. The boundary of the D2-brane is now an electric source for the 3-form field strength built out from this potential. Again, the gauge invariant combination is given by:

$$\mathcal{G}_3 = dV_2 - \hat{A}_3. \quad (4.12)$$

The analysis of the Goldstone modes of broken supersymmetry and of broken translation invariance, and the requirement that these bosonic and fermionic modes still fit into a representation of the unbroken supersymmetries, forces us to identify the two field strengths \mathcal{F}_2 and \mathcal{G}_3 by an electric-magnetic duality on the D4-brane world-volume:

$$\mathcal{F}_2 = \star \mathcal{G}_3. \quad (4.13)$$

Moreover, we could have analyzed instead the (more familiar) configuration of a fundamental string ending on the D4-brane. Preservation of the gauge invariance would have led to the expression (4.11). We would have found that the end point of the string behaves like an electric charge for the 2-form field strength and like a magnetic charge for the 3-form field strength. Thus we conclude that the boundaries of the string and of the membrane are electric-magnetic dual objects on the world-volume of the D4-brane.

We can now get back to the equation (4.10). The point is that if we formulate the world-volume effective theory of the D4-brane in terms of the 2-form field strength \mathcal{F}_2 , then preservation of gauge invariance of the D2-brane action, leading to (4.12), and the electric-magnetic duality (4.13) together imply the presence of a supplementary term:

$$I_{D4} = \dots + \mu_4 \int_{W_5} \mathcal{F}_2 \wedge \hat{A}_3. \quad (4.14)$$

This term is often called a Wess-Zumino term in the D4-brane action, and its presence can also be traced back to more general considerations on the gauge invariance of the D4-brane action in general backgrounds (see for instance [73, 139]).

The presence of this term is crucial since it implies that there is an additional term in the r.h.s. of the equation (4.7), of the form $\mu_4 \delta_5 \wedge \mathcal{F}_2$. This in turn means that the second term in the r.h.s. of the equation (4.9) now should read:

$$-\mu_4 \delta_5 \wedge (B_2 + \mathcal{F}_2) \equiv -\mu_4 \delta_5 \wedge dV_1,$$

where we have used (4.11).

The expression relating the tensions of the various objects can then be written as:

$$\mu_2 = \mu_4 \hat{\mu}_1, \quad \hat{\mu}_1 = \int_{S^2} dV_1. \quad (4.15)$$

The difference between the equation above and (4.10) is that now the singular part in \mathcal{F}_2 due to the string-like source is carried by the purely world-volume field, and not by the pull-back of the space-time field B_2 .

The equations of motion and Bianchi identities for the world-volume field strength \mathcal{F}_2 , completed with electric and magnetic sources, are then:

$$\begin{aligned} d \star \mathcal{F}_2 &= \hat{\mu}_0 \delta_4 - \hat{F}_4, \\ d \mathcal{F}_2 &= \hat{\mu}_1 \delta_3 - \hat{H}_3. \end{aligned}$$

4.2.2 Charge conservation using the Chern-Simons terms

We now go back to the general case, and we focus on the Chern-Simons terms, as in [138], since we have seen in the example above that they allow to precisely guess how the charge conservation occurs.

We distinguish three different types of configurations ($p_a \mapsto p_b$) of a p_a -brane ending on a p_b -brane: (electric \mapsto magnetic), (electric \mapsto electric) and (magnetic \mapsto magnetic), where by electric brane we mean a brane which couples minimally to a field of supergravity (all the field strengths are taken by convention to be of rank $n \leq D/2$). Note that the (magnetic \mapsto electric) never appears since it corresponds, in the supergravities we consider here, to $p_a \geq p_b + 2$.

Let us first consider the (electric \mapsto magnetic) case in detail. The equations of motion for the field strength corresponding to the p_a -brane take the generic form

$$d * F_{p_a+2} = F_{D-p_b-2} \wedge F_{p_b-p_a+1} + \mu_a^e \delta_{D-p_a-1}. \quad (4.16)$$

Here and in what follows wedge products are defined up to signs and numerical factors irrelevant to our discussion.

In the r.h.s. of (4.16) the first term comes from the variation of the Chern-Simons term

$$\int A_{p_a+1} \wedge F_{D-p_b-2} \wedge F_{p_b-p_a+1}, \quad (4.17)$$

where F_{D-p_b-2} is the field strength with magnetic coupling to the p_b -brane and $F_{p_b-p_a+1}$ is a $(p_b - p_a + 1)$ -form field strength present in the theory. Let us for the moment just postulate the presence of such a Chern-Simons term. The second term in (4.16) is the p_a -brane charge density, where δ_{D-p_a-1} is the Dirac delta function in the directions transverse to the p_a -brane. It should also take into account the fact that the brane is open.

The equation (4.16) requires a couple of comments. We want to study the deformation of intersecting closed brane configurations with zero binding energy when one of the branes is sliced open along the intersection. Firstly, notice that we have introduced an explicit source term for the electric brane since, to study its opening, we want to extend the validity of the usual closed brane solution on the branes. This term is required because the supergravity equations of motion from which the intersecting brane solutions are derived do not contain any source term and are therefore valid only outside the sources. Secondly we have to verify that, in the limit where the open brane closes, the contribution from the Chern-Simons term vanishes since the closed brane configuration depends only on F_{p_a+2} and F_{D-p_b-2} and not on the third field $F_{p_b-p_a+1}$. The equation for the latter is

$$d * F_{p_b-p_a+1} = F_{D-p_b-2} \wedge F_{p_a+2}.$$

It is compatible with a solution where $F_{p_b-p_a+1} = 0$ since, when the two closed branes are orthogonal, and the fields are smeared over the directions longitudinal to the branes (as in the classical solutions), the electric F_{p_a+2} and the magnetic F_{D-p_b-2} have necessarily a common index.

Taking into account that in the configuration considered,

$$F_{p_b-p_a+1} = dA_{p_b-p_a},$$

and

$$dF_{D-p_b-2} = \mu_b^m \delta_{D-p_b-1},$$

one can rearrange (4.16) in the following way:

$$d(*F_{p_a+2} - F_{D-p_b-2} \wedge A_{p_b-p_a}) = \mu_a^e \delta_{D-p_a-1} - \mu_b^m \delta_{D-p_b-1} \wedge A_{p_b-p_a}. \quad (4.18)$$

The integration of (4.18) over a S^{D-p_a-1} sphere, which intersects the p_a -brane only at a point and the p_b -brane on a $S^{p_b-p_a}$ sphere which surrounds the intersection, gives

$$0 = \mu_a^e - \mu_b^m \int_{S^{p_b-p_a}} A_{p_b-p_a}.$$

This equation can be rewritten as:

$$\mu_a^e = \mu_b^m \cdot \hat{\mu}_I, \quad \text{with} \quad \hat{\mu}_I \equiv \int_{S^{p_b-p_a}} A_{p_b-p_a}. \quad (4.19)$$

We see that the pull-back $\hat{A}_{p_b-p_a}^{(p_b+1)}$ of the potential $A_{p_b-p_a}$ on the closed p_b -brane behaves like a $(p_b - p_a)$ -form field strength, magnetically coupled to the boundary. We will see in the next subsection that to preserve gauge invariance one has to add to this field a $(p_b - p_a)$ -form $dW_{p_b-p_a-1}$ defined on the world volume of the p_b -brane. This procedure will actually change the definition of the charge $\hat{\mu}_I$, without nevertheless altering the present discussion. We thus define the field strength

$$\mathcal{G}_{p_b-p_a} = dW_{p_b-p_a-1} - \hat{A}_{p_b-p_a}^{(p_b+1)}. \quad (4.20)$$

The charge $\hat{\mu}_I$ is then simply a magnetic charge for \mathcal{G} . On the p_b -brane, and outside sources like $\hat{\mu}_I$, the field strength $\mathcal{G}_{p_b-p_a}$ satisfies the Bianchi identity:

$$d\mathcal{G}_{p_b-p_a} = -d\hat{A}_{p_b-p_a}^{(p_b+1)} = -\hat{F}_{p_b-p_a+1}^{(p_b+1)}.$$

For the (electric \mapsto electric) case the reasoning goes along the same lines starting from the equation of motion:

$$d * F_{p_a+2} = *F_{p_b+2} \wedge F_{p_b-p_a+1} + \mu_a^e \delta_{D-p_a-1}. \quad (4.21)$$

This time $d * F_{p_b+2} = \mu_b^e \delta_{D-p_b-1}$.

For the (magnetic \mapsto magnetic) case we have to consider the Bianchi identities instead of the equations of motion:

$$dF_{D-p_a-2} = F_{D-p_b-2} \wedge F_{p_b-p_a+1} + \mu_a^m \delta_{D-p_a-1}. \quad (4.22)$$

Here also, we have $dF_{D-p_b-2} = \mu_b^m \delta_{D-p_b-1}$. In both of these two last cases, the Chern-Simons like term comes from the additional terms that are present in the definitions of the gauge invariant field strengths in type II supergravities.

We should now emphasize that for all the intersections (4.2)–(4.5), a Chern-Simons term of the type described in (4.16), (4.21) or (4.22) is always present. This is really a non-trivial result since, as we had already stated, the Chern-Simons terms played absolutely no rôle in deriving the classical intersection rules. One sees that requiring the possibility of open brane configurations actually would lead to the introduction of these Chern-Simons terms which from the supergravity point of view are only required by supersymmetry.

We will review in Section 4.3 all the relevant cases in 10 and 11 dimensional maximal supergravities.

4.2.3 Electric-magnetic duality on the brane

We saw in the preceding subsection how charge conservation leads to the existence of a field \mathcal{G} whose magnetic source is the charge of the boundary. However there is an alternative description of this charge. It arises from the consideration of gauge invariance for the A_{p_a+1} potential and will lead to the identification of the world-volume degrees of freedom.

Consider the minimal coupling of the A_{p_a+1} potential to the open p_a -brane. The boundary of the p_a -brane breaks gauge invariance under $A_{p_a+1} \rightarrow A_{p_a+1} + d\Lambda_{p_a}$ as

$$\delta \int_{W_{p_a+1}} A_{p_a+1} = \int_{(\partial W)_{p_a}} \Lambda_{p_a}.$$

The space-time volume $(\partial W)_{p_a}$ swept by the boundary of the p_a -brane is the world-volume of a (p_a-1) -brane living on the host p_b -brane. In order to restore gauge invariance one has to introduce [55] a field V_{p_a} living on the p_b -brane and transforming like $V_{p_a} \rightarrow V_{p_a} + \Lambda_{p_a}$. The resulting gauge invariant field strength on the p_b -brane is

$$\mathcal{F}_{p_a+1} = dV_{p_a} - \hat{A}_{p_a+1}^{(p_b+1)}.$$

In string theory, the existence of vector field V_1 degrees of freedom stems from the zero mass excitations of open strings. In the context of supergravity V_{p_a} emerges from broken supersymmetry. Indeed, the introduction of a p_b -brane breaks half of the space-time supersymmetries. The broken supersymmetries give rise to eight massless on-shell fermionic degrees of freedom of the p_b -brane and their bosonic partners. $D - p_b - 1$ of them are the Goldstone translation modes of the p_b -brane and are world-volume scalars. Because of the remaining (unbroken) supersymmetry, we must have a total of eight on-shell bosonic degrees of freedom. The field strength \mathcal{F}_{p_a+1} exactly accounts for the remaining Goldstone degrees of freedom.

The charge of the boundary of the p_a -brane is measured by the integral:

$$\hat{\mu}_I = \int_{S^{p_b-p_a}} \star \mathcal{F}_{p_a+1}, \quad (4.23)$$

where \star indicates the Hodge dual on the p_b -brane and the $S^{p_b-p_a}$ encircles the (p_a-1) boundary. The $\hat{\mu}_I$ in (4.23) is naturally identified with the one appearing in (4.19). This

identification is in fact required by supersymmetry, as all massless degrees of freedom have already been accounted for. Using (4.20), the two expressions for the charge $\hat{\mu}_I$, (4.23) and (4.19), imply

$$\mathcal{F}_{p_a+1} = \star \mathcal{G}_{p_b-p_a}.$$

It remains to justify why we introduced a potential $W_{p_b-p_a-1}$ in the definition of \mathcal{G} given in (4.20). To this end notice that the Chern-Simons term (4.17) could have been written as

$$\int A_{p_b-p_a} \wedge F_{D-p_b-2} \wedge F_{p_a+2}.$$

In this form the Chern-Simons term is suitable to study the opening of a (p_b-p_a-1) -brane on the p_b -brane. This study goes along the same lines as before, with the interchange of the rôle of \mathcal{F} and \mathcal{G} . The introduction of W is now required to preserve gauge invariance under a transformation of $A_{p_b-p_a}$, gauge invariance which otherwise would be broken by the presence of the boundary of the (p_b-p_a-1) -brane. For any ‘host’ p_b -brane, open brane configurations always occur in dual pairs. Their boundary charges are respectively electric and magnetic sources for \mathcal{F} , and vice-versa for \mathcal{G} .

To summarize, and taking into account the world-volume sources produced by the boundaries of the open p_a - and (p_b-p_a-1) -brane, the equations for the gauge field living on the host p_b -brane are:

$$\begin{aligned} d \star \mathcal{F}_{p_a+1} &= \hat{\mu}_{p_a-1} \hat{\delta}_{p_b-p_a+1} - \hat{F}_{p_b-p_a+1}^{(p_b+1)}, \\ d \mathcal{F}_{p_a+1} &= \hat{\mu}_{p_b-p_a-2} \hat{\delta}_{p_a+2} - \hat{F}_{p_a+2}^{(p_b+1)}. \end{aligned}$$

4.3 Open branes and world-volume fields for the host brane

In this section we review all the cases leading to open branes, (4.2)–(4.5), identifying for each one the Chern-Simons term which participates to the charge conservation, and the world-volume field on the host brane under which the boundary of the open brane is charged.

We take the actions of type II and 11 dimensional supergravities as in (3.6), (3.8) and (3.3) respectively. The precise definitions of the field strengths are also given in Section 3.1.

The open brane configurations of type IIA, type IIB and $D = 11$ supergravities are listed respectively in Tables 4.1, 4.2 and 4.3. In each table we list in the first column the configurations, in the second column the relevant equations of motion or Bianchi identities, and in the third column the equations for the world-volume field strength (we neglect the sources in all the equations). The configurations are arranged by dual electric-magnetic pairs with respect to the host branes. In the tables, a superscript recalls the electric or magnetic nature of each brane (the D3-brane being self-dual), and in the type IIB case we have defined the RR 1-form field strength $F_1 = d\chi$.

$F1^e \mapsto D2^e$	$d * H_3 = *F'_4 \wedge F_2$	$d \star \mathcal{F}_2 = -\hat{F}_2^{(3)}$
$D0^e \mapsto D2^e$	$d * F_2 = *F'_4 \wedge H_3$	$d\mathcal{F}_2 = -\hat{H}_3^{(3)}$
$F1^e \mapsto D4^m$	$d * H_3 = F'_4 \wedge F'_4$	$d \star \mathcal{F}_2 = -\hat{F}_4^{(5)}$
$D2^e \mapsto D4^m$	$d * F'_4 = F'_4 \wedge H_3$	$d\mathcal{F}_2 = -\hat{H}_3^{(5)}$
$F1^e \mapsto D6^m$	$d * H_3 = F_2 \wedge *F'_4$	$d \star \mathcal{F}_2 = - * \hat{F}_4^{(7)}$
$D4^m \mapsto D6^m$	$dF'_4 = F_2 \wedge H_3$	$d\mathcal{F}_2 = -\hat{H}_3^{(7)}$
$D0^e \mapsto NS5^m$	$d * F_2 = H_3 \wedge *F'_4$	$d \star \mathcal{G}_1 = - * \hat{F}_4^{(6)}$
$D4^m \mapsto NS5^m$	$dF'_4 = H_3 \wedge F_2$	$d\mathcal{G}_1 = -\hat{F}_2^{(6)}$
$D2^e \mapsto NS5^m$	$d * F'_4 = H_3 \wedge F'_4$	$d \star \mathcal{G}_3 = d\mathcal{G}_3 = -\hat{F}_4^{(6)}$

Table 4.1: Open branes in type IIA

The configurations of type II theories fall into two classes. The first is associated with the configuration with the fundamental string ending on Dp-branes, its dual being always a D(p - 2)-brane [99]. The second class consists of D-branes ending on the solitonic NS5-brane. These are the only branes which can end on the NS5-brane.

Notice in Table 4.1 the “opening” of a D0-brane on the D2-brane and on the NS5-brane. It corresponds to an intersection of dimension -1 which does not make obvious sense. However we have seen in the previous chapter that the relation (4.1) stills holds for the D0-brane world-line intersections in the Euclidean provided the electric fields are given imaginary values.

In type IIB theory, we have seen in Section 2.4 that there exists an $SL(2, R)$ duality symmetry at the level of the supergravity which is broken to $SL(2, Z)$ in the full string theory. The last two cases of Table 4.2 are built up from objects which are $SL(2, R)$ -dual to each other. It follows that the configurations $D1^e \mapsto F1^e$ and $NS5^m \mapsto D5^m$ exist by S-duality. In fact there is a continuum of dyonic configurations [40] obtained by acting with the $SL(2, R)$ group¹. These configurations do not have duals. Indeed, the dual configurations would involve intersections with a (-1)-brane, which do not make sense even in the Euclidean.

Let us now review the outcome of the analysis above from the point of view of the world-volume effective theory of the host brane. The approach based on the Chern-

¹ Actually, due to the existence of (p, q) 1- and 5-branes, it is more convenient in this cases to think about string networks [140] and webs of (p, q) 5-branes [141] instead of strings ending on strings and 5-branes ending on 5-branes respectively.

$F1^e \mapsto D3^*$	$d * H_3 = F'_5 \wedge F'_3$	$d \star \mathcal{F}_2 = -\hat{F}_3^{(4)}$
$D1^e \mapsto D3^*$	$d * F'_3 = F'_5 \wedge H_3$	$d\mathcal{F}_2 = -\hat{H}_3^{(4)}$
$F1^e \mapsto D5^m$	$d * H_3 = F'_3 \wedge F'_5$	$d \star \mathcal{F}_2 = -\hat{F}_5^{(6)}$
$D3^* \mapsto D5^m$	$dF'_5 = F'_3 \wedge H_3$	$d\mathcal{F}_2 = -\hat{H}_3^{(6)}$
$F1^e \mapsto D7^m$	$d * H_3 = F_1 \wedge *F'_3$	$d \star \mathcal{F}_2 = - * \hat{F}_3^{(8)}$
$D5^m \mapsto D7^m$	$dF'_3 = F_1 \wedge H_3$	$d\mathcal{F}_2 = -\hat{H}_3^{(8)}$
$D1^e \mapsto NS5^m$	$d * F'_3 = H_3 \wedge F'_5$	$d \star \mathcal{G}_2 = -\hat{F}_5^{(6)}$
$D3^* \mapsto NS5^m$	$dF'_5 = H_3 \wedge F'_3$	$d\mathcal{G}_2 = -\hat{F}_3^{(6)}$
$F1^e \mapsto D1^e$	$d * H_3 = *F'_3 \wedge F_1$	$d \star \mathcal{F}_2 = -\hat{F}_1^{(2)}$
$D5^m \mapsto NS5^m$	$dF'_3 = H_3 \wedge F_1$	$d \star \mathcal{G}_6 = -\hat{F}_1^{(6)}$

Table 4.2: Open branes in type IIB

Simons terms actually allows us to guess the field content of this world-volume theory, and the guess is always confirmed when the theory is known by other means.

All the Dp -branes have a world-volume effective theory which can be formulated in terms of a 2-form field strength \mathcal{F}_2 . The electric charges are the end points of the fundamental strings, while the magnetic charges are the boundaries of the $D(p-2)$ -branes (as in [99]). Note the interesting case of the D3-brane on the world-volume of which the S-duality between fundamental strings and D-strings becomes electric-magnetic duality between their end points. In the case of the D-branes, the presence of the 2-form field strength is supported by the quantum stringy computation which gives super-Yang-Mills as the low energy effective action of the D-branes.

On the world-volume of the IIA NS5-brane, we can have the boundaries of the D2- and D4-brane. The boundary of the D2-brane is self-dual and thus couples to a self dual 3-form field-strength \mathcal{G}_3^+ , while the boundary of the D4-brane couples magnetically to a 1-form \mathcal{G}_1 deriving from a scalar potential. This scalar potential is nothing else than the scalar associated with the 11th direction which remains after ‘vertical’ reduction of the M5-brane action. Indeed, Table 4.3 shows that there is a self-dual 3-form living on the world-volume of the M5-brane.

For the IIB NS5-brane, again all the IIB D-branes can have boundaries on it. The D1- and the D3-brane boundaries are respectively the electric and magnetic charge related

$$M2^e \mapsto M5^m \quad d * F_4 = F_4 \wedge F_4 \quad d \star \mathcal{G}_3 = d\mathcal{G}_3 = -\hat{F}_4^{(6)}$$

Table 4.3: Open branes in $D = 11$ supergravity

to a 2-form field strength $\mathcal{G}_2 = d\tilde{V}_1 - \hat{A}_2$ which can be considered the S-dual of the \mathcal{F}_2 field on the D5-brane. The boundary of the D5-brane couples electrically to a (non-propagating) 6-form field strength \mathcal{G}_6 . This 6-form should be related to the mass term in the IIB NS5-brane action as discussed in [142], and could play a rôle in the definition of an $SL(2, Z)$ invariant IIB 5-brane action.

This world-volume field content of the NS5-branes was derived in [57] by directly identifying the Goldstone modes of the supergravity solution.

4.4 Two more cases of open branes

In this section we consider the two limiting cases of (4.3) and (4.5), which derive from consistent closed classical solutions, but seem at first sight inconsistent with charge conservation if there really is an open brane.

These two cases are $F1 \mapsto D0$ and $D6 \mapsto NS5$. It is straightforward to note that the arguments given in Section 4.2 do not apply to these cases, and this can be seen in two ways: first of all, there is no longer a topological obstruction for sliding off the S^7 (respectively the S^2) sphere, since the host brane now coincides with the boundary of the open brane; the charge seems thus no longer conserved. Secondly, gauge invariance is necessarily broken because in both cases one would need a world-volume gauge potential with rank equal to the dimension of the world-volume itself, and this has clearly no dynamics.

The resolution of this problem is to go to the framework of massive type IIA supergravity [137]. In this theory, the 1-form (RR) potential is no longer dynamical, but it is rather ‘eaten up’ by the 2-form (NSNS) potential that becomes massive. One can still write the theory in terms of a 1-form potential, but the gauge invariant 2-form field strength is modified to be:

$$F'_2 = dA_1 + mB_2. \quad (4.24)$$

This supergravity has a cosmological constant which is proportional to m^2 , and many more non-trivial terms with respect to the ‘massless’ IIA supergravity, but we will not need them for the present discussion.

Since the kinetic term for A_1 will be proportional to $F_2'^2$, the structure of (4.24) is enough to ensure charge conservation. However, before moving to the explicit proof, we should justify the use of massive type IIA supergravity. This can be seen as follows. One can go to a formulation in which m is no longer an arbitrary constant, but it is dynamically fixed to be a constant. This can be done as in [143] by the introduction

of a (non propagating) 9-form potential with 10-form field strength. The 0-form dual field strength then satisfies $dF_0 = 0$, fixing m to be $F_0 = m$. One can now see, in this formulation, that 8 dimensional objects can in fact be sources for the ‘cosmological’ mass. The equation for the 0-form field can indeed be modified by the addition of a source term:

$$dF_0 = \mu_8 \delta_1.$$

The source term will produce a discontinuous jump of magnitude μ_8 in the mass m when the 8-brane, which separates space-time in two disconnected parts, is crossed. In the context of type IIA theory, this domain wall is nothing else than a D8-brane. We thus see that massive IIA supergravity naturally appears as the low-energy effective action of type IIA string theory in a background containing D8-branes².

It is now a simple matter to prove charge conservation (see for instance [144, 145, 146]), actually simpler than for the cases treated in Section 4.2.

Consider first the fundamental string ending on a D0-brane, $F1 \mapsto D0$. because of the presence of B_2 in the expression (4.24), the equation of motion of B_2 is modified by an additional term in the r.h.s.:

$$d * H_3 = \mu_1 \delta_8 - m * F'_2. \quad (4.25)$$

We can take $m = \mu_8$ (we are always putting all numerical and sign factors to one, and neglecting the dilaton), signaling the presence of a D8-brane. The F1 can actually end on the D8-brane by the usual mechanism. The presence of the D0-brane at the end of the F1 is translated into the fact that:

$$d * F'_2 = \mu_0 \delta_9. \quad (4.26)$$

Integrating (4.25) over an S^8 sphere which intersects the string only once, we obtain that [144]:

$$0 = \mu_1 - \mu_8 \int_{S^8} * F'_2.$$

The integral in the second term gives nothing else than the charge of the D0-brane, μ_0 . We thus have $\mu_1 = \mu_8 \mu_0$, and the charge of the fundamental string is conserved. This can actually be seen the other way round, stating that in a non-trivial D8-brane background, a D0-brane can exist only if a string ends on it. The reverse is of course true when $m = 0$.

The case of the D6-brane ending on the NS5-brane is similar [146], except that now it is based on the Bianchi identities for the 2-form F'_2 . These read:

$$dF'_2 = \mu_6 \delta_3 + m H_3. \quad (4.27)$$

Identifying m with μ_8 and integrating over an S^3 cutting the D6 once, we obtain:

$$0 = \mu_6 + \mu_8 \int_{S^3} H_3,$$

² On one side of the D8-brane the cosmological mass m can actually vanish, giving back the usual massless type IIA theory in that region.

with the integral being the charge of the NS5-brane μ_5 .

We can also check that gauge invariance of the open brane is still preserved in these two cases. Here the compensating gauge variation will come from the bulk action instead than coming from the world volume of the host brane.

Under the gauge variation $\delta B_2 = d\Lambda_1$, the open string produces an anomalous term along the world-line of the D0-brane:

$$\delta \left(\mu_1 \int_{W_2} B_2 \right) = \mu_1 \int_{W_1} \hat{\Lambda}_1.$$

The variation of the bulk kinetic term for F'_2 exactly compensates it:

$$\delta \left(\frac{1}{2} \int_{M_{10}} *F'_2 \wedge F'_2 \right) = \int_{M_{10}} *F'_2 \wedge m\delta B_2 = \mu_8 \int_{M_{10}} d * F'_2 \wedge \Lambda_1 = \mu_8 \mu_0 \int_{W_1} \hat{\Lambda}_1,$$

where we have used (4.26). In the case of the D6-brane, although we have to reformulate the kinetic term in terms of the dual field strength F_8 , the conclusion is again similar.

The phenomenon leading to the charge conservation in the cases above is actually related to brane creation processes introduced in [64] and further discussed in [145, 147]. This can be seen easily as follows. Suppose there is a D8-brane which is such that we have $m = 0$ on its left and $m = \mu_8$ on its right. If we start with a configuration with an isolated D0-brane on the left, and we try to pass it through the D8-brane, we observe that a fundamental string stretching from the D8 to the D0-brane has to be created for the configuration with the D0-brane on the right to be stable. This actually happens because the D0-brane induces a charge on the world-volume of the D8-brane, and the induced charge changes sign when the D0-brane goes from one side to the other. The relevant equation on the world-volume of the D8-brane reads:

$$d \star \mathcal{F}_2 = \hat{\mu}_0 \hat{\delta}_8 - * \hat{F}'_2{}^{(9)},$$

and we see that the change in sign of the second term in the r.h.s. can be compensated by the introduction of a point like source, i.e. the end of a fundamental string.

More generally, the r.h.s. of the world-volume equations collected in the Tables 4.1–4.3 gives rise to an induced charge on the ‘host’ brane only if the brane producing the charge has a well-defined orientation in space-time. The configuration of these two branes (like the D0 and the D8-brane in the example above) can always be reconducted by dualities to a configuration of D-branes with $\nu = 8$ Neumann-Dirichlet directions. The common characteristic of all these configurations (which coincide with the ones discussed at the classical level in [134]) is that the two branes cannot avoid each other in space-time if one tries to move one past the other. A linking number of one brane with respect to the other can thus be defined [64].

4.5 Conclusive speculations

We have shown in this chapter that all the configurations with a closed p_a -brane intersecting a closed p_b -brane on a $p_a - 1$ dimensional intersection with vanishing binding

energy can be opened. Charge conservation and gauge invariance are simultaneously ensured by Goldstone field strengths appearing on the world-volume of the host p_b -brane. Each such intersection is related to another one involving the same host p_b -brane, and such that the respective boundaries of the open branes carry dual charges. Note that the field strengths in the branes have been generated from supergravity without any appeal to the apparatus of string theories.

We have seen that all D-branes carry fields \mathcal{F}_2 related to each other through dimensional reduction along the world-volume of the brane. The reason behind this universality is that all D-branes can receive charges from string boundaries. But this is not the case for the solitonic five-branes, who can host only ‘dual’ charges stemming from D-brane boundaries.

A particular case is that where a fundamental string opens on a D-brane. One can then consider the reverse process where open strings collapse to form closed strings with zero binding energy to the D-brane. Energy conservation permits the closed strings to separate from the brane. This process can be viewed as the classical limit of the well known quantum process of closed string emission by D-branes [62, 63]. Our analysis shows, at the classical level, that this string process can be extended to the emission of closed branes. Any host brane supports traveling open branes attached to it. These configurations of open branes can collapse through merging of the two boundaries into configurations containing closed branes with zero binding energy. This leads to the emission of supersymmetric closed branes.

The generality of this phenomenon points towards the existence of an underlying quantum theory where the emission of closed branes and that of closed strings occur as similar phenomena. This should in fact be M-theory. However, to describe in a more detailed way the emission of higher branes, a quantum theory of the latter is nevertheless still lacking (see however [148] for a proposal on open membranes in Matrix theory).

Chapter 5

Little theories on the world-volume of branes

The branes of M-theory are interesting objects not only for the rôle they play in 11 or 10 dimensional physics, but also because they are objects endowed with non-trivial dynamics on their world-volume. As we have seen in the preceding chapter, the branes have well-defined interactions with other branes, and these interactions result in the definition of world-volume charged objects and of world-volume effective theories. These world-volume theories, which can be ordinary field theories or, as we will see in this chapter, more complicated theories of extended objects, are especially interesting if they can be studied on their own. For this to be possible one generically tries to define a limit in M or in type II theories in which (most of) the bulk physics decouples, i.e. there are no interactions between the world-volume of the brane and the space-time in which it is embedded. Note that an important feature common to all the world-volume theories is that they do not contain gravity.

In this chapter we focus on theories in 6 and 7 dimensions, and which are moreover supersymmetric (with 16 or 8 supercharges). These are in some sense the highest dimensional non-trivial theories, and all the lower dimensional ones can in principle be recovered as particular limits.

The chapter is a slightly extended version of the paper [4].

5.1 Introduction

With the so-called Second Superstring Revolution, branes came to play an important rôle, as reviewed in Chapter 2. Accordingly, the attention on their world-volume dynamics grew up. It soon became apparent that, in addition to the conventional ‘embedding’ Nambu-Goto-like action, most of them have a far richer dynamics, differing in this respect from the fundamental type II superstrings and the 11 dimensional supermembrane. Namely, using a Goldstone mode analysis (see also Chapter 4) it appears that the solitonic 5-branes of type II theories, or NS5-branes, have world-volume low-energy effective actions given by two different 6 dimensional supersymmetric field theories, one chiral and

containing a self-dual tensor field¹, the other non-chiral and with a vector field, in addition to the usual embedding scalars [57]. We will discuss extensively these theories in this chapter.

In the following years, after the identification of the p -branes carrying Ramond-Ramond charge with the D-branes [48], it became clear that the world-volume low energy theory of these objects was simply given by the dimensional reduction to $p+1$ dimensions of 10 dimensional super Yang-Mills (SYM) theory [55]. The possibility to use in this case string perturbation theory techniques allowed actually to identify the effective theory of N coinciding branes, giving a SYM based on the group $U(N)$, while the Goldstone analysis could at most give the $U(1)$ ‘center of mass’ part.

By dualities, the effective theory of N coinciding M5-branes (and NS5-branes of type IIA), which are individually theories of self-dual tensors with no non-abelian generalization, was postulated by Strominger [99] to involve tensionless strings, arising actually from open M2-branes stretched between the M5-branes. The theory so obtained was closely related to the theory discussed in [150]. It has now to be stressed that in order to consider this and the other world-volume theories on their own, one had generically to decouple the bulk effects by taking some limits in the 10 or 11 dimensional theory. However taking too much of a strong limit can result in the omission of some higher energy effects which are nevertheless still confined to the world-volume of the brane.

A pioneering work on non-critical string theories living in 6 dimensions appeared in [151]. These theories will be called ‘little string theories’ because they have presumably much less degrees of freedom than the 10 dimensional, critical ones. In [152], Seiberg defined two sorts of little string theories as the theories which are left on the world-volume of the NS5-branes of type IIA and IIB theories when the string coupling is taken to vanish but the string scale (the string length l_s for instance) is kept fixed. In this limit it was argued that the bulk physics decouples, but nevertheless at energies of order m_s there is non-trivial ‘stringy’ dynamics in 6 dimensions, coming from the fundamental strings ‘trapped’ to the world-volume of the branes. The two little string theories were shown to be related by an analog of T-duality. They differ by their low-energy field theory limit which is defined by the one of the respective NS5-brane.

In the paper of Losev, Moore and Shatashvili [153], it was pointed out that adding transverse compact directions to the 5-branes, one could define in some limits little string theories with an additional dimensionless parameter, taken to be the coupling of the little string theory. Moreover, these theories could now be mapped by dualities to a 7 dimensional theory, with finite-energy membrane excitations, called m -theory. This 7 dimensional little theory was also discussed, in a more Matrix theory-oriented framework, in [154, 155].

In this chapter, we study in more details the little string theories and m -theory. We first revisit the theories in 6 and 7 dimensions with 16 supercharges, which is the amount of supersymmetry preserved by a BPS brane in a maximally supersymmetric theory like

¹ Strictly speaking, the theory should be defined by its equations of motion. Because of the self-duality condition, it is difficult to write a covariant action for this tensor field. See the recent progress on this issue in [149]. In the following, the use of the word ‘action’ should be considered merely as a shorthand.

M-theory or type II string theories. These theories lead to the definition of *ia*, *ib* little string theories and *m*-theory. The names beared by these theories will be justified by analogies with the 10 and 11 dimensional theories with capital letters. The different ways to obtain these little theories are analyzed. We start from 5 and 6 dimensional extended objects defined in M or type II theories and we take limits in which bulk modes decouple. This leads nevertheless to a non-trivial theory without gravity defined on the world-volume of the extended objects. We show the web of dualities between these little theories which exactly reproduces the scheme of the “big” theories in 10 and 11 dimensions. The spectrum of the BPS extended objects of these little theories is investigated and it is shown that it agrees with the U-duality group of M-theory compactified on T^6 . This will turn out to be of key importance in the application to the Matrix description of M-theory compactifications on higher dimensional tori. Also, for this application to be possible, the little theories as introduced above must be defined generically taking N parallel ‘host’ branes, although this feature will not enter into the precise definition of the little theories.

We will then turn to theories in 6 dimensions with 8 supercharges. These theories have (1,0) supersymmetry (i.e. the minimum in 6 dimensions), do not contain gravity and may have an additional integral parameter leading possibly to a gauge symmetry. Our strategy is to obtain them from the theories with 16 supercharges. We mimic the 10 dimensional procedure in which type I theory is obtained from IIB theory introducing an $\Omega 9$ orientifold and 16 D9-branes [51, 48, 26] (see also [53]). The two heterotic string theories are then found by chains of dualities. This procedure is reviewed in Chapter 2. Applying the same procedure to the 6 dimensional theories, we find one theory with open strings and two with closed strings, which we call respectively type *i*, h_b and h_a theories. These are in fact classes of theories. Unlike the 10 dimensional case, the group structure is not totally constrained, but also more subtle to define, especially for the h_a “little heterotic” theories. As a consistency check of the picture, the h_a theory can also be related to a particular compactification of *m*-theory.

The chapter is organized as follows. In Sections 5.2 and 5.3 we study respectively the theories with 16 and with 8 supercharges. We emphasize the web of dualities which relates all these theories, and its similarity with the dualities constituting M-theory. We also comment on the low-energy effective theories and on the decoupling of the little theories from the bulk. In Section 5.4 we apply the little theories to the Matrix theory description of M-theory on T^6 . We conclude in Section 5.5 with a discussion, mainly on the attempt to formulate the little theories using another Matrix model.

5.2 Theories with 16 supercharges

Supersymmetric theories with 16 supercharges naturally appear in type II string theories and M-theory as the effective theory on the world-volume of BPS branes. In order to have well-defined theories on world-volumes one has to take a limit in which the bulk modes decouple. This is achieved by sending the Planck mass, defined with respect to the non-compact space, to infinity.

We will consider here theories defined on the world-volume of 5 branes and 6 branes, and such that at least three of the transverse directions are non-compact (in order to keep the space asymptotically flat and thus the number N of parallel branes arbitrary). This allows for extra transverse compact directions, which will actually play a key rôle in defining the parameters of the little theories. Note that we could also consider lower dimensional branes, but they will in general have more transverse compact directions, eventually leading to a complicated theory which is nothing else than a toroidal compactification of one of the theories discussed below. The SYM theories are recovered as low energy limits of these more complicated theories.

In M-theory, we have the following two objects:

- M5-brane, with up to 2 transverse compact directions parametrized by R_1 and R_2 .
- KK6-brane, which has naturally a transverse compact direction, the so-called NUT direction (see [156] for a recent review on KK monopoles, and also Subsection 3.2.5).

In type IIA theory we have the following three objects:

- NS5(A)-brane, with 1 transverse compact direction parametrized by its radius R_A .
- KK5(A)-brane, with its transverse NUT compact direction.
- D6-brane, with no compact transverse directions.

The objects we have in type IIB theory are:

- NS5(B)-brane, with 1 transverse compact direction.
- KK5(B)-brane, with its NUT compact direction.
- D5-brane, with 1 compact transverse direction.

Note that we will sometimes call the above branes the ‘host’ branes, since they will host the dynamics of the little theories.

All these branes are related by the usual dualities relating type II and M-theory. We will however distinguish between dualities which leave the world-volume of the branes unaffected, as transverse T-dualities for NS branes (NS5 and KK5), IIB S-duality and transverse compactifications, and dualities which on the other hand act on the world-volume, as T-dualities and compactifications along a world-volume direction of NS branes or KK monopoles and T-dualities for D-branes (see Appendix B).

Considering the dualities leaving the world-volume unaffected leads to three different families of branes each one defining one theory:

- *ia*: $\text{KK5(A)} \leftrightarrow \text{NS5(B)} \leftrightarrow \text{D5}$
- *ib*: $\text{KK5(B)} \leftrightarrow \text{NS5(A)} \leftrightarrow \text{M5}$
- *m*: $\text{KK6} \leftrightarrow \text{D6}$

These three theories are related by dualities which affect the world-volume of the branes. A T-duality along the world-volume of a NS5 or a KK5 changes from IIA to IIB and thus also from *ia* to *ib*. Compactification of the KK6 on one of its world-volume directions yields the KK5(A), thus relating *ia* and *m* theories via compactification. The same duality between little theories is obtained acting with a T-duality on the world-volume of the D6, which gives the D5. Note also that the D4-brane, which defines a theory in 5 dimensions, can be obtained either by a T-duality from the D5, or by compactification from the M5. This shows that once compactified, there is no longer difference between *ia* and *ib* theories in 5 dimensions.

Although the relations discussed above are rather formal at this stage, they exactly reproduce the same pattern of dualities of the 10 and 11 dimensional theories. We will show hereafter that in the proper limits in which the above little theories make sense (i.e. when they decouple from the bulk), this structure still holds and acquires even more evidence.

We now turn to the description of the different little theories.

5.2.1 *ia* theory

As explained above, there are three ways to define type *ia* theory [153]. The six dimensional supersymmetry is $(1, 1)$. This is most easily found for the D5 brane from dimensional reduction of the $N = 1$ supersymmetry in $D = 10$ [26] (see also Section 2.5). For the NS5(B) and the KK5(A) it has been discussed respectively in [57] and [156]. The type *ia* theory is thus non-chiral.

The first approach is based on the D5 with a transverse compact direction of radius R_B .

We look for all the objects which from the D5 world-volume point of view have a finite tension, i.e. we rule out branes extending in transverse non-compact directions. The relevant configurations of branes intersecting with the D5, and breaking further 1/2 of the supersymmetry are: $D1 \subset D5$, $F1 \mapsto D5$, $D3 \mapsto D5$, $NS5(B) \mapsto D5$ and $KK5(B) \parallel D5$. The F1, D3 and NS5 can have a boundary on the D5, as discussed in Chapter 4, and their only dimension transverse to it wraps around the transverse compact direction.

Generically, supergravity solutions preserving 1/4 of the supersymmetries and representing two intersecting branes can be computed, following for instance the computations performed in Section 3.3 (see e.g. [101, 102, 119, 1]). Their existence can be deduced by the compatibility of the two supersymmetry projections which characterize the configuration. The supersymmetry projections characterizing the branes are discussed in Appendix B, where the notations are also recalled.

Before taking the limit in which the bulk decouples, we have to fix the tension and the coupling of the little string theory on the world-volume of the D5-brane. Since we have three parameters at hand, namely the string length l_s , the string coupling of IIB theory g_B and the radius R_B , it will be possible to send the 9 dimensional Planck mass²

² We have to consider the Planck mass in 9 dimensions because one of the transverse directions is compact. Furthermore its radius will be sent to zero in the limit discussed above. Note also that this limit does not depend on the size of the directions longitudinal to the D5-brane. For simplicity, we take

to infinity while keeping a non-trivial little theory on the brane characterized by two parameters, one of which dimensionless.

The only string-like object which lives on the D5-brane is the D1 string trapped to its world-volume [73]. We take it to define the fundamental little string of *ii*a theory. Accordingly, its tension is defined by (using (B.10) and neglecting here and in the following all numerical factors):

$$t_a \equiv T_{D1} = \frac{1}{g_B l_s^2}. \quad (5.1)$$

The boundaries of the F1, D3 and NS5, which are respectively 0-, 2- and 4-dimensional closed objects, act as little “d-branes” for the *f*1 little string. Their tension is postulated to be inversely proportional to the coupling of the little string theory g_a [153]. We have:

$$\begin{aligned} t_{d0} &\equiv T_{F1} R_B = \frac{R_B}{l_s^2} \equiv \frac{t_a^{1/2}}{g_a} \\ t_{d2} &\equiv T_{D3} R_B = \frac{R_B}{g_B l_s^4} \equiv \frac{t_a^{3/2}}{g_a} \\ t_{d4} &\equiv T_{NS5} R_B = \frac{R_B}{g_B^2 l_s^6} \equiv \frac{t_a^{5/2}}{g_a} \end{aligned} \quad (5.2)$$

The above definitions are consistent and, taking into account (5.1) we have:

$$g_a = \frac{l_s}{g_B^{1/2} R_B}. \quad (5.3)$$

The last object to consider is the KK5, which actually fills the world-volume of the D5. We can nevertheless define its tension using (B.9):

$$t_{s5} \equiv T_{KK5} = \frac{R_B^2}{g_B^2 l_s^8} = \frac{t_a^3}{g_a^2}. \quad (5.4)$$

The *d*4 and *s*5 branes were overlooked in the analysis of [153], they are however defined by perfectly well-behaved 10 dimensional configurations. They are important in the identification of this little theory as a model for a toroidal compactification of Matrix theory as we discuss at the end of this chapter.

It is worth pointing out that we were able to define the tensions of five little branes using only two combinations of the parameters g_B , l_s and R_B . This is really the crucial point which allows us to formulate the little theories.

We have now defined the string tension t_a and the string coupling g_a of the little theory. In order for this *ii*a theory to make sense, we have to take a limit in which the bulk modes decouple i.e. a limit in which the nine dimensional Planck Mass M_p is going to infinity at fixed t_a and g_a . The Planck Mass is given by:

$$M_p^7 = \frac{R_B}{g_B^2 l_s^8} = \frac{g_B t_a^{7/2}}{g_a}. \quad (5.5)$$

The limit defining type *ii*a is thus characterized by:

$$g_B \rightarrow \infty, \quad l_s \rightarrow 0, \quad R_B \rightarrow 0 \quad (5.6)$$

them to be infinite.

We can also find the *ia* theory starting with the NS5(B)-brane with one transverse compact direction of radius \tilde{R}_B . We call, in this case, the string coupling of type IIB \tilde{g}_B and the string length \tilde{l}_s . The 10 dimensional configurations breaking 1/4 supersymmetry which define the BPS objects living in 6 dimensions are simply obtained by S-duality from the ones discussed in the preceding approach. They are the following: $F1 \subset NS5(B)$, $D1 \rightarrow NS5(B)$, $D3 \rightarrow NS5(B)$, $D5 \rightarrow NS5(B)$ and $KK5(B) \parallel NS5(B)$. The little *ia* string is identified to the fundamental string of type IIB theory. Its tension is simply given by: $t_a = \tilde{l}_s^{-2}$. The little string coupling is:

$$g_a = \frac{\tilde{g}_B \tilde{l}_s}{\tilde{R}_B}. \quad (5.7)$$

It can be obtained for instance computing $t_{d0} = T_{D1} \tilde{R}_B$. The limit in which the Planck mass goes to infinity is defined by $\tilde{g}_B \rightarrow 0$, $\tilde{R}_B \rightarrow 0$ and \tilde{l}_s constant. This result is consistent with the S-duality transformations: $g_B \rightarrow \tilde{g}_B = 1/g_B$, $l_s^2 \rightarrow \tilde{l}_s^2 = g_B l_s^2$ and $\tilde{R}_B = R_B$ left unchanged.

This picture is maybe the heuristically more appealing, since the little strings are identified to fundamental type IIB strings trapped inside a NS5-brane, at vanishing type IIB coupling. Note also that when we decompactify the transverse direction, i.e. we take $\tilde{R}_B \rightarrow \infty$ instead of $\tilde{R}_B \rightarrow 0$, we obtain the zero coupling limit of the little type *ia* string theory, which thus coincides with the definition of one of the two little string theories of Seiberg [152]. Indeed, all the BPS states but the *f1* acquire an infinite tension and decouple.

The third object with a 6 dimensional (1,1) supersymmetric world-volume theory which can be used to define *ia* theory is the KK5 monopole of type IIA string theory, obtained by a T-duality on the transverse compact direction from the NS5(B)-brane. This direction becomes the NUT direction of the Euclidean Taub-NUT space transverse to the KK5 world-volume [156]. It appears that in this picture all the relevant configurations which preserve 1/4 supersymmetries are made up from branes of type IIA inside the world-volume of the KK5(A) [153]: $F1 \subset KK5(A)$, $D0 \subset KK5(A)$, $D2 \subset KK5(A)$, $D4 \subset KK5(A)$ and $NS5(A) \subset KK5(A)$. This makes the identification of t_a and g_a straightforward. The fundamental *ia* string coincides now with type IIA's F1, and thus $t_a = \tilde{l}_s^{-2}$. Since here also the "little" *d*-branes coincide with the D-branes of type IIA (with $p \leq 4$), also the little string coupling is given by the IIA one: $g_a = g_A$. It is easy to find by T-duality from the NS5(B) picture the limit in which the KK5 decouples from the bulk. Since under T-duality $g_B \rightarrow g_A = \frac{g_B \tilde{l}_s}{R_B}$, $R_B \rightarrow R_A = \frac{\tilde{l}_s^2}{R_B}$ and \tilde{l}_s is unchanged, in the KK5(A) picture we have g_A constant and $R_A \equiv R_{NUT} \rightarrow \infty$. The Riemann tensor of the Taub-NUT geometry vanishes in this limit, an indication that the KK monopole decouples from bulk physics (we comment on the decoupling issue at the end of the section).

We recapitulate the BPS spectrum of type *ia* theory in Table 5.1. We list the different little branes and their mass considering now a compact world-volume characterized by radii Σ_i with $i = 1 \dots 5$ and volume $\tilde{V}_5 = \Sigma_1 \dots \Sigma_5$. We include for later convenience the KK momenta w .

Brane	Mass
w	$\frac{1}{\Sigma_i}$
$f1$	$\Sigma_i t_a$
$d0$	$\frac{t_a^{1/2}}{g_a}$
$d2$	$\frac{\Sigma_i \Sigma_j t_a^{3/2}}{g_a}$
$d4$	$\frac{\tilde{V}_5 t_a^{5/2}}{\Sigma_i g_a}$
$s5$	$\frac{\tilde{V}_5 t_a^3}{g_a^2}$

Table 5.1: Mass of the BPS objects in *iaa* theory.

We also summarize in Table 5.2 the different ways to obtain *iaa* theory and the relation between the parameters.

To summarize, we have defined type *iaa* little string theory as a 6 dimensional string theory characterized by a tension t_a and a coupling g_a , and by a (1, 1) non-chiral supersymmetry. The low energy effective action of the little string theory is identified with the low-energy world-volume action of the brane on which it is defined, in this case 5+1 dimensional SYM. As anticipated, it does not contain gravity. If there are N host branes, the theory is based on the group $U(N)$. We collect at the end of this section the remarks on the low-energy limits of the little theories.

5.2.2 *iib* theory

We recall that there are three approaches to this 6 dimensional theory, using respectively the M5-brane with two transverse compact directions, the NS5-brane of type IIA with one compact transverse direction and the KK5 monopole of type IIB [153]. These three different branes all have a world-volume theory with (2,0) chiral supersymmetry [57, 150, 99, 156].

The procedure by which we analyze the structure of *iib* little string theory is similar to the one described in the preceding subsection. We will however meet here an interesting structure of *iib* which is its *s*-duality. We begin with the M5 approach, where this duality is geometric.

The M5-brane set up is characterized by the 11 dimensional Planck length L_p and the two radii R_1 and R_2 of the two transverse compact directions. The configurations breaking 1/4 supersymmetry in M-theory leading to finite tension objects on the world-volume of the M5 are the following: M2→M5 with the M2 direction orthogonal to the

	$D5$ $\frac{1}{g_B}, R_B, l_s \rightarrow 0$	$NS5(B)$ $\tilde{g}_B, \tilde{R}_B \rightarrow 0$	$KK5(A)$ $R_{NUT} \rightarrow \infty$
t_a	$\frac{1}{g_B l_s^2}$	$\frac{1}{\tilde{l}_s^2}$	$\frac{1}{\tilde{l}_s^2}$
g_a	$\frac{l_s}{g_B^{1/2} R_B}$	$\frac{\tilde{g}_B \tilde{l}_s}{\tilde{R}_B}$	g_A

Table 5.2: Definitions of *ia* parameters.

M5 wrapping either R_1 or R_2 ; $M5 \cap M5 = 3$; $KK6 \supset M5$ with the NUT direction of the KK6 identified either with R_1 or R_2 .

The boundaries of the M2-branes are strings on the M5, but we cannot immediately identify the fundamental *ib* little string because we have two different kinds of them. We simply choose one of the two (say, the boundary of the M2 wrapped on R_1) to be the fundamental and thus to have tension t_b , and the other to be the little *d1* brane with tension $\frac{t_b}{g_b}$. This defines g_b . *s*-duality of *ib* is then simply the interchange in M-theory of R_1 and R_2 (this can actually be extended to a full $SL(2, Z)$ duality group considering M2-branes wrapped on (p, q) cycles of the torus). We have thus (cfr. (B.4)):

$$\begin{aligned} t_{f1} &= T_{M2} R_1 = \frac{R_1}{L_p^3} \equiv t_b, \\ t_{d1} &= T_{M2} R_2 = \frac{R_2}{L_p^3} \equiv \frac{t_b}{g_b}. \end{aligned} \quad (5.8)$$

The little string coupling is then given by:

$$g_b = \frac{R_1}{R_2}. \quad (5.9)$$

We can now identify the other world-volume objects by their tension:

$$\begin{aligned} T_{M5} R_1 R_2 &= \frac{R_1 R_2}{L_p^6} = \frac{t_b^2}{g_b} \equiv t_{d3} \\ T_{KK6} R_2 &= \frac{R_1^2 R_2}{L_p^9} = \frac{t_b^3}{g_b} \equiv t_{d5} \\ T_{KK6} R_1 &= \frac{R_1 R_2^2}{L_p^9} = \frac{t_b^3}{g_b^2} \equiv t_{s5} \end{aligned} \quad (5.10)$$

Note that under *s*-duality the *d3* is inert and the *d5* and *s5* are exchanged.

We still have to find the limit in which the bulk physics decouples. Keeping t_b and g_b finite, the Planck mass in 9 dimensions is given by:

$$M_p^7 = \frac{R_1 R_2}{L_p^9} = \left(\frac{t_b^2}{g_b} \right) \frac{1}{L_p^3},$$

and goes to infinity when $L_p \rightarrow 0$. To keep the parameters of *iib* finite, we also have to take $R_1, R_2 \rightarrow 0$.

We now consider the NS5(A) approach. The parameters are the string length \tilde{l}_s , the string coupling \tilde{g}_A of type IIA theory and the radius \tilde{R}_A of the compact direction. The configurations, breaking 1/4 supersymmetry, leading to finite tension objects in the world-volume of the NS5(A) are: $F1 \cap NS5(A)$, $D2 \mapsto NS5(A)$, $D4 \mapsto NS(A)$, $D6 \mapsto NS(5)$ and $KK5(A) \parallel NS5(A)$. In this framework the little string tension t_b is defined by the fundamental string F1, namely $t_b = \tilde{l}_s^{-2}$. The little string coupling g_b is found by identifying the tension of the $d1$ -brane from the configuration with the D2. We have:

$$t_{d1} = T_{D2} \tilde{R}_A = \frac{\tilde{R}_A}{\tilde{g}_A \tilde{l}_s^3} \equiv \frac{t_b}{g_b}, \quad g_b = \frac{\tilde{g}_A \tilde{l}_s}{\tilde{R}_A} \quad (5.11)$$

We obtain this picture from the previous one by dimensional reduction on R_1 , $R_1 = \tilde{g}_A \tilde{l}_s$. The \tilde{R}_A here is the previous R_2 . In this case the limit is taken performing $\tilde{g}_A \rightarrow 0$ and $\tilde{R}_A \rightarrow 0$ at fixed t_b and g_b . Note that the s -duality in this picture is less straightforward to obtain from 10 dimensional string dualities (one has to operate a TST duality chain).

As for the type *iia* strings, the limit $\tilde{R}_A \rightarrow \infty$ of the picture above produces the decoupling of all the little ‘solitons’, leaving the second of the theories discussed by Seiberg [152]. Since in this case $g_b = 0$, s -duality ceases to be a symmetry of the theory (in the same way as there is no S-duality in perturbative type IIB string theory).

Turning now to the KK5(B) picture, we find that, as in the type *iia* case, the little string theory is the reduction to the world-volume of the KK5 of the physics of the objects that fit inside it. Thus we simply identify t_b with \tilde{l}_s^{-2} , g_b with g_B , s -duality with S-duality, $f1$ with F1 and so on. As in the previous KK5 case, the limit in which the bulk decouples involves taking the radius of the NUT direction to infinity.

We recapitulate the BPS spectrum of type *iib* theory in Table 5.3. As for the *iia* case, we list the different little branes and their mass considering now a compact world-volume characterized by radii Σ_i with $i = 1 \dots 5$ and volume $\tilde{V}_5 = \Sigma_1 \dots \Sigma_5$.

We also summarize in Table 5.4 the different ways to obtain *iib* theory and the relation between the parameters.

To recollect, type *iib* little string theory is also a non-critical 6 dimensional string theory characterized by a tension t_b and a coupling g_b , and with chiral (2,0) supersymmetry. It also displays a symmetry, s -duality, which acts on its BPS spectrum and which exchanges g_b with $1/g_b$. The low-energy effective theory is defined by the self-dual tensor multiplet. When there are N host branes defining the theory, we have N such multiplets, but the non-abelian generalization is still lacking, contrary to the former *iia* case. We will discuss this further shortly.

There is a t -duality relating *iia* and *iib* little string theories, as most easily seen in the pictures using the NS5 or the KK5 branes. It is simply the 10 dimensional T-duality between IIA and IIB, applied on a direction longitudinal to the world-volume of the above-mentioned branes. To be more specific, application of such a longitudinal T-duality maps, say, the NS5(A) picture of *iib* theory to the NS5(B) picture of *iia* theory, and similarly for the KK5 pictures. The behaviour of the BPS objects is the same as in

Brane	Mass
w	$\frac{1}{\Sigma_i}$
$f1$	$\Sigma_i t_b$
$d1$	$\frac{\Sigma_i t_b}{g_b}$
$d3$	$\frac{\tilde{V}_5 t_b^2}{\Sigma_i \Sigma_j g_b}$
$d5$	$\frac{\tilde{V}_5 t_b^3}{g_b}$
$s5$	$\frac{\tilde{V}_5 t_b^3}{g_b^2}$

Table 5.3: Mass of the BPS objects in iib theory.

type II string theories: KK momenta are exchanged with wound $f1$ strings (as in [152]), the $s5$ brane of one theory is mapped to the one of the other theory, and dp -branes become $d(p+1)$ - or $d(p-1)$ -branes for transverse or longitudinal t -dualities respectively. ii_a and iib theories are thus equivalent when reduced to 5 space-time dimensions or less.

5.2.3 m -theory

As stated at the beginning of this section, there are two objects with 7 dimensional world-volume in M/type II theories: the D6-brane in type IIA and the KK6 monopole in M-theory. The supersymmetry algebra is unique and obviously non-chiral.

	$M5$ $L_p, R_1, R_2 \rightarrow 0$	$NS5(A)$ $\tilde{g}_A, \tilde{R}_A \rightarrow 0$	$KK5(B)$ $R_{NUT} \rightarrow \infty$
t_b	$\frac{R_1}{L_p^3}$	$\frac{1}{l_s^2}$	$\frac{1}{l_s^2}$
g_b	$\frac{R_1}{R_2}$	$\frac{\tilde{g}_A \tilde{l}_s}{R_A}$	g_B

Table 5.4: Definitions of iib parameters.

We first consider the D6 approach. Note that for the transverse space to be asymptotically flat, we cannot have any compact transverse dimension. The free parameters are thus the string length l_s and type IIA string coupling g_A . Already at this stage we know that the theory on the world-volume will be characterized by only one parameter (one is lost taking the appropriate limit which decouples the bulk).

In this case, we have to consider configurations preserving 1/4 supersymmetries with a brane within the D6-brane. The only branes of type IIA for which this works are the D2- and the NS5-brane. We identify them with the $m2$ and $m5$ branes. As it is necessary for the definition of m -theory, only one parameter suffices to define both their tensions. Indeed we have:

$$\begin{aligned} t_{m2} &\equiv T_{D2} = \frac{1}{g_A l_s^3} \equiv \frac{1}{l_m^3} \\ t_{m5} &\equiv T_{NS5} = \frac{1}{g_A^2 l_s^6} = \frac{1}{l_m^6} \end{aligned} \tag{5.12}$$

l_m is thus the characteristic length of m -theory, the analog of the Planck length in M-theory.

In order to decouple gravity, we send the 10 dimensional Planck mass to infinity. Keeping l_m finite, we have:

$$M_p^8 = \frac{1}{g_A^2 l_s^8} = \frac{1}{(l_m^6) l_s^2},$$

and thus we have to take $l_s \rightarrow 0$ and $g_A \rightarrow \infty$.

In the KK6 approach, there are two configurations preserving 1/4 of supersymmetry: $M2 \subset KK6$ and $M5 \subset KK6$. M2 and M5 are thus respectively identified to $m2$ and $m5$, and $l_m = \tilde{L}_p$ where \tilde{L}_p is the eleven dimensional Planck length. The KK6 monopole can be seen as the M-theoretic origin (and thus the strong coupling limit) of the D6-brane. The radius of the NUT direction is thus given by $R_{NUT} = g_A l_s = g_A^{2/3} \tilde{L}_p$. Therefore, the limit above $g_A \rightarrow \infty$ becomes $R_{NUT} \rightarrow \infty$. Again, in this limit the geometry becomes that of flat space.

It is interesting to note that here as in the former cases of ii_a and ii_b theories, the KK monopole description is the more “economic” one, in the sense that one has to take only one limit. However, the other descriptions will be useful, for instance, to make contact with Matrix theory compactifications.

In Table 5.5 the masses of the different BPS objects of m -theory are listed. Again we consider a compact volume $\tilde{V}_6 = \Sigma_1 \dots \Sigma_6$.

The different ways to obtain m -theory are shown in Table 5.6, along with the relation between the parameters.

m -theory is thus a 7 dimensional theory in which the lowest dimensional object is a membrane³. Its low-energy effective field theory limit is taken to be defined by the world-volume low-energy field theory of the D6-brane, that is 6+1 dimensional SYM, with $U(N)$ group if there are N D6-branes.

The duality between m -theory and ii_a theory can now be made more precise. The relations between the parameters of m -theory compactified on the “7th” direction of

³As observed in [154, 155], 7 is one of the dimensions in which a supermembrane can be defined [157].

Brane	Mass
w	$\frac{1}{\Sigma_i}$
$m2$	$\frac{\Sigma_i \Sigma_j}{l_m^3}$
$m5$	$\frac{\tilde{V}_6}{\Sigma_i l_m^6}$

Table 5.5: Mass of the BPS objects in m -theory.

	$D6$ $\frac{1}{g_A}, l_s \rightarrow 0$	$KK6$ $R_{NUT} \rightarrow \infty$
l_m	$g_A^{\frac{1}{3}} l_s$	\tilde{L}_p

Table 5.6: Definitions of m -theory parameters.

radius R_c and ii a theory are easily found comparing the tensions of the wrapped and unwrapped $m2$ brane on one side, and of the $f1$ and $d2$ branes on the other side. One finds no surprises:

$$t_a = \frac{R_c}{l_m^3}, \quad g_a = \left(\frac{R_c}{l_m} \right)^{3/2}.$$

In the $KK5(A)$ and $KK6$ picture, this is a direct consequence of the similar relations between M and IIA theories. It is more amusing to see that they indeed correspond to T-duality relations between IIA and IIB when one goes to the $D5/D6$ picture.

We have thus completed the pattern of dualities existing between the little theories with 16 supercharges: m -theory is related to ii a theory by compactification, ii a and iib theories are related by t -duality, and finally iib theory has an s -duality. Accordingly, these three theories can be considered as three different ‘phases’, or corners in the moduli space, of the same theory. This is very reminiscent of the dualities defining M-theory, as they are reviewed in Chapter 2. The analogy will be further developed by the introduction of the little theories with 8 supercharges.

We can actually guess the maximal duality group of m -theory (or of any one of the type ii theories) when compactified down to, say, $p+1$ dimensions, with now $p < 5$. Since the branes of m -theory are actually the same (in number and in extension) as the ones of M-theory which can be accommodated on a 6-dimensional torus T^6 , both fit into the same duality multiplets when a compactification on T^{6-p} is considered. It is tempting

to speculate that like M-theory on T^6 , the fully compactified m -theory has a u -duality based on the group $E_6(Z)$, as discussed in [153] (see also [158]). The minimal evidence of this duality, i.e. that the BPS states fit into representations of $E_6(Z)$, will be important in providing the link between these little theories and Matrix theory compactified on T^6 .

5.2.4 Some comments on the low-energy effective actions of the little theories

Unlike 10 dimensional critical string theories, in the case of the little string theories it is impossible at the present state of the knowledge to directly compute their massless spectrum and the interactions, and then formulate the low-energy effective action that governs them. We have however another way to precisely guess these low-energy effective actions, even if this way is somewhat indirect and relies on the fact that the definition of the little theories uses some host brane. As anticipated when reviewing the little theories above, we assume that the low-energy effective action of the little theories coincides with the world-volume low-energy effective action of the host branes. It is actually a matter of consistency that the two low-energy limits coincide, the guess being that the low-energy limits of the little theories exist.

Let us begin reviewing the low-energy effective action of type iib little string theory. If we take the M5 picture to define iib strings, then we know that the low-energy theory is given by the $(2, 0)$ multiplet containing a self-dual tensor H^+ and 5 scalars ϕ_i (and the fermions which we do not consider here and in the following). Two of the scalars, say ϕ_1 and ϕ_2 , have a finite range since they correspond to the compact directions, while the three others have an infinite range. Due to the self-duality of the tensor, this theory is believed not to have a coupling at all [159] (it is fixed to $g = 1$). Accordingly all the fields have mass dimension 2 and the range of the scalars can be given by $0 \leq \phi_{1,2} \leq R_{1,2}/L_p^3$, as in [160]. In type iib variables, the ranges can be rewritten as:

$$0 \leq \phi_1 \leq t_b, \quad 0 \leq \phi_2 \leq \frac{t_b}{g_b}. \quad (5.13)$$

We already see at this stage that unlike ‘conventional’ string theories, here the little string coupling appears in the low-energy action in a non-trivial way. Note that when considering a NS5(A)-brane instead of a M5, the asymmetry⁴ between ϕ_1 and ϕ_2 can be ascribed to the fact that the former comes from the RR sector while the second from the NSNS one [160].

We can now ask ourselves the question of how we recover in the low-energy theory the BPS objects as solitons, in the same way as each brane of M-theory corresponds to a supergravity solution. Naively one could think that the 3-form H^+ naturally couples to a (self-dual) string, but then we run into problems because this would give rise to only one string soliton instead of the required two, according to Table 5.3. The resolution of this problem can be found in [161] (see also [162]). For a solution to be BPS, it cannot

⁴ Note that in [160] the fact that the range of ϕ_2 is proportional to R_A/g_A was taken as an indication that the direction effectively decompactifies when $g_A \rightarrow 0$. Here however the little string limit is such that the range is precisely kept finite.

carry only one charge, but there should be a balance of two forces that cancel in presence of two parallel BPS objects. The second field which carries a charge in this case is either ϕ_1 or ϕ_2 [161]. The construction of [161] is however not enough for our purpose, since in our case the same M2-brane ‘starts’ and ends on the same M5. One might be afraid of having an unstable situation with opposite charges, but this does not occur [162]: when both the field H^+ and the scalar have the opposite sign, the state is still BPS and the energies add, with no dependence on the distance between the two objects. Moreover the world-volume energy is proportional to the length of the intersecting brane [162] and thus in this case it should presumably be proportional to the range of the scalar. This is also why we restricted the solitons to have as non-trivial scalar fields only the ones with finite range. Taking into account (5.13), it is straightforward to identify the $f1$ solution with the soliton charged under ϕ_1 and the $d1$ to be the one charged under ϕ_2 . Note that in the NS5(A) picture, ϕ_1 is the scalar which does not correspond to an embedding coordinate. Let us stress that although we did not carry out any calculation to support all the above discussion, we find it likely to be correct.

The s -duality of iib theory then reflects in the low-energy action in the permutation of ϕ_1 and ϕ_2 , and it is actually a subgroup of the R-symmetry (which is $SO(5)$ [159]). The $d3$ -brane soliton can be found in a similar way [163]. This solution is now charged under the two scalars ϕ_1 and ϕ_2 at the same time, while the 3-form H^+ is taken to vanish. The classical solution is thus also clearly inert under (classical) s -duality.

Concerning the little 5-branes, the $d5$ and the $s5$, we do not expect them to correspond to classical solitons since they would have to couple to non-dynamical 6-form potentials. They have a similar status as the D9-brane in type IIB string theory.

The last remark about the $(2, 0)$ low-energy effective theory is that when there are N coinciding M5-branes, the theory is conjectured to possess tensionless strings [150, 99] which, upon dimensional reduction to 4+1 dimensions, yield the massless gauge bosons of a non-abelian $U(N)$ SYM theory. These tensionless strings arise from M2-branes stretched between two adjacent M5-branes, and should not be confused with the little string solitons, which have a finite tension since they originate from M2-branes wrapping on a transverse compact direction.

Going now to the definition of the low-energy effective action of iaa theory, we have already observed that it should coincide with the 5+1 SYM effective theory of the D5-brane. We can go a little further: in the case of the D5-brane, the SYM coupling is given by $g_{YM}^2 = g_B l_s^2$. Translating to iaa variables as given in Table 5.2, this gives for the coupling of the iaa low-energy action $g_{YM}^2 = 1/t_a$. One of the 4 scalars of the SYM multiplet, call it ϕ_B , represents a compact direction and its range is given by:

$$0 \leq \phi_B \leq \frac{R_B}{l_s^2} \equiv \frac{t_a^{1/2}}{g_a}. \quad (5.14)$$

Going to the NS5(B) picture, the $1/\tilde{g}_B$ behaviour in the range of ϕ_B derives from its NSNS origin [160]. As in the iib case, here also g_a does not appear in front of the action but rather in the definition of one of the fields.

The mapping between the iaa BPS objects (see Table 5.1) and classical solitons is as follows. The $f1$ corresponds to a SYM instanton and accordingly has a tension given by

$1/g_{YM}^2 = t_a$. The tension is independent of g_a since no scalar participates to the instanton solution. The $d0$ and the $d2$ are the electric and the magnetic charge with respect to the SYM vector field, and if they are to be BPS the scalar ϕ_B must balance their charge, as in [162]. If we have N D5-branes, the above little branes simply correspond to the massive gauge boson of broken gauge symmetry and to the 't Hooft-Polyakov monopole respectively, the mass of both of which depends on the expectation value (and thus here on the range) of a scalar field. Here the original $SO(4)$ R-symmetry is broken completely at the level of the solutions since the other scalars, being of infinite range, give rise to infinitely massive objects. The $d4$ brane should couple to the still conjectural 6-form field strength, which was also considered in the previous chapter to occur in the world-volume of the NS5(B)-brane.

The low-energy effective action of m -theory is defined by the world-volume low-energy effective theory of the D6-brane, which is 6+1 dimensional SYM. The SYM coupling is now given by $g_{YM}^2 = g_A l_s^3 \equiv l_m^3$ (see Table 5.6), while all of the three scalars belonging to the SYM multiplet have an infinite range. Accordingly, the usual classical electric and magnetic charged objects have here an infinite mass, and this is why there are no 0- and 3-branes in m -theory, contrary to what one could have expected from the low-energy point of view. Rather, the $m2$ brane is given by the instanton, with tension $1/g_{YM}^2 = 1/l_m^3$. It is less clear what is the SYM origin of the $m5$ brane, which has a tension going like $1/g_{YM}^4$. It could couple to a non-propagating 7-form field strength. Indeed, in [142] it was conjectured that a $(p+1)$ -form field strength existed on the world-volume of every Dp -brane.

Let us end this subsection observing that there is yet another way to find the BPS content of the little theories. It is simply based on the knowledge of the supersymmetry algebras with 16 supercharges in 6 and 7 dimensions, including the p -form 'central' charges [164]. This procedure is very similar to the one which allows to find all the BPS objects in M-theory [42, 53]. In the present case however one has to take into account that most of the automorphism group of the supersymmetry algebras is broken due to the presence of compact transverse directions which single out some of the (little) branes.

5.2.5 Some comments on the decoupling from the bulk

A potential problem in the definition of the little theories was pointed out in [69, 70] in a Matrix theory context. The argument, applied to the little theories, is basically the following: in the D6-brane definition of m -theory as summarized in Table 5.6, the mass of the D0-brane present in this type IIA theory vanishes. Indeed, in the m -theory limit, it is:

$$M_{D0} = \frac{1}{g_A l_s} = \frac{l_s^2}{l_m^3} \rightarrow 0.$$

This could be taken as a signal that there are (infinitely many) additional massless degrees of freedom, which moreover are free to propagate in the bulk. The m -theory limit should eventually represent a 6+1 dimensional theory coupled to the bulk through D0-branes.

The same problem can be reformulated for every little theory in any one of its formulations, simply acting with dualities on the D0–D6-brane system. In the pictures using the KK monopoles for instance, the problem is seen to arise since, as can be seen in Tables 5.2, 5.4 and 5.6, they are always embedded in theories where all the couplings (g_A , g_B or \tilde{L}_p) are finite.

We find however that the problem is subtler than it may seem at a first sight. Indeed there are several hints that all point towards the conclusion that the decoupling actually takes place. Going back to the D0–D6-brane system, it is well known that they do not form a supersymmetric bound state. Rather, the only supergravity solution carrying at the same time D0 and D6-brane charge is not a stable state (see Subsection 3.3.5), i.e. it has an excess of energy with respect to the masses of the separated D0 and D6-branes. This means that a D0-brane in the bulk will interact (at least in this way) with the D6-brane only if it is supplied some energy to do so. The same conclusion comes from the stringy computation of the interaction between D0- and D6-branes [113], where it is found that there is a repulsive force.

In the KK monopole picture of each little theory, the problem is reformulated as follows. In the limit $R_{NUT} \rightarrow \infty$, the geometry of a KK monopole approximates the one of an ALE space (see [165, 156]) with a singularity at the core if there are N coinciding KK monopoles. Moreover the ALE space appears effectively 4 dimensional since the NUT direction decompactifies. Since the string coupling, or the 11 dimensional Planck length, are finite, one has to check whether the bulk gravitons interact with the singularity (where the little theory should be located) or not. Heuristically, one can argue, following [153], that the gravitons propagating along the NUT direction feel, even if R_{NUT} is very large, a vanishingly small radius when they approach the core (see the metric of a KK monopole (3.116)). They are thus effectively very massive there and presumably do not penetrate the singularity, i.e. they decouple. Concerning the truly massless gravitons in the three remaining transverse directions, they see an effective 9 dimensional Planck mass which is proportional to $R_{NUT}^{1/7}$, like for instance in (5.5). It is thus very big and these gravitons also should decouple. These are certainly not definitive statements, and one should search for more quantitative evidence, but the above facts seem to point out that something subtle can happen which ensures the decoupling.

Related to the decoupling issue discussed above, there is also the problem of making sure that the little objects do not leave the host brane. In fact, all the configurations defining the little BPS objects are marginal bound states of the host brane with some other brane. There is thus no energy barrier which prevents the two branes to depart from each other. One way to address this problem is to consider, for instance in the case of D1⊂D5 or D2⊂D6 defining respectively the little *ia* string and the m_2 of m -theory, the low-energy effective theory of a Dp -brane inside a $D(p+4)$ -brane [73]. This was also considered in [153]. It appears that in this low-energy theory, the scalars belonging to the hypermultiplets (the matter multiplets, arising from the open strings going from the Dp -brane to the $D(p+4)$ -brane) describe the motion of the smaller brane inside the bigger, while the scalars belonging to the vector multiplet describe the motion of the smaller brane away from the bigger, in the bulk. In this terminology, the Coulomb branch of the Dp -brane theory describes motion in the bulk, while the Higgs branch describes motion

trapped to the world-volume of the $D(p+4)$ -brane. A trapped brane thus cannot escape to the bulk if the Higgs and Coulomb branches are decoupled, and this is what seems to happen in the little theory limit [153]. This is a tentative explanation (which also applies to the little string theories of Seiberg [152]), and we will also mention in the concluding section of this chapter a similar mechanism which seems to ensure the decoupling of the little theories from the bulk.

A different solution to this problem could be to take the NS5-brane picture to define the little string theories, and then to abandon the idea of having a transverse compact direction, the radius of which eventually goes to zero. Instead, one could take two (sets of) NS5-branes in flat space at a distance L , and then suspend the D-branes between them (as in a Hanany-Witten configuration [64]) in order to give rise to the little branes. The little type ii strings would still be given by type II strings trapped inside the NS5-branes. The limit defining the little string theory would then be the same, provided L replaces the rôle of R , except in the Planck mass that is now the 10 dimensional one. This approach has however some serious flaws: on one hand, there seems to be no direct way to recover the $s5$ branes in the little string theories (but this might not be a problem since we could conceivably discard them); on the other hand, and more seriously, there is no directly equivalent definition of m -theory. This is why we preferred in this chapter to define the little theories using a transverse compact direction, signaling however above the potential problems of this definition.

We should also mention the issue raised in [166] and which applies also to the little string theories of [152]. Stated very briefly, a NS5-brane has a non-vanishing Hawking temperature even in the extremal limit (see equation (3.96) and the discussion which follows it). Moreover, the dilaton, and thus the string coupling, explode near the location of the NS5-brane (see the equations (3.105) and (3.107)). Both of these facts would be an indication that the theory does not decouple from the bulk. However the background of an extremal NS5-brane (in the string frame) has also the peculiarity of having an infinite ‘throat’ which effectively disconnects the location of the brane from asymptotic infinity. The radiation which emanates from the brane seems thus not to reach the bulk of the flat spacetime surrounding the brane [166].

The issues discussed in this subsection should be an indication that the exact mechanism which allows us to consider the little theories as decoupled from the bulk clearly deserves a more detailed analysis. However, in our opinion it is not hopeless to think that the little theories treated in this chapter will eventually make sense.

5.3 Theories with 8 supercharges

We propose in this section to define the little string theories with (1,0) supersymmetry in 6 dimensions. Note that this is the highest dimension in which a theory with 8 supercharges can live. We construct the (1,0) theories by analogy with the 10 dimensional relation between $N = 1$ and $N = 2$ string theories.

In 10 dimensions, type I open string theory can be obtained from type IIB string theory [51, 48]. One adds to the IIB theory an $\Omega 9$ orientifold yielding SO open strings

[26], and then adds 16 D9-branes to have a vanishing net flux of D9 RR charge. This leads to an $N = 1$ supersymmetric theory with open strings carrying $SO(32)$ Chan-Paton factors. The two heterotic string theories are then obtained by dualities. The $SO(32)$ heterotic theory is found by S-duality from the type I (identifying the D1-brane in the latter to the fundamental heterotic string of the former [52]). The $E_8 \times E_8$ heterotic theory is obtained by T-duality from the $SO(32)$ one. The $E_8 \times E_8$ theory can also be derived from M-theory compactified on S^1/Z_2 [54].

Our strategy is the following: we define the theories with 8 supercharges using the 5-branes of the type ii little string theories, and we then show that the same pattern of dualities as in 10 dimensions arises.

5.3.1 Type i , h_b and h_a theories

Let us start with the iib little string theory, where we can define a procedure very close to that of [51, 48]. In this theory we have $d5$ -branes (cfr. Table 5.3), which are Dirichlet branes for the little iib fundamental strings, filling the 6-dimensional space-time. They are thus the analog of the D9-branes of type IIB theory. We now go to one of the precise pictures defining iib strings to analyze the structure of the theory defined by iib in presence of a certain number n of $d5$ -branes.

If we take the KK5(B) picture (see Table 5.4), the $d5$ -brane arises from the $D = 10$ D5-brane with its world-volume inside the KK5. It is now straightforward to identify which BPS states of the iib theory survive the “projection” due to the presence of the $d5$ -branes. From the 10-dimensional supersymmetry relations listed in Appendix B, we can see that only D1-branes can live at the same time within the KK5 and the D5-branes. The closed $f1$, coinciding with the F1, is no longer a BPS state, and the same occurs to the $d3$ and the $s5$. We are thus left with a theory of open little strings (the open IIB strings within the D5-brane), with a $d1$ -brane BPS state. We propose to call this theory type i .

Note that along with the n D5-branes, one can also add an $\Omega 5$ orientifold plane⁵ without breaking further supersymmetry. Since there are still 3 non-compact transverse directions, the SO or Sp nature of the orientifold and the number of D5-branes is not fixed by simple charge flux arguments. Therefore, unlike the 10 dimensional case, here we can have a priori arbitrary $U(n)$, $SO(2n)$ or $Sp(2n)$ gauge groups on the D5-branes. The $\Omega 5$ defines an $\omega 5$ little orientifold plane for the iib theory.

If there is only one KK5 brane, the gauge group discussed above corresponds to the gauge group of the little type i string theory. On the other hand, if there are N coinciding KK5 branes this issue is more subtle. Note also that we could have carried out all the above discussion starting from another configuration, for instance the one with N NS5(A)-branes within n D6-branes. The latter configuration is closer to a Hanany-Witten set up [64], and could thus be useful when considering the low-energy effective action of the type i strings. We return on this at the end of the section.

⁵ Much in the same way as it was introduced in [167] in the context of brane configurations describing field theory dualities involving SO and Sp groups.

Theory	Defined by:	BPS objects
i	$iib + d5$	$d1$
h_a	$iaa + s5$	$f1$
h_b	$iib + s5$	$f1$

Table 5.7: Main characteristics of the theories with 8 supercharges

In order to define a (1,0) closed string theory, we can simply apply the s -duality of iib strings to the type i theory. This duality maps the $d1$ branes to the $f1$ little strings, and most notably the $d5$ -branes to the $s5$ -branes. The only BPS states of this theory are thus the $f1$. We call this theory h_b . We could have directly found this h_b theory from the iib one by piling up n $s5$ -branes. If we are allowed to define the s -dual of the $\omega 5$ orientifold, then this procedure is reminiscent of the one used by Hull [53] to obtain the heterotic $SO(32)$ theory from type IIB. The possible groups characterizing the h_b theory should be the same as the ones for type i theory.

For future convenience, we recall that the above theory can be defined by a configuration with generically N NS5(A)-branes (the host branes) parallel to n KK5(A) monopoles (giving rise to the $s5$ -branes).

There is still a 5-dimensional object in the little string theories that could be used to define a new (1,0) theory, namely the $s5$ -brane of the type iaa theory. Taking the NS5(B) as the host brane, we obtain this theory piling up n KK5(B)-monopoles parallel to its world-volume. In this case the low-energy effective action, and its gauge symmetries, is more subtle to define, as we will see at the end of the section. We call this little string theory h_a . It is t -dual to the h_b one.

Using the duality between iaa and m -theory, we can see that iaa in presence of $s5$ -branes is dual to m -theory with $m5$ -branes, which are domain walls, or boundaries of the 7-dimensional ‘little’ space-time. Thus the h_a theory can be seen as an m -theory compactification in presence of $m5$ -branes. This description is very rough and schematic, but could be related to a 7-dimensional analog of the Horava-Witten mechanism [54] to obtain the $E_8 \times E_8$ heterotic string theory (although in [54] the 9-dimensional objects are really boundaries rather than branes). Note also that the m -theory configuration giving h_a theory can also be seen as NS5-branes within D6-branes, but this time we have generically n NS5-branes and N (host) D6-branes, that is the opposite than for the configuration leading to type i theory.

We thus see that the pattern of dualities that arises between the theories with 8 and 16 supercharges is very similar to the one between $N = 1$ and $N = 2$ string theories in 10 dimensions. We list in Table 5.7 the main characteristics of the (1,0) little string theories.

5.3.2 The low-energy effective actions and further comments

We now turn to the discussion of some speculative points related to the theories discussed above.

We would like to find the candidates for the low-energy effective theories of the little theories with $(1, 0)$ supersymmetry. This task is however more subtle than for the theories with 16 supercharges. Let us begin with h_a theory. Its low energy theory is a theory that has actually been extensively studied in the literature [168, 169, 170]. It can be alternatively seen as arising from one of the following two configurations, represented schematically as $(N) \text{ NS5(B)} \parallel (n) \text{ KK5(B)}$ and $(N) \text{ D6} \rightarrow (n) \text{ NS5(A)}$. Since the background of n KK5 monopoles can be related to an ALE space with a A_{n-1} singularity [156], the first configuration makes the link with the theories discussed in [168, 169]. Actually, in [169] two little string theories are defined as the theories on the NS5-branes at a A_{n-1} singularity. They are very likely to coincide with our h_b and h_a theories.

The second configuration with the D6-branes suspended between the NS5(A)-branes is closer to a Hanany-Witten configuration, and has indeed been considered in [170], with results coinciding with [168, 169]. As usual in this kind of configurations, the theory on the D6-brane is reduced to a 5+1 dimensional theory by the presence of the NS5-branes. In this case however there are two differences with respect to the cases considered in [64]. First of all, D6-brane charge conservation imposes that on every NS5-brane there are as many D6-branes on the right than on the left. This is clearly the case in our configuration since the N D6-branes actually wrap a compact direction. Secondly, since the world-volume of the NS5-brane is also 5+1 dimensional, it also contributes to the resulting effective field theory.

The low-energy field theory, for a generic set up in which the NS5-branes are at finite distance on the D6-branes, thus contains the $(1, 0)$ vector multiplets yielding a gauge group $U(N)^n$ (this is true if the direction of the D6-branes perpendicular to the NS5-branes is compact, as considered for lower dimensional D-branes in e.g. [171, 67]), some matter hypermultiplets coming from the massless modes of the open strings stretching between different D6-branes, and n self-dual tensor multiplets coming from the NS5-branes. This is precisely the right matter content to make the field theory above anomaly free (the account of the precise mechanism can be found in [168]). The naive argument that there was no constraint on n because the net flux could be non-vanishing was thus enough to ensure anomaly freedom at the level of the low-energy effective action.

In the limit in which all the NS5-branes coincide, the scalars in the tensor multiplets have vanishing expectation value (except one) and all but one of the SYM couplings become infinite. It is not clear what kind of theory is left, but there should be a restoration of a (global) $U(n)$ symmetry and tensionless strings (coming from D2-branes stretching from one NS5-brane to another) should also come into play.

Note also that in the case where there is only one host brane ($N = 1$), the effective theory should contain a single and decoupled $U(1)$ vector field and n tensor fields coming from the NS5-branes, but an explicit check of the anomaly cancellation should be performed.

Let us now consider the low-energy effective action of h_b theory. This should actually coincide with the low-energy action of the type i theory, since s -duality does not affect the world-volume theories. Depending on whether we see it as representing type i or h_b theory, the relevant configuration is respectively $(N)\text{NS5(A)} \subset (n)\text{D6}$ or $(N)\text{NS5(A)} \parallel (n)\text{KK5(A)}$. If we perform a T-duality along the NUT direction of the second configuration, we obtain $(N)\text{KK5(B)} \parallel (n)\text{NS5(B)}$. We thus obtain that the low-energy effective action of h_b/i theories is exactly the same as the one for h_a theory, but with N and n interchanged (see also [169, 172]). The theory has a $U(n)^N$ gauge group. It is now more compelling to understand what happens when all the host branes coincide. The difficulty is related to the existence of a non-abelian generalization of the theory of N self-dual tensors. Note that in this latter case, the configuration with only one host brane is actually the simplest precisely because the problem above is eluded. We recover a $U(n)$ gauge theory with a (decoupled) tensor multiplet.

The specularity of these two low-energy effective actions is puzzling and could well have a physical meaning. Indeed, because of the t -duality relating the h_a and h_b theories, also the two low-energy theories should coincide upon compactification on a circle. To help in clarifying this issue, one should certainly have a better understanding of the above mentioned limit in which all but one of the couplings take an infinite value.

Let us end this section with two further comments. Seiberg [152] defines (1,0) little string theories from the world-volume of the 5-branes in the two heterotic string theories. These little theories have however a global $SO(32)$ or $E_8 \times E_8$ symmetry, which is unlikely to arise in our cases. The (1,0) theories of [152] seem thus different from those discussed in this section (in the sense that it should not be possible to derive them from a pure type ii little string framework).

As a side remark, it is worth noting that the 5-branes of the little theories play a crucial rôle in the interplay between theories with 16 and 8 supercharges. By analogy, 9-branes in M-theory and in type II theories might be interesting to study. The existence of an M9-brane and NS-like 9-branes of type IIA and IIB theory was indeed discussed in [53].

5.4 Application to Matrix theory compactification

The little theories discussed above are relevant to the description of Matrix theory compactified on higher dimensional tori. We refer the reader to Section 2.6 for an introduction to Matrix theory.

In the original conjecture [24], M-theory in the infinite momentum frame (IMF) is described by the Matrix theory of a system of N D0-branes, in the large N limit. If some of the remaining 9 space directions are compactified (on T^d say), one has to correctly include in the Matrix description the additional BPS states that will fit into representations of the U-duality group of compactified M-theory. A way to achieve this is to take the system of D0-branes on T^d and transform it into a system of N Dd -branes completely wrapped on the dual torus. Then one could hope that all the physics of M-theory on T^d would be captured by the SYM theory in $d + 1$ dimensions which is the

low-energy effective action of this system of Dd -branes. This has been called the SYM prescription for Matrix theory compactifications [24, 173, 174, 175].

We will follow in this section Seiberg's formulation of Matrix theory [70] (see Section 2.6). We recall here that the two basic relations are (2.53) and (2.54):

$$\frac{R_s}{\tilde{l}_p^2} = \frac{R}{l_p^2}, \quad \frac{\tilde{L}_i}{\tilde{l}_p} = \frac{L_i}{l_p},$$

where R , l_p and L_i are the (finite) parameter of the original M-theory, and the Matrix theory limit is $R_s \rightarrow 0$, together with the large N limit. For completeness we list here the relations between quantities in the string theory in which the Dd -branes live, and Matrix theory variables (see [69, 70] and equations (2.61)–(2.65)):

$$\begin{aligned} l_s^2 &= \frac{\tilde{l}_p^3}{R_s} = R_s^{1/2} \frac{l_p^3}{R^{3/2}}, \\ \Sigma_i &= \frac{l_s^2}{\tilde{L}_i} = \frac{l_p^3}{RL_i}, \\ g_s &= R_s^{\frac{3-d}{4}} \frac{R^{\frac{3-d}{4}} l_p^{\frac{3}{2}(d-1)}}{V_d}, \\ g_{YM}^2 &= g_s l_s^{d-3} = \frac{R^{3-d} l_p^{3(d-2)}}{V_d}. \end{aligned} \tag{5.15}$$

l_s and g_s are respectively the string length and coupling; L_i and Σ_i are the sizes of the torus in M-theory and in the auxiliary (T-dualized) string theory picture respectively, and $V_d = L_1 \dots L_d$; g_{YM}^2 is the SYM coupling in $d+1$ dimensions, and it is dimensionful in $d \neq 3$. We have singled out in all the quantities above the dependence in R_s . In the Matrix theory limit, we always have that Σ_i and g_{YM}^2 are finite and $l_s \rightarrow 0$, but the string coupling blows up for $d \geq 4$.

Now for $d \geq 4$ the SYM is ill-defined because non-renormalizable (see e.g. [159]), and thus the SYM prescription for Matrix compactification seems to break down. However, what we should consider as a model for the description of M-theory on a torus is really the “theory on the D-brane” and not only its low-energy field theory limit. Furthermore, to be able to consider a system of N Dd -branes on its own, one has to take a limit in which the bulk physics in the auxiliary string theory decouples. This limit has to be compatible with the Matrix theory limit.

For Matrix theory on T^4 , it turns out [176] that the theory of D4-branes at strong string coupling coincides with the 6 dimensional (2,0) supersymmetric field theory (see [159]), which is the theory of N M5-branes in flat space (i.e. at $L_p \rightarrow 0$). For Matrix theory on T^5 , the theory of D5-branes at strong coupling is mapped [152] by a IIB S-duality to the theory of N NS5-branes at weak coupling but finite string tension, a theory which has string-like excitations. Finally, Matrix theory on T^6 is a theory of D6-branes which, at strong coupling, becomes a theory of KK6-monopoles [154, 155]. This 7-dimensional theory has membranes and, as we showed in Section 5.2, has a well-defined structure which has been called m -theory.

We will show in the remainder of this section how all the “phases” of m -theory (i.e. its 7- and 6-dimensional versions) describe M-theory on T^6 , and how some particular limits of them yield back the compactifications on T^5 and T^4 . In other words, we find the theories mentioned above [176, 152] as limits of the ia and ib little string theories.

Specializing now to $d = 6$, we consider first m -theory in the D6-brane picture. We have for the string coupling:

$$g_A = R_s^{-3/4} \frac{l_p^{15/2}}{R^{3/4} V_6}. \quad (5.16)$$

For the m -theory to be well-defined, its length scale l_m has to be a fixed parameter. Picking its value from Table 5.6, it takes the following expression in Matrix theory variables:

$$l_m^3 = g_A l_s^3 = \frac{l_p^{12}}{R^3 V_6}. \quad (5.17)$$

We thus see that Matrix theory on T^6 is described by the theory on the D6-branes at $g_A \rightarrow \infty$, $l_s \rightarrow 0$ and keeping l_m finite. This is exactly the limit used to define m -theory. We conclude that m -theory is the candidate for the Matrix theory description of M-theory on T^6 .

Knowing (5.17) and the relations between Σ 's and L 's, we can now translate the masses of the BPS states in m -theory into masses of M-theory objects. We know in advance to which kind of objects they will map to: since the BPS states break half of the supersymmetries of the little theories, they correspond to objects of M-theory in the IMF which break 1/4 of the supersymmetries. These are branes with travelling waves in the 11th direction, i.e. longitudinal branes. The remaining dimensions of these branes are wrapped on the T^6 . One could also have deduced this from the fact that the IMF energies of these states will be proportional to n the number of BPS little branes, and independent of N . Since these objects are string-like in the 5 dimensional supergravity to which M-theory is reduced, they should carry the 27 magnetic charges of this theory (i.e. they should fit into the $\mathbf{\bar{27}}$ of the U-duality group $E_6(Z)$ [20]). We indeed find the following identifications (see Table 5.5, and Appendix B for the tensions of the M-branes):

$$M_{m2} = \frac{\Sigma_i \Sigma_j}{l_m^3} = \frac{R V_6}{L_i L_j l_p^6},$$

the 15 $m2$ wrapped membranes are mapped to longitudinal M5-branes;

$$M_w = \frac{1}{\Sigma_i} = \frac{R L_i}{l_p^3},$$

the 6 momenta w are mapped to longitudinal M2-branes;

$$M_{m5} = \frac{\tilde{V}_6}{\Sigma_i l_m^6} = \frac{R V_6 L_i}{l_p^9}$$

the 6 $m5$ states are longitudinally wrapped KK6 monopoles (the NUT direction being always on the T^6). All these 27 states can be found also in the ia and ib pictures

to be discussed below, although the identification is less straightforward. This clearly convinces that the little string theories are 6-dimensional phases of a description of M-theory on T^6 .

We would also like to obtain the spectrum of the 27 electric charges in 5 dimensional supergravity (fitting into the **27** of $E_6(Z)$). These correspond to completely wrapped branes in M-theory, or transverse branes in the Matrix theory language (they can be represented as boosted branes). These objects preserve 16 supercharges in the Matrix model, and thus are totally supersymmetric states of the little theory. In the low-energy SYM picture of the little theories, some of these transverse branes can be associated to the electric and magnetic fluxes of the SYM [175, 132, 158]. However the transverse M5-branes are missing from this description, which is thus incomplete (note also that in the previous case there are no BPS states in the SYM which would represent the longitudinal KK6, or m -theory's $m5$). Going back to the D6-brane picture, one can find all these half-supersymmetric states by embedding in the D6-branes other branes of type IIA theory in a way that they make a non-threshold bound state (the archetype of these states is the supergravity solution of [129]). These states can be found by chains of dualities from [129] and are: $F1 \subset D6$, $D4 \subset D6$ and $KK5 \subset D6$. The energy of the F1, for instance, is given by:

$$E_{F1} = \sqrt{M_{D6}^2 + M_{F1}^2} - M_{D6} = \frac{M_{F1}^2}{2M_{D6}},$$

since in the Matrix theory limit $M_{D6} \gg M_{F1}$. We obtain:

$$E_{F1} = \frac{\sum_i^2 l_m^3}{2\tilde{V}_6} = \frac{R}{2L_i^2},$$

where in the first equality we have already substituted for the m -theory variables. This energy is to be compared with the energy of an object of mass M in the IMF, which is $E = \frac{1}{2}RM^2$. The $F1 \subset D6$ thus corresponds to transverse KK momenta on the T^6 . Much in the same way we can identify the $D4 \subset D6$ and the $KK5 \subset D6$ as corresponding respectively to the transverse M2 and M5. The energy of these states can be found e.g. in [158]. When there are N D6-branes and n other branes inside them, the IMF energy goes like n^2/N .

We now discuss the other pictures and the other little theories, along with the relations between their parameters and the Matrix theory variables. It is clear that the parameters of the little theories, once expressed in Matrix variables, will no longer depend on the picture by which the little theory was defined.

The *via* theory is most easily obtained going from the D6 to the D5 picture by T-duality. The reason to do this could be that one of the radii of the original T^6 is much bigger than the others, and we might want to decompactify it eventually. Then the parameters characterizing the IIB auxiliary theory in which the D5 lives are given by:

$$g_B = R_s^{-1/2} \frac{l_p^6}{R^{1/2} V_5}, \quad l_s^2 = R_s^{1/2} \frac{l_p^3}{R^{3/2}}, \quad R_B = R_s^{1/2} \frac{L_6}{R^{1/2}} (= \tilde{L}_6), \quad (5.18)$$

where $V_5 = L_1 \dots L_5$. The parameters of the little *ii*a string theory can be easily extracted using Table 5.2:

$$g_a = \frac{V_5^{1/2}}{l_p^{3/2} L_6}, \quad t_a = \frac{R^2 V_5}{l_p^9}. \quad (5.19)$$

In the Matrix theory limit $R_s \rightarrow 0$ the *ii*a parameters g_a and t_a are thus finite. Note however that if $L_6 \rightarrow \infty$ instead, then t_a remains fixed while g_a inevitably goes to zero. In this limit all the branes of *ii*a except the $f1$ acquire an infinite tension and thus decouple. We are left with a little string theory at zero coupling, which has exactly the right number of states to describe Matrix theory on T^5 . It has indeed 5 winding plus 5 momentum BPS states, which together make up the 10 longitudinal states of Matrix theory on T^5 [152].

To show more directly that the *ii* strings tend exactly to the description of Matrix theory on T^5 given by Seiberg [152], we go to the NS5(B) picture of *ii*a strings. This is performed by an S-duality, and we obtain for the IIB parameters:

$$\tilde{g}_B = \frac{1}{g_B} = R_s^{1/2} \frac{R^{1/2} V_5}{l_p^6}, \quad \tilde{l}_s^2 = g_B l_s^2 = \frac{l_p^9}{R^2 V_5}, \quad \tilde{R}_B = R_B = R_s^{1/2} \frac{L_6}{R^{1/2}}. \quad (5.20)$$

Now $\tilde{g}_B \rightarrow 0$ and \tilde{l}_s is finite in the Matrix theory limit. This was the original setting for the definition of the little string theories proposed by Seiberg to describe Matrix theory on T^5 . Consistently with the picture of Section 5.2, they are thus the zero coupling limit of the more complete type *ii* little string theories that describe Matrix theory on T^6 .

In order to go to the *iib* theory, we perform a T-duality along, say, the $\hat{5}$ direction. We obtain a NS5-brane in a IIA theory characterized by:

$$\tilde{g}_A = \frac{\tilde{g}_B \tilde{l}_s}{\Sigma_5} = R_s^{1/2} \frac{R^{1/2} V_4^{1/2} L_5^{3/2}}{l_p^{9/2}}, \quad \tilde{l}_s^2 = \frac{l_p^9}{R^2 V_4 L_5}, \quad \tilde{R}_A = R_s^{1/2} \frac{L_6}{R^{1/2}}, \quad (5.21)$$

with $V_4 = L_1 \dots L_4$. It is worth noting that from the *iib* point of view, the 5th direction has a radius:

$$\Sigma'_5 = \frac{\tilde{l}_s^2}{\Sigma_5} = \frac{l_p^6}{R V_4}. \quad (5.22)$$

This expression is finite and has forgotten all dependence on L_5 . Thus, we should no longer think of the fifth direction of the NS5(A) brane as related to the fifth direction of the original T^6 . Moreover, from (5.15) we can identify $\Sigma'_5 = g_{YM(4+1)}^2$, as in [176]. The parameters of the *iib* theory are given by:

$$g_b = \frac{L_5}{L_6}, \quad t_b = \frac{R^2 V_4 L_5}{l_p^9}. \quad (5.23)$$

Of course, we could have computed these parameters without leaving the little string theories, by t -duality from the *ii*a-theory. For $L_6 \rightarrow \infty$ and L_5 finite, t_b can be fixed but $g_b \rightarrow 0$ and we recover again the second string theory with 16 supercharges proposed by Seiberg [152]. When $L_5, L_6 \rightarrow \infty$, i.e. when we wish to consider a compactification

on T^4 , the tension of the little strings becomes very large, only the massless modes contribute, and we are left with a field theory of a special kind, which is however still 6 dimensional. To properly identify it, we can go to the M5 picture, since when L_5 is non compact we are at strong IIA coupling.

The M5 picture is easily obtained by decompactification of a new direction in the auxiliary M-theory, the radius of which we denote as R_1 . The parameters are thus:

$$R_1 = \tilde{g}_A \tilde{l}_s = \tilde{L}_5, \quad R_2 = \tilde{L}_6, \quad L_p^3 = R_s^{1/2} \frac{l_p^9}{R^{5/2} V_4}. \quad (5.24)$$

We clearly have $L_p \rightarrow 0$, which means that Matrix theory on T^4 is described by the theory on M5-branes in flat space. We have thus reproduced the results of [176].

As a last remark on this issue, note that we could have gone to the M5 picture from the D5 one through a T-duality on $\hat{5}$ which would have transformed the D5 into a D4, and then elevating the latter to an M5-brane. Though the labelling of the directions in the auxiliary theory is clearly different in this M5 from the one of the previous paragraph, when expressed in Matrix variables the quantities are exactly the same. This is related to the fact shown in (5.22) that in the *iib* picture the “base space” does not refer any more to the original L_5 .

5.5 Discussion

We have given in this chapter a description of little theories in 6 and 7 dimensions. Our analysis is entirely based on the spectrum of BPS states present in each one of these theories. The focus on BPS states was partly motivated by the application of these little theories to the Matrix theory description of M-theory compactifications, and to the necessity to recover the right U-duality group. It would be however very interesting to pursue the study of these non-critical string theories and m -theory beyond the BPS analysis. A full quantum and possibly non-perturbative formulation of these theories will elucidate the direct relation between the little string theories or m -theory and their low-energy effective actions, which were argued not to contain gravity. In other words, this formulation should reproduce the low-energy effective action of the host branes used to define the little theories. It may also help in understanding the full structure of the (2,0) field theory in 6 dimensions.

A way to approach these theories is to give them a Matrix-like formulation. This means that one would hope to study them non-perturbatively by formulating them in the infinite momentum frame or in the discrete light cone quantization. This is particularly appealing for the (2, 0) theory since the low-energy field theory has no perturbative limit. This approach has been initiated in [177, 178, 155] and was then applied to several other related theories. We will here present it only very schematically.

The basic idea is to build the Matrix model for m -theory along the same lines as the Matrix model for M-theory, i.e. Matrix theory. In this line of thought, the $d0$ -branes of type *ia* little strings could be the partons of m -theory in the IMF (or in the DLCQ). We can actually guess the hamiltonian governing their dynamics, following [155]. If we

take for instance the NS5(B) picture for the *iii* theory, then the $d0$ -branes are given by the D1-branes winding around the transverse compact direction. Generically, we have N NS5-branes with n D1-branes suspended between each pair of NS5-branes. This is again a configuration similar to a Hanany-Witten set up [64], and the theory of the $d0$ -branes can then be formulated along the same lines. The theory will have 8 supercharges, and scalars both in the adjoint of the gauge groups, describing the separation of the D1-branes between themselves (parametrizing the Coulomb branch), and in the (bi)fundamental, representing the motion away from the NS5-branes (parametrizing the Higgs branch). The Matrix model of m -theory is then the quantum mechanics on the Coulomb branch of the theory above. For the model to be well defined, the Coulomb branch has to decouple from the Higgs branch (which is equivalent in this picture to asking that the D1-branes do not leak into the bulk), and this limit has to coincide with the region of validity of m -theory.

Let us only mention that the approach of [177, 178] was somewhat different, also because they attempted to describe the $(2, 0)$ field theory and the related little string theory of [152] respectively. There the partons of the IMF formulation of the $(2, 0)$ field theory were taken to be the instantons in 4+1 dimensional SYM, i.e. the theory reduces to the dynamics of D0-branes within D4-branes. This corresponds to the quantum mechanics on the Higgs branch of a definite theory. It is possible to relate the two approaches by acting with some dualities on the former one and then taking some limits, as discussed in [155].

The little theories described in this chapter are clearly not yet firmly established, nevertheless they seem not only interesting on their own, but also important in the more vast context of M-theory. The knowledge of the little theories could be crucial in understanding some particular problem in M-theory, or alternatively they could be used as a tool for studying some other issue. Their last, but in our opinion not the least, attractive feature is that, as they seem to closely reproduce the web of dualities of M-theory, they could be used as a fairly accurate toy model for it.

Chapter 6

The four dimensional Schwarzschild black hole as an extremal brane configuration

In Chapter 3 we have shown that using intersecting brane configurations, one can build five and four dimensional black holes which, despite being extremal and supersymmetric, have a non-vanishing horizon area. The latter in turn results in a non-vanishing Bekenstein-Hawking entropy. These black holes can be considered from the standpoint of an equivalent configuration (i.e. a configuration defined by the same conserved charges) in the framework of perturbative string theory with D-branes. Precisely the supersymmetric properties of these black holes allow a computation of their statistical entropy in the weakly coupled D-brane picture. The two entropies thus computed match exactly.

This being an enormous advance in the understanding of black hole entropy, it is however still disappointing to have such an explanation only for extremal (or near-extremal) black holes. After all, one would like to understand black hole entropy in the simplest case in which it arises, namely for the Schwarzschild black hole.

In this chapter we show that it is possible, performing several boosts and dualities on a neutral black brane of M-theory to which the Schwarzschild black hole is related by trivial compactification, to map it onto a configuration of intersecting branes with four charges. The infinite boost limit is well-defined and corresponds to extremality for the intersecting brane configuration. The Bekenstein-Hawking entropy of the four dimensional Schwarzschild black hole is thus exactly reproduced by the statistical entropy of the D-brane configuration.

The chapter closely follows the paper [5].

6.1 Black hole entropy in M-theory

The formulation of some long-standing problems of black hole quantum physics in the new framework of M-theory has proven successful. Using D-brane techniques a statistical explanation of the entropy for some black holes has been discovered. The entropy has

been computed in terms of the degeneracy of D-brane configurations describing at weak coupling, charged black holes in the extremal and near-extremal limit [61, 62, 179, 180, 181, 182, 103]. Unfortunately, this systematic approach cannot be applied directly to neutral Schwarzschild black holes. Microscopic considerations based on Matrix theory have however been discussed recently [183, 184, 185, 186, 187, 188]. Other considerations [189] involve a connection between the Schwarzschild black hole and the 2+1 dimensional BTZ black hole [190], which has been given a microscopic description in [191, 192, 193, 194, 195, 196].

In [197], it is proposed to view a Schwarzschild black hole as the compactification of a black brane in 11 dimensional supergravity and to relate it to a charged black hole with the same thermodynamic entropy. The charged black hole is obtained by subjecting the black brane to a boost [198] in an internal uncompactified direction followed by Kaluza-Klein reduction on a different radius [197] (see also [186]). Following this idea, a quantitative analysis of Schwarzschild black holes applying some M-theory concepts has been suggested in [199]. There, a near extremal limit is defined, in which the Schwarzschild radius remains arbitrarily large at infinite boost. It is proposed to use this limit to obtain the entropy of Schwarzschild black holes from the microscopic entropy of the charged ones, viewed as systems of D-branes. This was applied in [199] to the seven dimensional black hole, mapped onto a near extremal system of D3-branes.

In this chapter we apply this proposal to four dimensional Schwarzschild black holes. In order to relate a four dimensional Schwarzschild black hole to a “countable” D-brane configuration, the procedure is more involved because we will have to perform several boosts and dualities [197]. More precisely, each boost creates a (Ramond-Ramond) charge, and we will end up with a configuration of intersecting branes with four charges. In this case, the infinite boost limit leads exactly to extremality where the configuration is marginal and a standard microscopic counting of the entropy can be safely performed (because it is protected by BPS arguments). In this way the Bekenstein-Hawking entropy of the four dimensional Schwarzschild black hole is exactly recovered.

6.2 Mapping a neutral black brane on a configuration with four charges

The metric of a four dimensional Schwarzschild black hole is:

$$ds^2 = -f dt^2 + f^{-1} dr^2 + r^2 d\Omega_2^2, \quad f = 1 - \frac{r_0}{r}. \quad (6.1)$$

This metric can be trivially embedded in an 11 dimensional space-time simply by taking its product with a flat 7-dimensional (compact) space. It thus corresponds to a neutral black seven-brane¹ compactified on a $T^6 \times S^1$ characterized by sizes L_i ($i = 1 \dots 6$) and $2\pi R$:

$$ds^2 = -f dt^2 + dz^2 + \sum_{i=1}^6 dx_i^2 + f^{-1} dr^2 + r^2 d\Omega_2^2. \quad (6.2)$$

¹ Note that we will always take $r_0 \gg R, L_i$. In this regime the black brane is entropically favoured with respect to a higher dimensional black hole (see for a recent discussion [186, 197]).

The four and eleven dimensional Newton constants, respectively G_4 and $G_{11} \equiv l_p^9$, are related by:

$$G_4 = \frac{l_p^9}{2\pi L_1 \dots L_6 R}. \quad (6.3)$$

The Bekenstein-Hawking entropy of the black hole described by (6.1), or equivalently by (6.2), is given by:

$$S_{BH} = \frac{A_h}{4G_4} \equiv \frac{\pi r_0^2}{G_4} = 2\pi^2 \frac{L_1 \dots L_6 R}{l_p^9} r_0^2. \quad (6.4)$$

We now consider this neutral black seven-brane in the framework of M-theory. Most generally, we recall the precise relation between the parameters of M-theory compactified on a circle S^1 (i.e. the eleven dimensional Planck length l_p and the radius R) and the parameters of type IIA string theory, i.e. the string coupling g_s and the string length $l_s \equiv \sqrt{\alpha'}$ (see Section 2.4):

$$R = g_s l_s, \quad (6.5)$$

$$l_p^3 = 2^{4/3} \pi^{7/3} g_s l_s^3. \quad (6.6)$$

This gives the 10 dimensional Newton coupling constant:

$$G_{10} = \frac{l_p^9}{2\pi R} = 8\pi^6 g_s^2 l_s^8. \quad (6.7)$$

Using boosts (in a sense to be defined below) and dualities, we will map the above black brane onto a configuration of intersecting branes carrying 4 Ramond-Ramond charges. We will then show that there exists a limit in which the latter configuration approaches extremality in such a way that the statistical evaluation of its entropy is well-defined.

We now proceed to the careful description of all the steps leading to the final configuration corresponding to the intersection $D4 \cap D4 \cap D4 \cap D0$, which is a marginal bound state in the extremal limit.

Suppose that the coordinate z is (momentarily) non-compact. For instance, we can simply consider that we have gone to the covering space of the S^1 factor of the compact space over which the neutral black brane is wrapped. We can now perform a boost of rapidity α in that direction:

$$\begin{aligned} t &= \cosh \alpha t' + \sinh \alpha z', \\ z &= \sinh \alpha t' + \cosh \alpha z'. \end{aligned} \quad (6.8)$$

The metric of the Schwarzschild black hole embedded in 11 dimensions is of course not invariant under the boost (6.8). It acquires a non-vanishing off-diagonal $g_{t'z'}$ component:

$$ds^2 = -H_\alpha^{-1} f dt'^2 + H_\alpha \left(dz' + H_\alpha^{-1} \frac{r_0}{r} \sinh \alpha \cosh \alpha dt' \right)^2 + \sum_{i=1}^6 dx_i^2 + f^{-1} dr^2 + r^2 d\Omega_2^2,$$

$$H_\alpha = 1 + \frac{r_0}{r} \sinh^2 \alpha. \quad (6.9)$$

The resulting configuration, as viewed in 10 dimensions, is a non-extremal set of D0₁-branes (smeared on the T^6); the subscript indicates that the D0-branes have been created by the first boost. The 10 dimensional fields are the following:

$$ds^2 = -H_\alpha^{-7/8} f dt^2 + H_\alpha^{1/8} \left(\sum_{i=1}^6 dx_i^2 + f^{-1} dr^2 + r^2 d\Omega_2^2 \right),$$

$$e^\phi = H_\alpha^{3/4}, \quad A_t = H_\alpha^{-1} \frac{r_0}{r} \sinh \alpha \cosh \alpha, \quad (6.10)$$

where the metric is in the Einstein frame and we have dropped the primes. The fields above are the ones of a black brane carrying electric charge under the 1-form RR potential.

We have now to pay attention to the physical meaning of the boost, remembering that the coordinates z and z' are eventually compactified. The boost (6.8) is a coordinate transformation if z parametrizes the covering space of the S^1 , but identifications at constant t or t' clearly give two different compactified 10 dimensional theories.

If we compactify at constant t' , i.e. in the boosted frame, the length R is rescaled to the value:

$$R' = \frac{R}{\cosh \alpha}. \quad (6.11)$$

We thus define a new compactification identifying the boosted coordinate z' on intervals of length $2\pi R'$. This is not a constant time compactification as viewed from the unboosted z frame. The 10 dimensional theories defined respectively by compactification on R and R' are thus different (for instance, the 10 dimensional Newton constants will differ). They nevertheless describe the same physics at the horizon where all time intervals are blueshifted to zero. This can be seen also from the fact that if we compute the proper length of the circle parametrized by z' in (6.9), we find $2\pi R$ precisely if the calculation is performed at the horizon.

Transformations like (6.8) are thus not coordinate transformations for the 10 dimensional theory, neither ordinary dualities since from the 10 dimensional point of view a charge is created, i.e. a new parameter comes into the solution. This can also be seen by the fact that in 11 dimensions they mix a spatial coordinate with time. Note that these transformations can nevertheless be considered as dualities if one compactifies also the time coordinate [198, 200].

This ‘new’ IIA string theory is characterized by parameters g'_s and l'_s which have a well-defined dependence on the boost parameter α given by (6.5) and (6.6) where R is replaced by R' .

In the following step we perform T-dualities in the directions, say, $\hat{1}\hat{2}\hat{3}\hat{4}$ of T^6 to obtain a non-extremal configuration of D4₁-branes. We use the standard T-duality relations:

$$L_i \rightarrow \frac{4\pi^2 l_s^2}{L_i}, \quad g_s \rightarrow g_s \frac{2\pi l_s}{L_i}. \quad (6.12)$$

We now uplift this IIA configuration to 11 dimensions. Note that this is a “new” M-theory in the sense that the Planck length is now a function of α and the dependence on α of the radius of compactification has changed. We have now a non-extremal M5-brane.

To create a second charge, we perform a boost of parameter β on the eleventh direction, which is longitudinal to the M5-brane. The M5-brane is however not invariant under the boost due to its non-extremality. Following the same procedure as for the first boost, the radius of compactification of the new M-theory is rescaled by $1/\cosh\beta$. After compactification, the resulting IIA configuration corresponds to the non-extremal version of the marginal bound state $D4_1\cap D0_2$.

We then T-dualize on the $\hat{1}\hat{2}\hat{5}\hat{6}$ directions, leading to a $D4_1\cap D4_2$ configuration. The first set of D4-branes lies now in the $\hat{3}\hat{4}\hat{5}\hat{6}$ directions. Note that the T-dualities are always performed in order to obtain a configuration such that, when going to M-theory, the 11th direction is common to all the branes in the 11 dimensional configuration.

Uplifting to eleven dimension for the second time, the parameters now depend on the two boosts α and β . We are now ready to create a third charge, performing a third boost of parameter γ .

This results, after compactification and T-dualities over $\hat{1}\hat{2}\hat{3}\hat{4}$, to a non-extremal configuration $D4_1\cap D4_2\cap D4_3$, lying respectively in the $\hat{1}\hat{2}\hat{5}\hat{6}$, $\hat{3}\hat{4}\hat{5}\hat{6}$ and $\hat{1}\hat{2}\hat{3}\hat{4}$ directions.

A last uplift–boost–compactification procedure characterized by a boost parameter δ leads to our final configuration $D4_1\cap D4_2\cap D4_3\cap D0_4$. The corresponding metric in the Einstein frame is (see e.g. [120, 121, 122]):

$$\begin{aligned} ds^2 = & -H_\alpha^{-\frac{3}{8}}H_\beta^{-\frac{3}{8}}H_\gamma^{-\frac{3}{8}}H_\delta^{-\frac{7}{8}}f dt^2 + H_\alpha^{-\frac{3}{8}}H_\beta^{\frac{5}{8}}H_\gamma^{-\frac{3}{8}}H_\delta^{\frac{1}{8}}(dy_1^2 + dy_2^2) \\ & + H_\alpha^{\frac{5}{8}}H_\beta^{-\frac{3}{8}}H_\gamma^{-\frac{3}{8}}H_\delta^{\frac{1}{8}}(dy_3^2 + dy_4^2) + H_\alpha^{-\frac{3}{8}}H_\beta^{-\frac{3}{8}}H_\gamma^{\frac{5}{8}}H_\delta^{\frac{1}{8}}(dy_5^2 + dy_6^2) \\ & + H_\alpha^{\frac{5}{8}}H_\beta^{\frac{5}{8}}H_\gamma^{\frac{5}{8}}H_\delta^{\frac{1}{8}}(f^{-1}dr^2 + r^2d\Omega_2^2), \end{aligned} \quad (6.13)$$

where

$$f = 1 - \frac{r_0}{r}, \quad H_\alpha = 1 + \frac{r_0}{r} \sinh^2 \alpha, \quad (6.14)$$

and similarly for H_β , H_γ and H_δ . The non-trivial components of the RR field strengths are:

$$\begin{aligned} \tilde{F}_{ty_1y_2y_5y_6r} &= -\partial_r \left(H_\alpha^{-1} \frac{r_0}{r} \cosh \alpha \sinh \alpha \right), \\ \tilde{F}_{ty_3y_4y_5y_6r} &= -\partial_r \left(H_\beta^{-1} \frac{r_0}{r} \cosh \beta \sinh \beta \right), \\ \tilde{F}_{ty_1y_2y_3y_4r} &= -\partial_r \left(H_\gamma^{-1} \frac{r_0}{r} \cosh \gamma \sinh \gamma \right), \\ F_{tr} &= -\partial_r \left(H_\delta^{-1} \frac{r_0}{r} \cosh \delta \sinh \delta \right), \end{aligned} \quad (6.15)$$

where \tilde{F}_6 is the 10 dimensional Hodge dual of the 4-form RR field strength.

The string coupling and string length of the type $\widehat{\text{IIA}}$ theory in which this configuration is embedded are, in terms of the original quantities appearing in (6.3) and (6.4):

$$\hat{g}_s = 4\pi^{\frac{1}{2}} \frac{l_p^{\frac{9}{2}}}{L_1 L_2 L_5 L_6 R^{\frac{1}{2}}} \left(\frac{\cosh \alpha \cosh \beta \cosh \gamma}{\cosh^3 \delta} \right)^{\frac{1}{2}}, \quad (6.16)$$

$$\hat{l}_s^2 = \frac{1}{2^{\frac{4}{3}}\pi^{\frac{7}{3}}R} l_p^3 \cosh \alpha \cosh \beta \cosh \gamma \cosh \delta \quad (= l_s^2 \cosh \alpha \dots \cosh \delta). \quad (6.17)$$

The lengths of the final 6-torus over which the above configuration is wrapped are:

$$\begin{aligned} \hat{L}_{1,2} &= \frac{2^{\frac{2}{3}}}{\pi^{\frac{1}{3}}} \frac{l_p^3}{RL_{1,2}} \cosh \alpha \cosh \gamma, \\ \hat{L}_{3,4} &= L_{3,4} \cosh \beta \cosh \gamma, \\ \hat{L}_{5,6} &= \frac{2^{\frac{2}{3}}}{\pi^{\frac{1}{3}}} \frac{l_p^3}{RL_{5,6}} \cosh \alpha \cosh \beta. \end{aligned} \quad (6.18)$$

We now compute the charge densities of the D-branes in this configuration:

$$\begin{aligned} Q_{D4_1} &= \frac{1}{16\pi\hat{G}_{10}} \int_{\hat{T}_{(3,4)}^2 \times S^2} F_4 = \frac{\pi^{\frac{7}{3}}}{2^{\frac{11}{3}}} \frac{L_1^2 L_2^2 L_3 L_4 L_5^2 L_6^2 R^5 r_0 \tanh \alpha}{l_p^{21} \cosh^3 \alpha \cosh^3 \beta \cosh^3 \gamma \cosh \delta}, \\ Q_{D4_2} &= \frac{1}{16\pi\hat{G}_{10}} \int_{\hat{T}_{(1,2)}^2 \times S^2} F_4 = \frac{\pi^{\frac{5}{3}}}{2^{\frac{7}{3}}} \frac{L_1 L_2 L_5^2 L_6^2 R^3 r_0 \tanh \beta}{l_p^{15} \cosh^3 \alpha \cosh^3 \beta \cosh^3 \gamma \cosh \delta}, \\ Q_{D4_3} &= \frac{1}{16\pi\hat{G}_{10}} \int_{\hat{T}_{(5,6)}^2 \times S^2} F_4 = \frac{\pi^{\frac{5}{3}}}{2^{\frac{7}{3}}} \frac{L_1^2 L_2^2 L_5 L_6 R^3 r_0 \tanh \gamma}{l_p^{15} \cosh^3 \alpha \cosh^3 \beta \cosh^3 \gamma \cosh \delta}, \\ Q_{D0_4} &= \frac{1}{16\pi\hat{G}_{10}} \int_{\hat{T}^6 \times S^2} \star F_2 = \frac{\pi}{2} \frac{L_1 \dots L_6 R r_0 \sinh \delta}{l_p^9 \cosh \alpha \cosh \beta \cosh \gamma}. \end{aligned} \quad (6.19)$$

The charge densities above are normalized in such a way that the elementary D-branes have a charge density equal to their tension [26].

The tensions of elementary D0 and D4-branes are given by (see Appendix B):

$$\begin{aligned} T_{D0} &= \frac{1}{\hat{g}_s \hat{l}_s} = \frac{\pi^{\frac{2}{3}}}{2^{\frac{4}{3}}} \frac{L_1 L_2 L_5 L_6 R}{l_p^6} \frac{\cosh \delta}{\cosh \alpha \cosh \beta \cosh \gamma}, \\ T_{D4} &= \frac{1}{2^4 \pi^4 \hat{g}_s \hat{l}_s^5} = \frac{\pi^{\frac{4}{3}}}{2^{\frac{8}{3}}} \frac{L_1 L_2 L_5 L_6 R^3}{l_p^{12}} \frac{1}{\cosh^3 \alpha \cosh^3 \beta \cosh^3 \gamma \cosh \delta}. \end{aligned} \quad (6.20)$$

Using (6.19) and (6.20), we can now compute the different numbers of constituent D-branes of each type:

$$\begin{aligned} N_1 &= Q_{D4_1} T_{D4}^{-1} = \frac{\pi}{2} \frac{L_1 \dots L_6 R^2}{l_p^9} r_0 \tanh \alpha, \\ N_2 &= Q_{D4_2} T_{D4}^{-1} = (2\pi)^{\frac{1}{3}} \frac{L_5 L_6}{l_p^3} r_0 \tanh \beta, \\ N_3 &= Q_{D4_3} T_{D4}^{-1} = (2\pi)^{\frac{1}{3}} \frac{L_1 L_2}{l_p^3} r_0 \tanh \gamma, \\ N_4 &= Q_{D0_4} T_{D0}^{-1} = (2\pi)^{\frac{1}{3}} \frac{L_3 L_4}{l_p^3} r_0 \tanh \delta. \end{aligned} \quad (6.21)$$

Strictly speaking, these numbers represent the number of branes only in the extremal limit, but can be interpreted more generally as the difference between the number of branes and anti-branes. Note that it is also possible to compute the numbers (6.21) by evaluating after every boost the number of D0 branes created. Indeed, the numbers are strictly invariant under all the subsequent dualities and further boosts.

6.3 Infinite boost limit, extremality, and statistical entropy

We will now show that taking all the boost parameters to infinity with r_0 kept fixed is equivalent to taking the extremal limit on the intersecting D-brane configuration. In order to do this, we compute the ADM mass applying (3.82) to (6.13):

$$M = \frac{\pi L_1 \dots L_6 R}{4 l_p^9} r_0 \frac{\cosh 2\alpha + \cosh 2\beta + \cosh 2\gamma + \cosh 2\delta}{\cosh \alpha \cosh \beta \cosh \gamma \cosh \delta}. \quad (6.22)$$

This formula is to be compared with the extremal value derived from the charge densities (6.19):

$$\begin{aligned} M_{ext} &= Q_{D4_1} \hat{L}_1 \hat{L}_2 \hat{L}_5 \hat{L}_6 + Q_{D4_2} \hat{L}_3 \hat{L}_4 \hat{L}_5 \hat{L}_6 + Q_{D4_3} \hat{L}_1 \hat{L}_2 \hat{L}_3 \hat{L}_4 + Q_{D0_4} \\ &= \frac{\pi L_1 \dots L_6 R}{4 l_p^9} r_0 \frac{\sinh 2\alpha + \sinh 2\beta + \sinh 2\gamma + \sinh 2\delta}{\cosh \alpha \cosh \beta \cosh \gamma \cosh \delta}. \end{aligned} \quad (6.23)$$

It is then easy to show that when all the boost parameters are equal and are taken to infinity, the departure from extremality rapidly goes to zero as:

$$\frac{M - M_{ext}}{M_{ext}} \sim e^{-4\alpha}. \quad (6.24)$$

Note that in the same limit both M and M_{ext} go to zero as $e^{-2\alpha}$. However we have to take into account the formulas (6.16) and (6.17) which tell us that \hat{g}_s remains finite and $\hat{m}_s \equiv \hat{l}_s^{-1}$ goes to zero as $e^{-2\alpha}$. Thus the masses above are finite in string units. Note also that for all the internal directions the ratio \hat{L}_i/\hat{l}_s is finite².

Despite the fact that \hat{m}_s is vanishingly small in the limit discussed, we can neglect the massive string modes because the Hawking temperature goes to zero even faster. Indeed, applying (3.85) to (6.13), we have:

$$T_H = \frac{1}{4\pi r_0 \cosh \alpha \cosh \beta \cosh \gamma \cosh \delta}, \quad (6.25)$$

which gives a T_H/\hat{m}_s of order $e^{-2\alpha}$ in the extremal limit.

For the supergravity description to be valid and not corrected by higher order curvature terms, we have to check that the typical length scale derived from the Riemann

²This result and the finiteness of \hat{g}_s imply that the string units and the Planck units in 4 dimensions are of the same order in the boost parameters α .

tensor is much bigger than the string length. This can indeed be computed, and it holds as soon as $r_0 \gg l_p$.

In this infinite boost limit, the numbers of constituent D-branes (6.21) tend to a finite value, which moreover is large if we take $r_0 \gg l_p$ as above and we assume that R and all L_i are of order l_p or bigger. The fact that the N 's are large but tend to a finite value is a crucial element for the validity of our mapping procedure.

We are now in position to compute the microscopic degeneracy of this extremal $D4 \cap D4 \cap D4 \cap D0$ configuration. This configuration can be related to the one considered in [180] (see also [201]) by a series of T- and S-dualities (see Appendix B). Indeed, performing a T-duality in the $\hat{5}$ direction, followed by a type IIB S-duality, and then by three more T-dualities in the directions labelled by $\hat{5}$, $\hat{1}$ and $\hat{2}$, one ends up with a configuration consisting of N_2 D6-branes wrapped on the whole T^6 , N_3 NS5-branes lying in the $\hat{1}\hat{2}\hat{3}\hat{4}\hat{5}$ directions, N_1 D2-branes in the $\hat{5}\hat{6}$ directions (which become actually $N_1 N_3$ after breaking on the NS5-branes) and finally N_4 quanta of momentum in the $\hat{5}$ direction. Going to flat space, the degeneracy of these momentum excitations can be computed as in [180].

Let us review very briefly how this counting goes (a complete and introductory account can be found in [201]; it is also useful to keep in mind the D-brane physics, as reviewed in Section 2.5). Consider first the system in which N_{D2} D2-branes lie parallel to N_{D6} D6-branes. There will be all sorts of open strings connecting the various D-branes. The low-energy theory associated with this system will be based on the gauge group $U(N_{D2}) \times U(N_{D6})$, and will have 8 supercharges (thus corresponding to a $\mathcal{N}=2$ theory in $D=4$). We now search for the massless modes that can contribute to the n units of momentum, which flow along one of the directions of the D2-branes. The scalars in the vector multiplets will parametrize motion in which the D2-branes are away from the D6-branes, while the scalars in the hypermultiplets will parametrize motion of the D2-branes within the D6-branes. It turns out (see [201]) that one can take into account only the hypermultiplets in the bifundamental, which correspond to strings going from D2-branes to D6-branes and vice-versa. A careful study of the quantization of open strings with $\nu=4$ Neumann-Dirichlet directions tells us that there is a total of $N_B = 4N_{D2}N_{D6}$ such massless bosonic degrees of freedom, accompanied by an equal number N_F of fermionic degrees of freedom.

Now the N_{NS5} NS5-branes come into play. The configuration is such that the D2-branes can open and end on them (see Chapter 4). We thus end up having a total amount of $N_{NS5}N_{D2}$ D2-branes, each one suspended between two NS5-branes. The momentum flows along the intersection. We now have $N_B = N_F = 4N_{NS5}N_{D2}N_{D6}$ massless modes which contribute to it. The logarithm of the degeneracy of this 2 dimensional gas of massless modes is then given, for $n \gg N_B$, by (see e.g. [6]):

$$\ln d(n) = 2\pi \sqrt{\frac{1}{6} \left(N_B + \frac{1}{2} N_F \right) n} \equiv S_{micro}. \quad (6.26)$$

Reinserting the numbers of each set of component branes as given before, we have that $N_B = N_F = 4N_1 N_2 N_3$ and $n = N_4$. The statistical entropy is then given for large

N 's by³:

$$S_{micro} = 2\pi\sqrt{N_1 N_2 N_3 N_4}. \quad (6.27)$$

Note that all the N 's, which by (6.21) are proportional to r_0 , can be taken arbitrarily large because our limit has been defined keeping r_0 fixed and, in fact, arbitrarily large. Using the values (6.21) in the infinite boost limit, we find:

$$S_{micro} = 2\pi^2 \frac{L_1 \dots L_6 R}{l_p^9} r_0^2 = \frac{\pi r_0^2}{G_4} = \frac{A_h}{4G_4}, \quad (6.28)$$

in perfect agreement with the Bekenstein-Hawking entropy of the original four dimensional Schwarzschild black hole given in (6.4).

We end this chapter with some comments on the above result.

At first sight, counting states of a Schwarzschild black hole through a mapping onto an extreme BPS black hole seems a very indirect procedure. However the mapping crucially rests on the relation between compactification radii (see e.g. Eq. (6.11)) ensuring equality of semi-classical thermodynamic entropies. As we have seen, this mapping is equivalent, on the horizon only, to a coordinate transformation in eleven dimensions. The physics outside the horizon is different and in spacetime we have two distinct physical systems. Nevertheless, the very fact that entropy was obtained by a counting of quantum states strongly suggests that the two different systems can be related by a reshuffling of degrees of freedom defined on the horizon.

Another crucial element of our computation is the fast convergence in the infinite boost limit to a configuration of extremal BPS D-branes. This is in sharp contradistinction to the case examined in [199] where all the entropy comes from a departure from extremality. This arises because the slow vanishing of the excess mass $\Delta M \equiv M - M_{ext}$ in that case is exactly compensated by the growth of the internal volume to give a finite value to the entropy of the non-BPS excitations of the D3-branes. In our case the situation is different. The product $\Delta M L$ goes to zero in the limit, leaving a pure BPS state which can then be safely extrapolated to flat space.

Let us conclude by mentioning that in [203] the mapping procedure described in this chapter between a Schwarzschild black hole and a black hole with four charges was used to relate at low energies neutral Hawking radiation in the former to D0-brane emission in the latter. The low-energy absorption cross-sections were shown to agree, however the grey body factors were argued to be out of reach essentially because of the strong near-extremality condition.

³ The condition $N_4 \gg N_1 N_2 N_3$ can be released following the usual trick of considering multiply wrapped branes [202, 201].

Appendix A

Kaluza-Klein reduction and the resulting action

In this appendix we start from a theory including gravity in $D + p$ dimensions and we reduce it on a general p -dimensional torus with non-trivial metric. The main assumption in doing this is that all the fields do not depend on the p ‘internal’ coordinates, i.e. we will only consider the zero modes of these reduced fields. We call this procedure ‘Kaluza-Klein reduction’, and take this as a synonym of dimensional reduction.

The aim is to derive the **general action in D dimensions** which includes the **new fields arising from the reduction procedure**, namely the field components with part of their indices in the compactified space.

Let us start with a theory in $D + p$ dimensions containing only gravity. If we **single out the p internal directions**, we can recast the metric in the following form:

$$ds^2 \equiv G_{MN} dz^M dz^N = g_{\mu\nu} dx^\mu dx^\nu + h_{ij} (dy^i + A_\mu^i dx^\mu) (dy^j + A_\nu^j dx^\nu). \quad (\text{A.1})$$

Here $z^M = (x^\mu, y^i)$, $\mu = 0 \dots D - 1$ and $i = 1 \dots p$; all the functions appearing in (A.1) depend only on **x^μ** . The components of the $(D + p)$ -dimensional metric G_{MN} can thus be expressed as:

$$G_{\mu\nu} = g_{\mu\nu} + h_{ij} A_\mu^i A_\nu^j, \quad G_{\mu i} = h_{ij} A_\mu^j, \quad G_{ij} = h_{ij}. \quad (\text{A.2})$$

The inverse metric components G^{MN} , which are such that **$G^{MP} G_{PN} = \delta_N^M$** , can be easily computed to be:

$$G^{\mu\nu} = g^{\mu\nu}, \quad G^{\mu i} = -A^{i\mu} \equiv -g^{\mu\nu} A_\nu^i, \quad G^{ij} = h^{ij} + A^{i\lambda} A_\lambda^j. \quad (\text{A.3})$$

The inverse metric components $g^{\mu\nu}$ and h^{ij} are such that, respectively, **$g^{\mu\lambda} g_{\lambda\nu} = \delta_\nu^\mu$** and **$h^{ik} h_{kj} = \delta_j^i$** .

The starting point is the following action for pure gravity in $D + p$ dimensions:

$$I = \int d^{D+p} z \sqrt{-G} R[G]. \quad (\text{A.4})$$

$R[G]$ denotes the curvature scalar constructed from the metric G_{MN} . The relation between the determinants is straightforward, namely:

$$G = gh, \quad \leftrightarrow \quad \sqrt{-G} = \sqrt{-g}\sqrt{h}. \quad (\text{A.5})$$

The curvature scalar can also be recast in D -dimensional quantities. Using the equations (A.2) and (A.3), one obtains after quite a lengthy computation:

$$R[G] = R[g] - \frac{3}{4}\partial_\mu h^{ij}\partial^\mu h_{ij} - \frac{1}{4}h^{ij}\partial_\mu h_{ij} h^{kl}\partial^\mu h_{kl} - h^{ij}\square h_{ij} - \frac{1}{4}h_{ij}F_{\mu\nu}^i F^{j\mu\nu}, \quad (\text{A.6})$$

where $R[g]$ and \square are the D -dimensional curvature and D'Alembertian operator respectively, built from the metric $g_{\mu\nu}$, the field strength is given by $F_{\mu\nu}^i = \partial_\mu A_\nu^i - \partial_\nu A_\mu^i$ and all the greek indices are raised with $g^{\mu\nu}$.

It is in general better to get rid of the term containing the D'Alembertian, transforming it into a total divergence. In order to achieve this, we note that in (A.4) the above expression is multiplied by \sqrt{h} , which satisfies:

$$\frac{1}{\sqrt{h}}\square\sqrt{h} = \frac{1}{2}h^{ij}\square h_{ij} + \frac{1}{2}\partial_\mu h^{ij}\partial^\mu h_{ij} + \frac{1}{4}h^{ij}\partial_\mu h_{ij} h^{kl}\partial^\mu h_{kl}. \quad (\text{A.7})$$

The action (A.4) is thus equivalent, modulo a total divergence, to:

$$I = \int d^D x \sqrt{-g}\sqrt{h} \left\{ R[g] + \frac{1}{4}\partial_\mu h^{ij}\partial^\mu h_{ij} + \frac{1}{4}h^{ij}\partial_\mu h_{ij} h^{kl}\partial^\mu h_{kl} - \frac{1}{4}h_{ij}F_{\mu\nu}^i F^{j\mu\nu} \right\}. \quad (\text{A.8})$$

In order to put this action into its canonical form (i.e. to go to the 'Einstein frame'), we now want to eliminate the factor of \sqrt{h} in front of $R[g]$. We thus have to operate a Weyl rescaling of the metric. Most generally, if one transforms the metric by:

$$g_{\mu\nu} = e^{2\varphi}\tilde{g}_{\mu\nu}, \quad (\text{A.9})$$

the scalar curvature in D dimensions transforms accordingly and gives:

$$R[g] = e^{-2\varphi} \{ R[\tilde{g}] - 2(D-1)\square\varphi - (D-1)(D-2)\partial_\mu\varphi\partial^\mu\varphi \}, \quad (\text{A.10})$$

where the metric entering in the quantities on the r.h.s. is always $\tilde{g}_{\mu\nu}$.

The Weyl rescaling we have to do in (A.8) in order to go to the canonical action is:

$$g_{\mu\nu} = e^{-\frac{1}{D-2}\log h}\tilde{g}_{\mu\nu}. \quad (\text{A.11})$$

We can again drop a total divergence arising from the term proportional to $\square\log h$, and if we use $\partial_\mu\log h = h^{ij}\partial_\mu h_{ij}$, we finally obtain the Kaluza-Klein reduced action in D dimension, in its canonical form (for readability we drop the tildes):

$$I = \int d^D x \sqrt{-g} \left\{ R + \frac{1}{4}\partial_\mu h^{ij}\partial^\mu h_{ij} - \frac{1}{4(D-2)}h^{ij}\partial_\mu h_{ij} h^{kl}\partial^\mu h_{kl} - \frac{1}{4}h^{\frac{1}{D-2}}h_{ij}F_{\mu\nu}^i F^{j\mu\nu} \right\}. \quad (\text{A.12})$$

We can now specialize to the simplest case in which we reduce on a single direction y . This case will also turn out to be the most useful. Here we can write $h_{yy} \equiv h = e^{2\sigma}$, and the action (A.12) for this case becomes:

$$I = \int d^D x \sqrt{-g} \left\{ R - \frac{D-1}{D-2} \partial_\mu \sigma \partial^\mu \sigma - \frac{1}{4} e^{2\frac{D-1}{D-2}\sigma} F_{\mu\nu} F^{\mu\nu} \right\}. \quad (\text{A.13})$$

The kinetic term of the scalar can be put to its canonical form simply defining a new scalar $\phi = a\sigma$ such that its kinetic term is multiplied by $\frac{1}{2}$. This requirement gives for a the value:

$$a = \sqrt{\frac{2(D-1)}{D-2}}, \quad (\text{A.14})$$

and the resulting action is simply:

$$I = \int d^D x \sqrt{-g} \left\{ R - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{4} e^{a\phi} F_{\mu\nu} F^{\mu\nu} \right\}. \quad (\text{A.15})$$

It is worthwhile at this stage to recall two physically relevant cases for the expression (A.14). The first one is when 4 dimensional physics is seen as a **reduction from 5 dimensions**. In this case $D = 4$ and thus $a = \sqrt{3}$. It is this class of **4 dimensional dilatonic black holes** that will be relevant to the study of Kaluza-Klein monopoles.

The second case is the reduction of 11 dimensional supergravity to 10 dimensions. Here we have $D = 10$ and thus $a = \frac{3}{2}$. In this case however we have a physical interpretation of the scalars involved: $\langle e^\phi \rangle \equiv g_s$ the **string coupling** constant, and $\langle e^\sigma \rangle \sim R_{11}/l_p$ the size of the 11th direction in 11 dimensional Planck units. The relation $\phi = \frac{3}{2}\sigma$ thus entails $R_{11} \sim g_s^{2/3} l_p$.

For the sake of completeness, we also have to find how the matter fields, and in particular the antisymmetric tensor fields, reduce under **Kaluza-Klein compactification**. For simplicity, we proceed in the simplest case of the reduction on a single direction.

Suppose in $D+1$ dimensions we have a n -form field strength deriving from a potential:

$$H_{M_1 \dots M_n} = n \partial_{[M_1} C_{M_2 \dots M_n]} = \partial_{M_1} C_{M_2 \dots M_n} + \text{cyclic permutations}. \quad (\text{A.16})$$

Upon reduction on y , the $(n-1)$ -form potential C gives rise to two potentials $A^{(n-1)}$ and $A^{(n-2)}$, respectively an $(n-1)$ - and an $(n-2)$ -form. The relation is simply:

$$A_{\mu_1 \dots \mu_{n-1}}^{(n-1)} = C_{\mu_1 \dots \mu_{n-1}}, \quad A_{\mu_1 \dots \mu_{n-2}}^{(n-2)} = C_{\mu_1 \dots \mu_{n-2} y}. \quad (\text{A.17})$$

The corresponding field strengths are also easily found to be:

$$F_{\mu_1 \dots \mu_n}^{(n)} = n \partial_{[\mu_1} A_{\mu_2 \dots \mu_n]}^{(n-1)} = n \partial_{[\mu_1} C_{\mu_2 \dots \mu_n]} = H_{\mu_1 \dots \mu_n}, \quad (\text{A.18})$$

$$\begin{aligned} F_{\mu_1 \dots \mu_{n-1}}^{(n-1)} &= (n-1) \partial_{[\mu_1} A_{\mu_2 \dots \mu_{n-1}]}^{(n-2)} = \\ &= (n-1) \partial_{[\mu_1} C_{\mu_2 \dots \mu_{n-1}] y} = n \partial_{[\mu_1} C_{\mu_2 \dots \mu_{n-1} y]} = H_{\mu_1 \dots \mu_{n-1} y}. \end{aligned} \quad (\text{A.19})$$

The starting Lagrangian in $D + 1$ dimensions is:

$$L = -\frac{1}{2n!} H_{M_1 \dots M_n} H^{M_1 \dots M_n}. \quad (\text{A.20})$$

Using again the expressions (A.3) for $p = 1$, $h_{yy} = e^{2\sigma}$, and the above formulas (A.18) and (A.19), we find:

$$L = -\frac{1}{2(n-1)!} e^{-2\sigma} F_{\mu_1 \dots \mu_{n-1}}^{(n-1)} F^{(n-1)\mu_1 \dots \mu_{n-1}} - \frac{1}{2n!} F_{\mu_1 \dots \mu_n}^{(n)} F'^{(n)\mu_1 \dots \mu_n}, \quad (\text{A.21})$$

where we have defined the following modified n -form field strength:

$$F'_{\mu_1 \dots \mu_n}{}^{(n)} = F_{\mu_1 \dots \mu_n}^{(n)} - n F_{[\mu_1 \dots \mu_{n-1}}^{(n-1)} A_{\mu_n]}. \quad (\text{A.22})$$

The Kaluza-Klein procedure thus introduces this Chern-Simons-like coupling.

The action for the n -form can then be reexpressed in the Einstein frame like in (A.15):

$$\begin{aligned} I_{forms} &= \int d^{D+1}z \sqrt{-G} \left\{ -\frac{1}{2n!} H^2 \right\} \\ &= \int d^Dx \sqrt{-g} \left\{ -\frac{1}{2(n-1)!} e^{a_{n-1}\phi} F_{(n-1)}^2 - \frac{1}{2n!} e^{a_n\phi} F'^2_{(n)} \right\}, \end{aligned} \quad (\text{A.23})$$

where the couplings to the scalar are:

$$a_{n-1} = -(D-n) \sqrt{\frac{2}{(D-1)(D-2)}}, \quad a_n = (n-1) \sqrt{\frac{2}{(D-1)(D-2)}}. \quad (\text{A.24})$$

Note that the metric appearing in (A.23) is the Weyl rescaled one, namely the $\tilde{g}_{\mu\nu}$ in (A.11).

Again, taking the example of the reduction of the 4-form appearing in 11 dimensional supergravity, we obtain for the 3- and 4-form of the 10-dimensional type IIA supergravity respectively $a_3 = -1$ and $a_4 = \frac{1}{2}$.

Appendix B

Summary of the properties of the branes in 10 and 11 dimensions

In this appendix, we give a list of the projections imposed on the supersymmetric parameters of the theory when there is a brane in the background. We also give the tensions of the branes, and their transformation laws under the various dualities.

In 11 dimensions, the Majorana condition can be imposed on the spinors¹. Moreover, a representation of the Clifford algebra exists in which all the Γ_M matrices are real. The charge conjugation matrix C , which is such that $C\Gamma_M^T = -\Gamma_M C$ and $C^T = -C$, can be taken to be $C = \Gamma_0$. The Γ_M matrices are such that the one corresponding to the 11th direction satisfies $\Gamma_{11} = \Gamma_0 \dots \Gamma_9$.

The supersymmetry algebra of 11 dimensional supergravity (or of M-theory), including all possible p -form charges, is given by (see for instance [42, 53]):

$$\{Q, Q^T\} = \Gamma^M C P_M + \Gamma^{MN} C Z_{MN} + \Gamma^{MNPQR} C Z_{MNPQR}, \quad (\text{B.1})$$

where Q is the supersymmetry generator, which is Majorana, and $\Gamma^{M_1 \dots M_p} = \Gamma^{[M_1} \dots \Gamma^{M_p]}$. Note that we could still add p -form charges for $p = 6, 9, 10$ but these can be accounted for by a Hodge duality followed by a redefinition of the charges with $p = 5, 2, 1$ respectively.

We call BPS states those for which the matrix $\{Q, Q^T\}$ has some zero eigenvalues. This means that they are supersymmetric since some combinations of the supercharges are represented trivially by zero on them. Taking the determinant of the r.h.s. of (B.1) and asking that it vanishes, one can see (see e.g. [204]) that the supersymmetry condition will impose at the same time an equality between the mass and one component (or more) of a p -form charge, and a supersymmetry projection on the supersymmetric parameter of the theory.

In 11 dimensional supergravity (or in M-theory), we have one Majorana supersymmetric parameter ϵ . The outcome of the analysis discussed above leads to the following relations (the numbers between brackets indicate the directions longitudinal to the brane, and $W[1]$ stands for a travelling wave or KK momentum in the direction $\hat{1}$):

$$W[1]: \quad \epsilon = \Gamma_0 \Gamma_1 \epsilon$$

¹We refer the reader to e.g. [25] for a general discussion on the spinor representations and the extended supersymmetry algebras in arbitrary D space-time dimensions.

$$\begin{aligned}
\text{M2}[1,2]: & \quad \epsilon = \Gamma_0 \Gamma_1 \Gamma_2 \epsilon \\
\text{M5}[1..5]: & \quad \epsilon = \Gamma_0 \dots \Gamma_5 \epsilon \\
\text{KK6}[1..6]: & \quad \epsilon = \Gamma_0 \dots \Gamma_6 \epsilon \\
\text{M9}[1..9]: & \quad \epsilon = \Gamma_0 \dots \Gamma_9 \epsilon
\end{aligned} \tag{B.2}$$

Note that there are no other combinations of the Γ_M matrices which square to the identity. Also, the signs in the relations above have been arbitrarily chosen and fix the convention which makes the distinction between a brane and an anti-brane.

The tensions of these objects are given as follows. The quantum of mass of a KK momentum on a compact direction of radius R is:

$$M_W = \frac{1}{R}. \tag{B.3}$$

If L_p is the 11 dimensional Planck length (related to the 11 dimensional Newton constant by $G_{11} = L_p^9$), the tensions of the M2 and M5 branes are:

$$T_{M2} = \frac{\pi^{1/3}}{2^{2/3} L_p^3}, \quad T_{M5} = \frac{1}{2^{7/3} \pi^{1/3} L_p^6}. \tag{B.4}$$

The precise numerical factors are derived by dualities from type II string theories. The tension of a KK6 monopole with a transverse NUT direction of radius R_N is:

$$T_{KK6} = \frac{\pi R_N^2}{4L_p^9}. \tag{B.5}$$

This can be easily obtained from the tension of a D6-brane, which is given below. We will not discuss here the tension of the M9.

In type II theories, there are 2 Majorana-Weyl spinors ϵ_L and ϵ_R (with reference to the string origin of the relative supersymmetry generators). They satisfy the chirality conditions:

$$\epsilon_L = \Gamma_{11} \epsilon_L, \quad \epsilon_R = \eta \Gamma_{11} \epsilon_R,$$

with $\eta = +1$ for IIB theory and $\eta = -1$ for IIA theory (the overall sign is again fixed by convention). The supersymmetry projections are the following (we denote by F1 the fundamental strings of each theory):

$$\begin{aligned}
\text{F1}[1]: & \quad \begin{cases} \epsilon_L = \Gamma_0 \Gamma_1 \epsilon_L \\ \epsilon_R = -\Gamma_0 \Gamma_1 \epsilon_R \end{cases} \\
\text{W}[1]: & \quad \begin{cases} \epsilon_L = \Gamma_0 \Gamma_1 \epsilon_L \\ \epsilon_R = \Gamma_0 \Gamma_1 \epsilon_R \end{cases} \\
\text{NS5}[1..5]: & \quad \begin{cases} \epsilon_L = \Gamma_0 \dots \Gamma_5 \epsilon_L \\ \epsilon_R = -\eta \Gamma_0 \dots \Gamma_5 \epsilon_R \end{cases} \\
\text{KK5}[1..5]: & \quad \begin{cases} \epsilon_L = \Gamma_0 \dots \Gamma_5 \epsilon_L \\ \epsilon_R = \eta \Gamma_0 \dots \Gamma_5 \epsilon_R \end{cases} \\
\text{D}p[1..p]: & \quad \epsilon_L = \Gamma_0 \dots \Gamma_p \epsilon_R
\end{aligned} \tag{B.6}$$

Note that the relations for IIA theory are obtained from those of M-theory compactifying on the 11th direction. Γ_{11} plays thus the rôle of the chiral projector in 10 dimensions, and the supersymmetry parameters are related by $\epsilon_{L(R)} = \frac{1}{2}(1 \pm \Gamma_{11})\epsilon$. Also the relations of IIA and IIB theories are related by T-duality, namely under a T-duality over the \hat{i} direction the supersymmetry parameters transform (see e.g. [26] and Section 2.5) as $\epsilon_L \rightarrow \epsilon_L$ and $\epsilon_R \rightarrow \Gamma_i \epsilon_R$ (up to a sign). We recapitulate at the end of this appendix how all the branes are related by compactification and dualities.

The mass of a KK mode W is as in (B.3). Type II string theories are both characterized by the string length $l_s = \sqrt{\alpha'}$ and by the string coupling constant g . The tension of the fundamental string is:

$$T_{F1} = \frac{1}{2\pi l_s^2}. \quad (\text{B.7})$$

The tension of the solitonic NS5 branes is given by:

$$T_{NS5} = \frac{1}{(2\pi)^5 g^2 l_s^6}. \quad (\text{B.8})$$

The KK5 monopole has a tension of:

$$T_{KK5} = \frac{R_N^2}{(2\pi)^5 g^2 l_s^8}, \quad (\text{B.9})$$

where R_N is the radius of the NUT direction. Finally the tensions of the Dp -branes are given by:

$$T_{Dp} = \frac{1}{(2\pi)^p g l_s^{p+1}}. \quad (\text{B.10})$$

Let us comment on how the precise numerical factors in the tensions are found. The expression (B.7) is taken as a definition. Then the dualities are precisely defined as follows. The type IIB S-duality acts as:

$$g \rightarrow \frac{1}{g}, \quad l_s^2 \rightarrow g l_s^2,$$

while T-duality on a circle of radius R (and thus of length $L = 2\pi R$) acts as:

$$R \rightarrow \frac{l_s^2}{R} \quad g \rightarrow g \frac{l_s}{R}.$$

Using S- and T-dualities, all the tensions (B.8), (B.9) and (B.10) are found. Note that if one considers the NUT direction of the KK5 monopole as one of the directions of its world-volume, then one has to define a new tension dividing (B.9) by $2\pi R_N$.

The tensions above completely fix the 10 dimensional Newton constant. Indeed, the tensions of all electric-magnetic dual objects must satisfy:

$$T_p T_{D-p-4} = \frac{2\pi}{16\pi G_D}. \quad (\text{B.11})$$

In 10 dimensions we thus have $G_{10} = 8\pi^6 g^2 l_s^8$.

Going to 11 dimensions, one has the two conditions $G_{10} = \frac{G_{11}}{2\pi R}$ and $M_{D0} = M_W^{(11)}$ (which gives $R = gl_s$), where R is now the radius of the 11th direction. The expressions (B.4) and (B.5) are thus recovered, and can be shown to satisfy (B.11) with $G_{11} = L_p^9$.

We end with a summary of how the various branes are related by dualities and compactifications from 11 to 10 dimensions.

Under a T-duality along a direction longitudinal to the world-volume of the brane the winding modes and the KK momentum of a fundamental string are exchanged, i.e. the F1 and the W are mapped to each other; a Dp -brane becomes a $D(p-1)$ -brane; the NS5-brane and the KK5-monopole are inert.

Under a T-duality along a direction transverse to the world-volume of the brane the F1 and the W are inert; a Dp -brane is mapped to a $D(p+1)$ -brane; the NS5-brane is mapped to a KK5-monopole and vice-versa. Note that when an object in IIA (IIB) theory is said to be inert under T-duality, it exactly means that it is mapped to its analog in IIB (IIA) theory.

Under S-duality in IIB theory the F1 and the D1-brane are mapped to each other; the D5- and the NS5-branes are mapped to each other; the D3-brane, the wave W and the KK5-monopole are inert. We will not discuss the action of S-duality on the higher branes, namely the D7 and the D9.

Compactification along one of the world-volume directions or transverse to them gives, for the 11 dimensional wave W, the D0-brane and the 10 dimensional wave W respectively; for the M2-brane, the F1 and the D2; for the M5, the D4 and the NS5; for the KK6, the KK5 and the D6 (in this last case the compactification is performed along the NUT direction). We will not consider the compactification of the hypothetical M9.

Appendix C

Computation of curvature tensors

In this appendix we work out the explicit form of the components of the Ricci tensor for a general metric which is compatible with the expected form of the sought for p -brane solutions.

We will consider general diagonal metrics with components depending on only d coordinates, the other p and the time being translational invariant. To begin with we consider a set of coordinates for which the metric is isotropic in the d -dimensional space, but we do not impose any symmetry in that space. Then we consider a special case in which we have indeed spherical symmetry. In this latter case it is possible, by a reparametrization of the radial coordinate, to generalize the Ricci tensor to a non-isotropic set of coordinates.

We will use the vielbein formalism because it has several interesting by-products, such as giving without extra effort the full Riemann tensor in orthonormal coordinates, and it also uses the spin connection which will be needed when considering supersymmetry.

The vielbein formalism

The first step in the vielbein formalism is to rewrite the metric in terms of a $D \times D$ matrix $e^{\hat{\mu}}{}_{\nu}$, where the ‘hatted’ indices are indices of the D dimensional Lorentz group, while the ‘unhatted’ ones are covariant indices for D dimensional diffeomorphisms. The relations between the metric and the vielbein are the following:

$$g_{\mu\nu} = \eta_{\hat{\rho}\hat{\sigma}} e^{\hat{\rho}}{}_{\mu} e^{\hat{\sigma}}{}_{\nu}, \quad (\text{C.1})$$

$$\eta^{\hat{\rho}\hat{\sigma}} = g^{\mu\nu} e^{\hat{\rho}}{}_{\mu} e^{\hat{\sigma}}{}_{\nu}, \quad (\text{C.2})$$

where $\eta_{\hat{\rho}\hat{\sigma}}$ is the metric of flat Minkowski space-time.

If we define the 1-forms:

$$e^{\hat{\mu}} = e^{\hat{\mu}}{}_{\nu} dx^{\nu}, \quad (\text{C.3})$$

the metric can be rewritten in the orthonormal frame:

$$ds^2 = g_{\mu\nu} dx^{\mu} dx^{\nu} = \eta_{\hat{\mu}\hat{\nu}} e^{\hat{\mu}} e^{\hat{\nu}}. \quad (\text{C.4})$$

We now introduce the spin connection $\omega^{\hat{\mu}\hat{\nu}}{}_{\rho}$, which is the ‘Yang-Mills’ connection for the local Lorentz transformations. This is because the Lorentz transformations acting

on the vielbein have to be gauged away in order to keep the same number of physical degrees of freedom. Since it belongs to the adjoint representation, the spin connection is antisymmetric in the Lorentz indices.

We impose as a constraint on the vielbein that it is covariantly constant:

$$\hat{D}_\lambda e^\mu{}_\nu = \partial_\lambda e^\mu{}_\nu + \omega^\mu{}_{\rho\lambda} e^\rho{}_\nu - \Gamma_{\lambda\nu}^\sigma e^\mu{}_\sigma = 0. \quad (\text{C.5})$$

This equation implies that the metric is covariantly constant, as it should be.

Introducing also the 1-form version of the spin connection:

$$\omega^{\hat{\mu}}{}_{\hat{\nu}} = \omega^{\hat{\mu}}{}_{\hat{\nu}\lambda} dx^\lambda, \quad (\text{C.6})$$

the relation (C.5) implies the following equation:

$$de^{\hat{\mu}} + \omega^{\hat{\mu}}{}_{\hat{\nu}} \wedge e^{\hat{\nu}} = 0, \quad (\text{C.7})$$

which completely determines the spin connection.

The solution to the equation above can be easily computed using that:

$$de^{\hat{\mu}} = h^{\hat{\mu}}{}_{\hat{\nu}\hat{\rho}} e^{\hat{\nu}} \wedge e^{\hat{\rho}}, \quad h_{\hat{\mu}\hat{\nu}\hat{\rho}} = -h_{\hat{\rho}\hat{\nu}\hat{\mu}}. \quad (\text{C.8})$$

Then writing $\omega^{\hat{\mu}}{}_{\hat{\nu}} = f^{\hat{\mu}}{}_{\hat{\nu}\hat{\rho}} e^{\hat{\rho}}$ the equation (C.7) becomes:

$$h_{\hat{\mu}\hat{\nu}\hat{\rho}} = \frac{1}{2}(f_{\hat{\mu}\hat{\nu}\hat{\rho}} - f_{\hat{\rho}\hat{\nu}\hat{\mu}}) = \frac{1}{2}(f_{\hat{\mu}\hat{\nu}\hat{\rho}} + f_{\hat{\rho}\hat{\mu}\hat{\nu}}), \quad (\text{C.9})$$

which gives:

$$f_{\hat{\mu}\hat{\nu}\hat{\rho}} = h_{\hat{\mu}\hat{\nu}\hat{\rho}} - h_{\hat{\nu}\hat{\mu}\hat{\rho}} - h_{\hat{\rho}\hat{\mu}\hat{\nu}}. \quad (\text{C.10})$$

The final expression for the spin connection is thus:

$$\omega^{\hat{\mu}}{}_{\hat{\nu}} = \frac{1}{2} \left\{ e_{\hat{\nu}}{}^\sigma e_{\hat{\rho}}{}^\lambda (\partial_\sigma e^{\hat{\mu}}{}_\lambda - \partial_\lambda e^{\hat{\mu}}{}_\sigma) - e^{\hat{\mu}\sigma} e_{\hat{\rho}}{}^\lambda (\partial_\sigma e_{\nu\lambda} - \partial_\lambda e_{\nu\sigma}) - e^{\hat{\mu}\sigma} e_{\hat{\nu}}{}^\lambda (\partial_\sigma e_{\rho\lambda} - \partial_\lambda e_{\rho\sigma}) \right\} e^{\hat{\rho}}. \quad (\text{C.11})$$

However, it is often easier to determine the spin connection simply inspecting the expression for $de^{\hat{\mu}}$ and guessing which $\omega^{\hat{\mu}}{}_{\hat{\nu}}$ would satisfy eq. (C.7).

Having the spin connection, it is straightforward to determine the curvature, which is defined by:

$$\mathcal{R}^{\hat{\mu}}{}_{\hat{\nu}} = d\omega^{\hat{\mu}}{}_{\hat{\nu}} + \omega^{\hat{\mu}}{}_{\hat{\rho}} \wedge \omega^{\hat{\rho}}{}_{\hat{\nu}}, \quad \mathcal{R}_{\hat{\mu}\hat{\nu}} = -\mathcal{R}_{\hat{\nu}\hat{\mu}}. \quad (\text{C.12})$$

This curvature tensor conveys exactly the same information as the Riemann tensor. It is indeed proportional to it. Writing $\mathcal{R}^{\hat{\mu}}{}_{\hat{\nu}} = \frac{1}{2} \mathcal{R}^{\hat{\mu}}{}_{\hat{\nu}\sigma\rho} dx^\sigma \wedge dx^\rho$, one can show, using only (C.5) and (C.1)–(C.2), that the following equality holds:

$$e_{\hat{\mu}}{}^\alpha e_{\hat{\nu}}{}^\beta \mathcal{R}^{\hat{\mu}}{}_{\hat{\nu}\sigma\rho} = \partial_\sigma \Gamma_{\beta\rho}^\alpha - \partial_\rho \Gamma_{\beta\sigma}^\alpha + \Gamma_{\gamma\sigma}^\alpha \Gamma_{\beta\rho}^\gamma - \Gamma_{\gamma\rho}^\alpha \Gamma_{\beta\sigma}^\gamma \equiv \mathcal{R}^\alpha{}_{\beta\sigma\rho}. \quad (\text{C.13})$$

Note that we can also express the components of the curvature tensor in the orthonormal frame, $\mathcal{R}^{\hat{\mu}}{}_{\hat{\nu}} = \frac{1}{2} \mathcal{R}^{\hat{\mu}}{}_{\hat{\nu}\hat{\sigma}\hat{\rho}} e^{\hat{\sigma}} \wedge e^{\hat{\rho}}$. The Ricci tensor, which enters the Einstein equations, is then defined by:

$$R_{\hat{\mu}\hat{\nu}} = \mathcal{R}^{\hat{\lambda}}{}_{\hat{\mu}\hat{\lambda}\hat{\nu}}, \quad R_{\hat{\mu}\hat{\nu}} = R_{\hat{\nu}\hat{\mu}}. \quad (\text{C.14})$$

The curvature in a special case of interest

We now specialize to the computation of the curvature for a particular metric relevant to the study of p -branes. The aim of the rest of this appendix is to provide an expression for the l.h.s. of the Einstein equations (3.13).

The general metric we consider is the following:

$$ds^2 = -B^2 dt^2 + \sum_i C_i^2 dy_i^2 + G^2 \delta_{ab} dx^a dx^b. \quad (\text{C.15})$$

All the functions B , C_i ($i = 1 \dots p$) and G depend only on the x^a ($a = 1 \dots d$) coordinates, though no additional symmetry is postulated in the d dimensional space spanned by the x 's. We refer to Chapter 3 for the physical motivation leading to a metric of the form above. We call these x^a coordinates isotropic since the metric of the d dimensional subspace is the flat metric multiplied by an overall function.

Since the metric (C.15) is diagonal, the vielbeins are very easily determined:

$$e^{\hat{t}} = B dt, \quad e^{\hat{i}} = C_i dy^i, \quad e^{\hat{a}} = G dx^a. \quad (\text{C.16})$$

The non-vanishing spin connection components are then the following:

$$\omega^{\hat{t}}_{\hat{a}} = \frac{1}{G} \partial_a \ln B e^{\hat{t}}, \quad (\text{C.17})$$

$$\omega^{\hat{i}}_{\hat{a}} = \frac{1}{G} \partial_a \ln C_i e^{\hat{i}}, \quad (\text{C.18})$$

$$\omega^{\hat{a}}_{\hat{b}} = \frac{1}{G} (\partial_b \ln G e^{\hat{a}} - \partial_a \ln G e^{\hat{b}}). \quad (\text{C.19})$$

The Riemann tensor components in the orthonormal frame are:

$$\mathcal{R}^{\hat{t}\hat{i}}_{\hat{t}\hat{j}} = -\frac{1}{G^2} \partial_a \ln B \partial_a \ln C_i \delta_{ij}, \quad (\text{C.20})$$

$$\mathcal{R}^{\hat{t}\hat{a}}_{\hat{t}\hat{b}} = \frac{1}{G^2} \{ -\partial_a \partial_b \ln B - \partial_a \ln B \partial_b \ln B + \partial_a \ln B \partial_b \ln G + \partial_b \ln B \partial_a \ln G - \delta_b^a \partial_c \ln B \partial_c \ln G \}, \quad (\text{C.21})$$

$$\mathcal{R}^{\hat{i}\hat{j}}_{\hat{k}\hat{l}} = -\frac{1}{G^2} \partial_a \ln C_i \partial_a \ln C_j (\delta_k^i \delta_l^j - \delta_l^i \delta_k^j), \quad (\text{C.22})$$

$$\mathcal{R}^{\hat{i}\hat{a}}_{\hat{j}\hat{b}} = \frac{1}{G^2} \{ -\partial_a \partial_b \ln C_i - \partial_a \ln C_i \partial_b \ln C_i + \partial_a \ln C_i \partial_b \ln G + \partial_b \ln C_i \partial_a \ln G - \delta_b^a \partial_c \ln C_i \partial_c \ln G \} \delta_j^i, \quad (\text{C.23})$$

$$\begin{aligned} \mathcal{R}^{\hat{a}\hat{b}}_{\hat{c}\hat{d}} = & \frac{1}{G^2} \{ -\delta_c^a \partial_b \partial_d \ln G + \delta_d^a \partial_b \partial_c \ln G + \delta_c^b \partial_a \partial_d \ln G - \delta_d^b \partial_a \partial_c \ln G \\ & + \delta_c^a \partial_b \ln G \partial_d \ln G - \delta_d^a \partial_b \ln G \partial_c \ln G - \delta_c^b \partial_a \ln G \partial_d \ln G + \delta_d^b \partial_a \ln G \partial_c \ln G \\ & - (\delta_c^a \delta_d^b - \delta_d^a \delta_c^b) \partial_e \ln G \partial_e \ln G \}. \end{aligned} \quad (\text{C.24})$$

Repeated indices relative to the x^a coordinates are summed. Note also that for a geometry with a metric like (C.15), one can build from the Riemann tensor and the metric

a scalar showing singular behaviour if and only if at least one of the components above also has a singular behaviour. The above expressions are thus useful in searching for the singularities of a determined geometry.

We now proceed to the computation of the Ricci tensor (C.14). For our simplified problem, only three sets of components will be non-vanishing. They are the following (remember that $\delta_a^a = d$):

$$R_i^i = \frac{1}{G^2} \left\{ -\partial_a \partial_a \ln B - \partial_a \ln B \partial_a \left[\ln B + \sum_i \ln C_i + (d-2) \ln G \right] \right\}, \quad (\text{C.25})$$

$$R_j^j = \frac{1}{G^2} \left\{ -\partial_a \partial_a \ln C_i - \partial_a \ln C_i \partial_a \left[\ln B + \sum_i \ln C_i + (d-2) \ln G \right] \right\} \delta_j^i, \quad (\text{C.26})$$

$$\begin{aligned} R_b^{\hat{a}} = \frac{1}{G^2} & \left\{ -\partial_a \partial_b \left[\ln B + \sum_i \ln C_i + (d-2) \ln G \right] \right. \\ & + \partial_a \ln G \partial_b \left[\ln B + \sum_i \ln C_i + (d-2) \ln G \right] \\ & + \partial_b \ln G \partial_a \left[\ln B + \sum_i \ln C_i + (d-2) \ln G \right] \\ & - \partial_a \ln B \partial_b \ln B - \sum_i \partial_a \ln C_i \partial_b \ln C_i - (d-2) \partial_a \ln G \partial_b \ln G \\ & \left. - \delta_b^a \partial_c \partial_c \ln G - \delta_b^a \partial_c \ln G \partial_c \left[\ln B + \sum_i \ln C_i + (d-2) \ln G \right] \right\}. \quad (\text{C.27}) \end{aligned}$$

The quantity $\varphi \equiv \ln B + \sum_i \ln C_i + (d-2) \ln G$ appears ubiquitously in the expressions above. We will indeed see that it plays an important rôle in the simplification of the Einstein equations, in some particular cases (i.e. for extremal branes). Note also that the only restriction on d is that it should be 1 or bigger, otherwise the metric is constant.

Since the metric is of the special form (C.15), the Ricci tensor with Lorentz (hatted) indices actually coincides with the one carrying covariant (unhatted) indices.

The Ricci tensor components displayed in (C.25)–(C.27) are the ones which will be used when searching for multi- p -brane solutions, including the intersecting configurations.

These equations are thus most suitable for finding solutions which are static despite the fact that there might be several ‘objects’ (p -branes of different kinds which are pointlike from the point of view of the d -dimensional space of the x ’s) at a finite distance from each other. This kind of configuration is likely to reflect a balance of forces typical of a BPS saturated state, i.e. of extremal p -branes.

On the other hand, a non-extremal configuration most likely consists of a single collapsed object in the whole of d -space. It might thus turn out to be useful to reformulate the problem requiring the solution to be spherically symmetric.

The Ricci tensor for a metric with spherical symmetry

The metric (C.15) actually incorporates also the spherically symmetric case, and the non-isotropic one by a simple change of coordinates. However in General Relativity a good choice of coordinates can often substantially simplify the equations of motion and the search for their solutions. It is thus not a worthless task to rewrite the Ricci tensor components for a metric belonging to the same class of geometries, but in a slightly different coordinate system.

The metric we will consider is the following:

$$ds^2 = -B^2 dt^2 + \sum_i C_i^2 dy_i^2 + F^2 dr^2 + G^2 r^2 d\Omega_{d-1}^2. \quad (\text{C.28})$$

It has spherical symmetry in the d space, and whenever $F \neq G$ it is not isotropic in the sense that the metric (C.15) was. All the functions of course depend only on r , and $d\Omega_{d-1}^2$ is the metric on the $(d-1)$ -sphere S^{d-1} .

We could compute the Ricci tensor using the same straightforward strategy as before. We will instead use the results (C.25)–(C.27) and perform the change of coordinates, in agreement with the fact that physically this new case is a particular case and not a generalization of the previous one. We will operate in two steps: first specialize to spherical symmetry, and secondly perform a reparametrization of the radial coordinate to go to a metric like (C.28).

The first step is to take the isotropic metric (C.15) and to go to spherical coordinates. The transformation between the set of x^a ($a = 1 \dots d$) coordinates and the set $\{r, \theta_\alpha$ ($\alpha = 1 \dots d-1$) $\}$ can be written in condensed form:

$$\begin{aligned} x_a &= r \sin \theta_1 \dots \sin \theta_{a-1} \cos \theta_a, & a &= 1 \dots d-1 \\ x_d &= r \sin \theta_1 \dots \sin \theta_{d-1}, \end{aligned} \quad (\text{C.29})$$

or inversely:

$$r = \sqrt{x_1^2 + \dots + x_d^2} \equiv \sqrt{x_a x_a}, \quad (\text{C.30})$$

$$\cot \theta_\alpha = \frac{x_\alpha}{\sqrt{x_{\alpha+1}^2 + \dots + x_d^2}}, \quad \alpha = 1 \dots d-1. \quad (\text{C.31})$$

The range of the angular variables is $0 \leq \theta_1, \dots, \theta_{d-2} \leq \pi$ and $0 \leq \theta_{d-1} \leq 2\pi$. Actually to cover the full d dimensional space R^d and not half of it, one should add to (C.31) the supplementary equation $\sin \theta_{d-1} = \frac{x_d}{\sqrt{x_{d-1}^2 + x_d^2}}$, in order to take into account the sign of x_d . However it will be not necessary here since we focus on a local change of coordinates.

If f is a function of r , then the following relations hold for its derivatives, where $' \equiv \frac{d}{dr}$:

$$\begin{aligned} \partial_a f &= \frac{x_a}{r} f', \\ \partial_a \partial_b f &= \delta_b^a \frac{1}{r} f' + \frac{x_a x_b}{r^2} \left(f'' - \frac{1}{r} f' \right), \\ \partial_a \partial_a f &= f'' + \frac{d-1}{r} f'. \end{aligned} \quad (\text{C.32})$$

The Ricci tensor components (C.25)–(C.27) can thus be specialized to the spherically symmetric case:

$$R_{\hat{t}}^{\hat{t}} = \frac{1}{G^2} \left\{ -(\ln B)'' - \frac{d-1}{r}(\ln B)' - (\ln B)'\varphi' \right\}, \quad (\text{C.33})$$

$$R_{\hat{j}}^{\hat{i}} = \frac{1}{G^2} \left\{ -(\ln C_i)'' - \frac{d-1}{r}(\ln C_i)' - (\ln C_i)'\varphi' \right\} \delta_j^i, \quad (\text{C.34})$$

$$R_{\hat{b}}^{\hat{a}} = \delta_b^a \frac{1}{G^2} \left\{ -(\ln G)'' - \frac{d-1}{r}(\ln G)' - (\ln G)'\varphi' - \frac{1}{r}\varphi' \right\} \\ + \frac{x_a x_b}{r^2} \frac{1}{G^2} \left\{ -\varphi'' + \frac{1}{r}\varphi' + 2(\ln G)'\varphi' - (\ln B)'^2 - \sum_i (\ln C_i)'^2 - (d-2)(\ln G)'^2 \right\}, \quad (\text{C.35})$$

where φ was defined just after eq. (C.27).

We now want to write the components of the Ricci tensor in terms of the indices relative to the spherical coordinates. This only involves the $R_{\hat{b}}^{\hat{a}}$ components. Noting that the latter are of the form $R_{\hat{b}}^{\hat{a}} = R_{(1)}\delta_b^a + R_{(2)}\frac{x_a x_b}{r^2}$, it is easy to show, using the relations (C.29), (C.30) and (C.31), that the components in spherical coordinates are:

$$R_{\hat{r}}^{\hat{r}} = R_{(1)} + R_{(2)}, \quad R_{\hat{\beta}}^{\hat{\alpha}} = R_{(1)}\delta_{\beta}^{\alpha} \quad (\alpha, \beta = 1 \dots d-1), \quad R_{\hat{\alpha}}^{\hat{\alpha}} = 0,$$

thus giving:

$$R_{\hat{r}}^{\hat{r}} = \frac{1}{G^2} \left\{ -\varphi'' + (\ln G)'\varphi' - (\ln G)'' - \frac{d-1}{r}(\ln G)' \right. \\ \left. - (\ln B)'^2 - \sum_i (\ln C_i)'^2 - (d-2)(\ln G)'^2 \right\}, \quad (\text{C.36})$$

$$R_{\hat{\beta}}^{\hat{\alpha}} = \frac{1}{G^2} \left\{ -(\ln G)'' - \frac{d-1}{r}(\ln G)' - (\ln G)'\varphi' - \frac{1}{r}\varphi' \right\} \delta_{\beta}^{\alpha}. \quad (\text{C.37})$$

We are now ready to proceed to the second step, which is a reparametrization of the radial coordinate r . We have to go from the (isotropic) metric:

$$ds^2 = \dots + G^2 (dr^2 + r^2 d\Omega_{d-1}^2)$$

to the more general metric:

$$ds^2 = \dots + F^2 d\bar{r}^2 + \bar{G}^2 \bar{r}^2 d\Omega_{d-1}^2.$$

We will drop the bars only at the end of the computation, when it will be no longer ambiguous. We also define for the time being $\dot{f} \equiv \frac{d}{d\bar{r}} f$.

The relation above between r and \bar{r} is thus such that:

$$Gr = \bar{G}\bar{r}, \quad F = G \frac{dr}{d\bar{r}}. \quad (\text{C.38})$$

This implies the following relations for some of the quantities appearing in the Ricci tensor components (f is an arbitrary function independent from G ; note that φ is not a function of this kind):

$$\begin{aligned}
f' &= \frac{G}{F} \dot{f} \\
(\ln G)' + \frac{1}{r} &= \frac{G}{F} \left[(\ln \bar{G})' + \frac{1}{\bar{r}} \right] \\
f'' + \frac{1}{r} f' &= \frac{G^2}{F^2} \left[\ddot{f} + \dot{f} (\ln \bar{G})' + \frac{1}{\bar{r}} \dot{f} - (\ln F)' \dot{f} \right] \\
(\ln G)'' + \frac{1}{r} (\ln G)' &= \frac{G^2}{F^2} \left[(\ln \bar{G})'' + \frac{2}{\bar{r}} (\ln \bar{G})' + (\ln \bar{G})'^2 - (\ln F)' (\ln \bar{G})' - \frac{1}{\bar{r}} (\ln F)' \right].
\end{aligned}$$

These relations can now be plugged into the Ricci tensor components (C.33), (C.34), (C.36) and (C.37).

We write the result dropping the bars (thus conforming to the notation of (C.28)):

$$\begin{aligned}
R_{\hat{t}}^{\hat{t}} &= \frac{1}{F^2} \left\{ -(\ln B)'' - (\ln B)' \left[\ln B + \sum_i \ln C_i - \ln F + (d-1)(\ln G + \ln r) \right]' \right\}, \\
R_{\hat{j}}^{\hat{i}} &= \delta_j^i \frac{1}{F^2} \left\{ -(\ln C_i)'' - (\ln C_i)' \left[\ln B + \sum_i \ln C_i - \ln F + (d-1)(\ln G + \ln r) \right]' \right\}, \\
R_{\hat{r}}^{\hat{r}} &= \frac{1}{F^2} \left\{ -(\ln B)'' - \sum_i (\ln C_i)'' - (\ln B)'^2 - \sum_i (\ln C_i)'^2 + (\ln B)' (\ln F)' \right. \\
&\quad \left. + \sum_i (\ln C_i)' (\ln F)' - (d-1) \left[(\ln G)'' + (\ln G)'^2 + \frac{2}{r} (\ln G)' \right. \right. \\
&\quad \left. \left. - (\ln G)' (\ln F)' - \frac{1}{r} (\ln F)' \right] \right\}, \\
R_{\hat{\beta}}^{\hat{\alpha}} &= \delta_{\beta}^{\alpha} \frac{1}{F^2} \left\{ - \left[(\ln G)' + \frac{1}{r} \right] \left[\ln B + \sum_i \ln C_i - \ln F + (d-1)(\ln G + \ln r) \right]' \right. \\
&\quad \left. - (\ln G)'' + \frac{1}{r^2} + (d-2) \frac{F^2}{r^2 G^2} \right\}. \tag{C.39}
\end{aligned}$$

These are the Ricci tensor components which will be used when searching for general black p -brane solutions. Here also the ansätze made to solve the equations will strongly depend on the structure of the expressions above.

Bibliography

- [1] R. Argurio, F. Englert and L. Houart, “Intersection rules for p -branes”, *Phys. Lett.* **B398** (1997) 61, hep-th/9701042.
- [2] R. Argurio, F. Englert, L. Houart and P. Windey, “On the opening of branes”, *Phys.Lett.* **B408** (1997) 151, hep-th/9704190.
- [3] R. Argurio, “Intersection rules and open branes”, to be published in the proceedings of NATO Advanced Study Institute on Strings, Branes and Dualities, Cargèse, France, 26 May-14 June 1997, hep-th/9712170.
- [4] R. Argurio and L. Houart, “Little theories in six and seven dimensions”, *Nucl.Phys.* **B517** (1998) 205, hep-th/9710027.
- [5] R. Argurio, F. Englert and L. Houart, “Statistical entropy of the four dimensional Schwarzschild black hole”, to appear in *Phys.Lett. B*, hep-th/9801053.
- [6] M. B. Green, J. H. Schwarz and E. Witten, *Superstring theory*, Vol. 1 and 2, Cambridge University Press (Cambridge, 1987).
- [7] E. Kiritsis, “Introduction to superstring theory”, hep-th/9709062.
- [8] G. Veneziano, “Construction of a crossing-symmetric, Regge behaved amplitude for linearly rising trajectories”, *Nuovo Cim.* **A57** (1968) 190.
- [9] Y. Nambu, Lectures at the Copenhagen symposium, unpublished (1970);
T. Goto, “Relativistic quantum mechanics of one-dimensional mechanical continuum and subsidiary condition of dual resonance model”, *Prog.Theor.Phys.* **46** (1971) 1560.
- [10] P. Ramond, “Dual theory for free fermions”, *Phys.Rev.* **D3** (1971) 2415;
A. Neveu and J. H. Schwarz, “Factorizable dual model of pions”, *Nucl.Phys.* **B31** (1971) 86.
- [11] J. Scherk and J. H. Schwarz, “Dual models for nonhadrons”, *Nucl.Phys.* **B81** (1974) 118.

- [12] F. Gliozzi, J. Scherk and D. Olive, “Supergravity and the spinor dual model”, *Phys.Lett.* **65B** (1976) 282; “Supersymmetry, supergravity theories and the dual spinor model”, *Nucl.Phys.* **B122** (1977) 253.
- [13] M. B. Green and J. H. Schwarz, “Supersymmetrical string theories”, *Phys.Lett.* **109B** (1982) 444 (see also “Supersymmetrical dual string theory”, *Nucl.Phys.* **B181** (1981) 502; *Nucl.Phys.* **B198** (1982) 252; *Nucl.Phys.* **B198** (1982) 441).
- [14] D. J. Gross, J. A. Harvey, E. Martinec and R. Rohm “Heterotic string theory”, *Nucl.Phys.* **B256** (1985) 253; *Nucl.Phys.* **B267** (1986) 75.
- [15] P. Candelas, G. T. Horowitz, A. Strominger and E. Witten, “Vacuum configurations for superstrings”, *Nucl.Phys.* **B258** (1985) 46.
- [16] L. Dixon, J. A. Harvey, C. Vafa and E. Witten, “Strings on orbifolds”, *Nucl.Phys.* **B261** (1985) 678; *Nucl.Phys.* **B274** (1986) 285.
- [17] E. Witten, “String theory dynamics in various dimensions”, *Nucl.Phys.* **B443** (1995) 85, hep-th/9503124.
- [18] A. Sen, “Electric magnetic duality in string theory”, *Nucl.Phys.* **B404** (1993) 109, hep-th/9207053; “Strong-weak coupling duality in four-dimensional string theory”, *Int.J.Mod.Phys.* **A9** (1994) 3707, hep-th/9402002; “Dyon-monopole bound states, selfdual harmonic forms on the multi-monopole moduli space, and $SL(2, Z)$ invariance in string theory”, *Phys.Lett.* **B329** (1994) 217, hep-th/9402032.
- [19] A. Font, L. E. Ibáñez, D. Lüst and F. Quevedo, “Strong-weak coupling duality and nonperturbative effects in string theory”, *Phys.Lett.* **B249** (1990) 35;
- [20] C. M. Hull and P. K. Townsend, “Unity of superstring dualities”, *Nucl.Phys.* **B438** (1995) 109, hep-th/9410167.
- [21] C. Montonen and D. Olive, “Magnetic monopoles as gauge particles?”, *Phys.Lett.* **72B** (1977) 117.
- [22] H. Osborn, “Topological charges for $N=4$ supersymmetric gauge theories and monopoles of spin 1”, *Phys.Lett.* **83B** (1979) 321.
- [23] A. Giveon, M. Porrati and E. Rabinovici, “Target space duality in string theory”, *Phys.Rept.* **244** (1994) 77, hep-th/9401139.
- [24] T. Banks, W. Fischler, S. H. Shenker and L. Susskind, “M theory as a matrix model: a conjecture”, *Phys.Rev.* **D55** (1997) 5512, hep-th/9610043.
- [25] J. Strathdee, “Extended Poincare supersymmetry”, *Int.J.Mod.Phys.* **A2** (1987) 273.
- [26] J. Polchinski, “TASI Lectures on D-branes”, hep-th/9611050.

- [27] S. Coleman, “The quantum sine-Gordon equation as the massive Thirring model”, *Phys.Rev.* **D11** (1975) 2088.
- [28] J. H. Schwarz, “Evidence for nonperturbative string symmetries”, *Lett.Math.Phys.* **34** (1995) 309, hep-th/9411178.
- [29] J. Wess and J. Bagger, *Supersymmetry and supergravity*, Princeton University Press (1983).
- [30] E. B. Bogomol’nyi, “Stability of classical solutions”, *Sov.J.Nucl.Phys.* **24** (1976) 449;
M. K. Prasad and C. M. Sommerfield, “An exact classical solution for the ’t Hooft monopole and the Julia-Zee dyon”, *Phys.Rev.Lett.* **35** (1975) 760;
S. Coleman, S. Parke, A. Neveu and C. M. Sommerfield, “Can one dent a dyon?”, *Phys.Rev.* **D15** (1977) 544.
- [31] E. Witten and D. Olive, “Supersymmetry algebras that include topological charges”, *Phys.Lett.* **B78** (1978) 97.
- [32] K. S. Narain, “New heterotic string theories in uncompactified dimensions ≥ 10 ”, *Phys.Lett.* **169B** (1986) 41;
K. S. Narain, M. H. Sarmadi and E. Witten, “A note on toroidal compactification of heterotic string theory”, *Nucl.Phys.* **B279** (1987) 369.
- [33] P. Ginsparg, “Comment on toroidal compactification of heterotic superstrings”, *Phys.Rev.* **D35** (1987) 648.
- [34] M. Dine, P. Huet and N. Seiberg, “Large and small radius in string theory”, *Nucl.Phys.* **B322** (1989) 301.
- [35] J. Dai, R. G. Leigh and J. Polchinski, “New connections between string theories”, *Mod.Phys.Lett.* **A4** (1989) 2073.
- [36] E. Cremmer and B. Julia, “The N=8 supergravity theory. 1. The Lagrangian”, *Phys.Lett.* **80B** (1978) 48; “The SO(8) supergravity”, *Nucl.Phys.* **B159** (1979) 141.
- [37] B. Julia, “Group disintegrations”, in *Supergravity and superspace*, S. W. Hawking and M. Rocek (eds.), Cambridge University Press (Cambridge, 1981).
- [38] E. Bergshoeff, C. Hull and T. Ortin, “Duality in the type II superstring effective action”, *Nucl.Phys.* **B451** (1995) 547, hep-th/9504081.
- [39] J. H. Schwarz, “Covariant field equations of chiral N=2 D=10 supergravity”, *Nucl.Phys.* **B226** (1983) 269;
P. S. Howe and P. C. West, “The complete N=2, D=10 supergravity”, *Nucl.Phys.* **B238** (1984) 181.

- [40] J. H. Schwarz, “An $SL(2, Z)$ multiplet of type IIB superstrings”, *Phys.Lett.* **B360** (1995) 13, hep-th/9508143; “The power of M theory”, *Phys.Lett.* **B367** (1996) 97, hep-th/9510086.
- [41] P. K. Townsend, “The eleven-dimensional supermembrane revisited”, *Phys.Lett.* **B350** (1995) 184, hep-th/9501068.
- [42] P. K. Townsend, “P-brane democracy”, proceedings of the March 95 PAS-COS/John Hopkins Conference, hep-th/9507048.
- [43] E. Cremmer, B. Julia and J. Scherk, “Supergravity theory in eleven-dimensions”, *Phys.Lett.* **76B** (1978) 409.
- [44] M. J. Duff, P. S. Howe, T. Inami and K. S. Stelle, “Superstrings in $D=10$ from supermembranes in $D=11$ ”, *Phys.Lett.* **191B** (1987) 70.
- [45] G. T. Horowitz and A. Strominger, “Black strings and p -branes”, *Nucl.Phys.* **B360** (1991) 197.
- [46] R. I. Nepomechie, “Magnetic monopoles from antisymmetric tensor gauge fields”, *Phys.Rev.* **D31** (1985) 1921.
- [47] C. Teitelboim, “Monopoles of higher rank”, *Phys.Lett.* **167B** (1986) 69.
- [48] J. Polchinski, “Dirichlet-branes and Ramond-Ramond charges”, *Phys.Rev.Lett.* **75** (1996) 4724, hep-th/9510017.
- [49] P. S. Aspinwall, “Some relationships between dualities in string theory”, *Nucl.Phys.Proc.Suppl.* **46** (1996) 30, hep-th/9508154.
- [50] A. Strominger, “Massless black holes and conifolds in string theory”, *Nucl.Phys.* **B451** (1995) 96, hep-th/9504090.
- [51] A. Sagnotti, “Open strings and their symmetry groups”, in *Cargèse '87, Non-perturbative quantum field theory*, G. Mack et al. (eds.), Pergamon Press (1988) p.521;
M. Bianchi and A. Sagnotti, “On the systematics of open-string theories”, *Phys.Lett.* **B247** (1990) 517;
P. Horava, “Strings on world sheet orbifolds”, *Nucl.Phys.* **B327** (1989) 461; “Background duality of open string models”, *Phys.Lett.* **B231** (1989) 251.
- [52] J. Polchinski and E. Witten, “Evidence for heterotic-type I string duality”, *Nucl.Phys.* **B460** (1996) 525, hep-th/9510169.
- [53] C. M. Hull, “Gravitational duality, branes and charges”, *Nucl.Phys.* **B509** (1998) 216, hep-th/9705162.

- [54] P. Horava and E. Witten, “Heterotic and type I string dynamics from eleven dimensions”, *Nucl.Phys.* **B460** (1996) 506, hep-th/9510209.
- [55] E. Witten, “Bound states of strings and p -branes”, *Nucl.Phys.* **B460** (1996) 335, hep-th/9510135.
- [56] R. G. Leigh, “Dirac-Born-Infeld action from Dirichlet sigma model”, *Mod.Phys.Lett.* **A4** (1989) 2767.
- [57] C. G. Callan, J. A. Harvey and A. Strominger, “Worldsheet Approach to Heterotic Instantons and Solitons”, *Nucl.Phys.* **B359** (1991) 611; “Worldbrane Actions for String Solitons”, *Nucl. Phys.* **B367** (1991) 60.
- [58] U. H. Danielsson, G. Ferretti and B. Sundborg, “D particle dynamics and bound states”, *Int.J.Mod.Phys.* **A11** (1996) 5463, hep-th/9603081.
- [59] D. Kabat and P. Pouliot, “A comment on zero-brane quantum mechanics”, *Phys.Rev.Lett.* **77** (1996) 1004, hep-th/9603127.
- [60] M. R. Douglas, D. Kabat, P. Pouliot S. H. Shenker, “D-branes and short distances in string theory”, *Nucl.Phys.* **B485** (1997) 85, hep-th/9608024.
- [61] A. Strominger and C. Vafa, “Microscopic origin of the Bekenstein-Hawking entropy”, *Phys.Lett.* **B379** (1996) 99, hep-th/9601029.
- [62] C. G. Callan and J. M. Maldacena, “D-brane approach to black hole quantum mechanics”, *Nucl.Phys.* **B472** (1996) 591, hep-th/9602043.
- [63] A. Hashimoto and I. R. Klebanov, “Decay of Excited D-Branes”, *Phys.Lett.* **B381** (1996) 437, hep-th/9604065;
 S. R. Das and S. D. Mathur, “Comparing decay rates for black holes and D-branes”, *Nucl.Phys.* **B478** (1996) 561, hep-th/9606185;
 S. S. Gubser and I. R. Klebanov, “Emission of charged particles from four-dimensional and five-dimensional black holes”, *Nucl.Phys.* **B482** (1996) 173, hep-th/9608108.
 A. Hashimoto, “Perturbative dynamics of fractional strings on multiply wound D-strings”, *Int.J.Mod.Phys.* **A13** (1998) 903, hep-th/9610250.
- [64] A. Hanany and E. Witten, “Type IIB superstrings, BPS monopoles, and three-dimensional gauge dynamics”, *Nucl.Phys.* **B492** (1997) 152, hep-th/9611230.
- [65] S. Elitzur, A. Giveon and D. Kutasov, “Branes and N=1 duality in string theory”, *Phys.Lett.* **B400** (1997) 269, hep-th/9702014.
- [66] A. Giveon and D. Kutasov, “Brane dynamics and gauge theory”, hep-th/9802067.
- [67] E. Witten, “Solutions of four-dimensional field theories via M theory”, *Nucl.Phys.* **B500** (1997) 3, hep-th/9703166.

- [68] L. Susskind, “Another conjecture about M(atrrix) theory”, hep-th/9704080.
- [69] A. Sen, “D0-branes on T^n and Matrix theory”, hep-th/9709220.
- [70] N. Seiberg, “Why is the Matrix model correct?”, *Phys.Rev.Lett.* **79** (1997) 3577, hep-th/9710009.
- [71] T. Banks, “Matrix theory”, hep-th/9710231;
D. Bigatti and L. Susskind, “Review of Matrix theory”, hep-th/9712072;
W. Taylor, “Lectures on D-branes, gauge theory and M(atrices)”, hep-th/9801182.
- [72] S. Hellerman and J. Polchinski, “Compactification in the lightlike limit”, hep-th/9711037;
A. Bilal, “A comment on compactification of M theory on an (almost) lightlike circle”, hep-th/9801047; “DLCQ of M theory as the light-like limit”, hep-th/9805070.
- [73] M. R. Douglas, “Branes within branes”, hep-th/9512077.
- [74] P. K. Townsend, “D-branes from M-branes”, *Phys.Lett.* **B373** (1996) 68, hep-th/9512062.
- [75] B. de Wit, J. Hoppe and H. Nicolai, “On the quantum mechanics of supermembranes”, *Nucl.Phys.* **B305** (1988) 545;
B. de Wit, M. Luscher and H. Nicolai, “The supermembrane is unstable”, *Nucl.Phys.* **B320** (1989) 135;
B. de Wit, U. Marquard and H. Nicolai, “Area preserving diffeomorphisms and supermembrane Lorentz invariance”, *Commun.Math.Phys.* **128** (1990) 39.
- [76] M. Dine and A. Rajaraman, “Multigraviton scattering in the Matrix model”, hep-th/9710174.
- [77] M. R. Douglas and H. Ooguri, “Why Matrix theory is hard”, hep-th/9710178.
- [78] M. Berkooz and M. R. Douglas, “Five-branes in M(atrrix) theory”, *Phys.Lett.* **B395** (1997) 196, hep-th/9610236.
- [79] T. Banks, N. Seiberg and S. Shenker, “Branes from Matrices”, *Nucl.Phys.* **B490** (1997) 91, hep-th/9612157.
- [80] N. Seiberg and E. Witten, “Electric-magnetic duality, monopole condensation, and confinement in N=2 supersymmetric Yang-Mills theory”, *Nucl.Phys.* **B426** (1994) 19, hep-th/9407087; “Monopoles, duality and chiral symmetry breaking in N=2 supersymmetric QCD”, *Nucl.Phys.* **B431** (1994) 484, hep-th/9408099.

- [81] M. Huq and M. A. Namazie, “Kaluza-Klein supergravity in ten-dimensions”, *Class.Quant.Grav.* **2** (1985) 293;
F. Giani and M. Pernici, “N=2 supergravity in ten-dimensions”, *Phys.Rev.* **D30** (1984) 325;
I. C. G. Campbell and P. C. West, “N=2 D=10 nonchiral supergravity and its spontaneous compactification”, *Nucl.Phys.* **B243** (1984) 112.
- [82] K. S. Stelle, “Lectures on supergravity p -branes”, hep-th/9701088.
- [83] G. W. Gibbons and K. Maeda, “Black holes and membranes in higher-dimensional theories with dilaton fields”, *Nucl.Phys.* **B298** (1988) 741.
- [84] M. J. Duff, H. Lü and C. N. Pope, “The black branes of M theory”, *Phys.Lett.* **B382** (1996) 73, hep-th/9604052.
- [85] G. W. Gibbons, G. T. Horowitz and P. K. Townsend, “Higher dimensional resolution of dilatonic black hole singularities”, *Class.Quant.Grav.* **12** (1995) 297, hep-th/9410073.
- [86] C. W. Misner, K. S. Thorne and J. A. Wheeler, *Gravitation*, Freeman (San Francisco, 1973).
- [87] J. X. Lu, “ADM masses for black strings and p -branes”, *Phys.Lett.* **B313** (1993) 29, hep-th/9304159.
- [88] G. W. Gibbons and S. W. Hawking, “Action integrals and partition functions in quantum gravity”, *Phys. Rev.* **D15** (1977) 2752.
- [89] S. W. Hawking, G. T. Horowitz and S. F. Ross, “Entropy, area, and black hole pairs”, *Phys. Rev.* **D51** (1995) 4302, gr-qc/9409013.
- [90] S. W. Hawking and S. F. Ross, “Duality between electric and magnetic black holes”, *Phys.Rev.* **D52** (1995) 5865, hep-th/9504019.
- [91] M. Bañados, C. Teitelboim and J. Zanelli, “Black hole entropy and the dimensional continuation of the Gauss-Bonnet theorem”, *Phys. Rev. Lett.* **72** (1994) 957, gr-qc/9309026.
- [92] T. Regge and C. Teitelboim, “Role of surface integrals in the Hamiltonian formulation of general relativity”, *Annals Phys.* **88** (1974) 286.
- [93] I. R. Klebanov and A. A. Tseytlin, “Entropy of near extremal black p -branes”, *Nucl.Phys.* **B475** 164, hep-th/9604089.
- [94] M. J. Duff and K. S. Stelle, “Multimembrane solutions of D=11 supergravity”, *Phys.Lett.* **B253** (1991) 113.

- [95] R. Güven, “Black p -brane solutions of D=11 supergravity theory”, *Phys.Lett.* **B276** (1992) 49.
- [96] A. Dabholkar, G. W. Gibbons, J. A. Harvey and F. Ruiz-Ruiz, “Superstrings and solitons”, *Nucl.Phys.* **B340** (1990) 33.
- [97] S. W. Hawking, “Gravitational instantons”, *Phys.Lett.* **60A** (1977) 81.
- [98] G. Papadopoulos and P. K. Townsend, “Intersecting M-branes”, *Phys.Lett.* **B382** (1996) 65, hep-th/9603087.
- [99] A. Strominger, “Open p -branes”, *Phys.Lett.* **B383** (1996) 44, hep-th/9512059.
- [100] K. Becker and M. Becker, “Boundaries in M-theory”, *Nucl.Phys.* **B472** (1996) 221, hep-th/9602071.
- [101] A. A. Tseytlin, “Harmonic superpositions of M-branes”, *Nucl.Phys.* **B475** (1996) 149, hep-th/9604035.
- [102] J. P. Gauntlett, D. A. Kastor and J. Traschen, “Overlapping branes in M-theory”, *Nucl.Phys.* **B478** (1996) 544, hep-th/9604179.
- [103] V. Balasubramanian and F. Larsen, “On D-branes and black holes in four dimensions”, *Nucl.Phys.* **B478** (1996) 199, hep-th/9604189.
- [104] E. Bergshoeff, M. de Roo, E. Eyras, B. Janssen and J. P. van der Schaar, “Multiple intersections of D-branes and M-branes”, *Nucl.Phys.* **B494** (1997) 119, hep-th/9612095.
- [105] A. A. Tseytlin, “No-force condition and BPS combinations of p -branes in 11 and 10 dimensions”, *Nucl.Phys.* **B487** (1997) 141, hep-th/9609212.
- [106] I. Ya. Aref’eva and O. A. Rytchkov, “Incidence matrix description of intersecting p -brane solutions”, hep-th/9612236.
- [107] N. Khviengia, Z. Khviengia, H. Lü and C. N. Pope, “Intersecting M-branes and bound states”, *Phys.Lett.* **B388** (1996) 21, hep-th/9605077.
- [108] G. W. Gibbons, M. B. Green and M. J. Perry, “Instantons and seven-branes in type IIB superstring theory”, *Phys.Lett.* **B370** (1996) 37, hep-th/9511080.
- [109] B. R. Greene, A. Shapere, C. Vafa and S.-T. Yau, “Stringy cosmic strings and noncompact Calabi-Yau manifolds”, *Nucl.Phys.* **B337** (1990) 1.
- [110] T. Banks and L. Susskind, “Brane–anti-brane forces”, hep-th/9511194.
- [111] I. Ya. Aref’eva, M. G. Ivanov and O. A. Rytchkov, “Properties of intersecting p -branes in various dimensions”, hep-th/9702077.

- [112] M. J. Duff, B. E. W. Nilsson and C. N. Pope, “Kaluza-Klein supergravity”, *Phys.Rept.* **130** (1986) 1.
- [113] G. Lifschytz, “Comparing D-branes to black-branes”, *Phys.Lett.* **B388** (1996) 720, hep-th/9604156.
- [114] H. J. Sheinblatt, “Statistical entropy of an extremal black hole with 0-brane and 6-brane charge”, *Phys.Rev.* **D57** (1998) 2421, hep-th/9705054.
- [115] R. R. Khuri and T. Ortin, “A nonsupersymmetric dyonic extreme Reissner-Nordstrom black hole”, *Phys.Lett.* **B373** (1996) 56, hep-th/9512178.
- [116] W. Taylor, “Adhering 0-branes to 6-branes and 8-branes”, *Nucl.Phys.* **B508** (1997) 122, hep-th/9705116.
- [117] N. S. Manton, “The force between ’t Hooft-Polyakov monopoles”, *Nucl.Phys.* **B126** (1977) 525.
- [118] A. Dabholkar, “Microstates of nonsupersymmetric black holes”, *Phys.Lett.* **B402** (1997) 53, hep-th/9702050.
- [119] E. Bergshoeff, M. de Roo, E. Eyras, B. Janssen and J. P. van der Schaar, “Intersections involving waves and monopoles in eleven dimensions”, *Class.Quant.Grav.* **14** (1997) 2757, hep-th/9704120.
- [120] M. Cvetič and A. A. Tseytlin, “Non-extreme black holes from non-extreme intersecting M-branes”, *Nucl.Phys.* **B478** (1996) 181, hep-th/9606033.
- [121] I. Ya. Aref’eva, M. G. Ivanov and I. V. Volovich, “Non-extremal intersecting p -branes in various dimensions”, *Phys.Lett.* **B406** (1997) 44, hep-th/9702079.
- [122] N. Ohta, “Intersection rules for non-extreme p -branes”, *Phys.Lett.* **B403** (1997) 218, hep-th/9702164.
- [123] A. Brandhuber, N. Itzhaki, J. Sonnenschein and S. Yankielowicz, “More on probing branes with branes”, *Phys.Lett.* **B423** (1998) 238, hep-th/9711010.
- [124] A. Dhar and G. Mandal, “Probing 4-dimensional nonsupersymmetric black holes carrying D0- and D6-brane charges”, hep-th/9803004.
- [125] M. Berkooz, M. R. Douglas and R. G. Leigh, “Branes intersecting at angles”, *Nucl.Phys.* **B480** (1996) 265, hep-th/9606139.
- [126] J. C. Breckenridge, G. Michaud and R. C. Myers, “New angles on D-branes”, *Phys.Rev.* **D56** (1997) 5172, hep-th/9703041.
- [127] V. Balasubramanian, F. Larsen and R. G. Leigh, “Branes at angles and black holes”, *Phys.Rev.* **D57** (1998) 3509, hep-th/9704143.

- [128] N. Hambli, “Comments on Dirichlet branes at angles”, *Phys.Rev.* **D56** (1997) 2369, hep-th/9703179.
- [129] J. M. Izquierdo, N. D. Lambert, G. Papadopoulos and P. K. Townsend, “Dyonic Membranes”, *Nucl.Phys.* **B460** (1996) 560, hep-th/9508177.
- [130] M. S. Costa, “Composite M-branes”, *Nucl.Phys.* **B490** (1997) 202, hep-th/9609181.
- [131] M. S. Costa, “Black composite M-branes”, *Nucl.Phys.* **B495** (1997) 195, hep-th/9610138.
- [132] Z. Guralnik and S. Ramgoolam, “Torons and D-brane bound states”, *Nucl.Phys.* **B499** (1997) 241, hep-th/9702099.
- [133] R. R. Khuri, “Remark on string solitons”, *Phys.Rev.* **D48** (1993) 2947; hep-th/9305143.
- [134] J. D. Edelstein, L. Tataru and R. Tatar, “Rules for localized overlappings and intersections of p -branes”, hep-th/9801049.
- [135] N. Itzhaki, A. A. Tseytlin and S. Yankielowicz, “Supergravity solutions for branes localized within branes”, hep-th/9803103.
- [136] A. A. Tseytlin, “Composite BPS configurations of p -branes in ten-dimensions and eleven-dimensions”, *Class.Quant.Grav.* **14** (1997) 2085, hep-th/9702163.
- [137] L. J. Romans, “Massive N=2a supergravity in ten-dimensions”, *Phys.Lett.* **169B** (1986) 374.
- [138] P. K. Townsend, “Brane surgery”, *Nucl.Phys.Proc.Suppl.* **58** (1997) 163, hep-th/9609217.
- [139] M. B. Green, C. M. Hull and P. K. Townsend, “D-brane Wess–Zumino actions, T-duality and the cosmological constant”, *Phys.Lett.* **B382** (1996) 65, hep-th/9604119.
- [140] A. Sen, “String network”, *JHEP* 03(1998)005, hep-th/9711130.
- [141] O. Aharony, A. Hanany and B. Kol, “Webs of (p,q) 5-branes, five dimensional field theories and grid diagrams”, *JHEP* 01(1998)002, hep-th/9710116.
- [142] P. K. Townsend, “Membrane tension and manifest IIB S-duality”, *Phys.Lett.* **B409** (1997) 131, hep-th/9705160.
- [143] E. Bergshoeff, M. de Roo, M. B. Green, G. Papadopoulos and P. K. Townsend, “Duality of type II 7-branes and 8-branes”, *Nucl.Phys.* **B470** (1996) 113, hep-th/9601150.

- [144] J. Polchinski and A. Strominger, “New vacua for type II string theory”, *Phys.Lett.* **B388** (1996) 736, hep-th/9510227.
- [145] O. Bergman, M. R. Gaberdiel and G. Lifschytz, “Branes, orientifolds and the creation of elementary strings”, *Nucl.Phys.* **B509** (1998) 194, hep-th/9705130.
- [146] A. Hanany and A. Zaffaroni, “Chiral Symmetry from Type IIA Branes”, *Nucl.Phys.* **B509** (1998) 145, hep-th/9706047.
- [147] C. Bachas, M. R. Douglas and M. B. Green, “Anomalous creation of branes”, *JHEP* 07(1997)002, hep-th/9705074;
U. H. Danielsson, G. Ferretti and I. R. Klebanov, “Creation of fundamental strings by crossing D-branes”, *Phys.Rev.Lett.* **79** (1997) 1984, hep-th/9705084;
S. P. de Alwis, “A note on brane creation”, *Phys.Lett.* **B413** (1997) 49, hep-th/9706142.
- [148] M. Li, “Open membranes in Matrix theory”, *Phys.Lett.* **B397** (1997) 37, hep-th/9612144.
- [149] I. Bandos, K. Lechner, A. Nurmagambetov, P. Pasti, D. Sorokin and M. Tonin, “Covariant action for the super-five-brane of M-theory”, *Phys.Rev.Lett.* **78** (1997) 4332, hep-th/9701149;
M. Aganagic, J. Park, C. Popescu and J. H. Schwarz, “World-volume action of the M theory five-brane”, *Nucl.Phys.* **B496** (1997) 191, hep-th/9701166.
- [150] E. Witten, “Some comments on string dynamics”, in *Los Angeles 1995, Future perspectives in string theory* 501-523, hep-th/9507121.
- [151] R. Dijkgraaf, E. Verlinde and H. Verlinde, “BPS spectrum of the 5-brane and black-hole entropy”, *Nucl.Phys.* **B486** (1997) 77, hep-th/9603126; “BPS quantization of the 5-brane”, *Nucl.Phys.* **B486** (1997) 89, hep-th/9604055.
- [152] N. Seiberg, “New theories in six-dimensions and Matrix description of M-theory on T^5 and T^5/Z_2 ”, *Phys.Lett.* **B408** (1997) 98, hep-th/9705221.
- [153] A. Losev, G. Moore and S. L. Shatashvili, “M&m’s”, hep-th/9707250.
- [154] I. Brunner and A. Karch, “Matrix description of M-theory on T^6 ”, *Phys.Lett.* **B416** (1998) 67, hep-th/9707259.
- [155] A. Hanany and G. Lifschytz, “M(atrix) theory on T^6 and a m(atrix) theory description of KK monopoles”, hep-th/9708037.
- [156] A. Sen, “Dynamics of multiple Kaluza-Klein monopoles in M and string theory”, *Adv.Theor.Math.Phys.* **1** (1998) 115, hep-th/9707042; “A note on enhanced gauge symmetries in M and string theory”, *JHEP* 09(1997)001, hep-th/9707123.

- [157] A. Achucarro, J. M. Evans, P. K. Townsend and D. L. Wiltshire, “Super p -branes”, *Phys.Lett.* **198B** (1987) 441.
- [158] S. Elitzur, A. Giveon, D. Kutasov and E. Rabinovici, “Algebraic aspects of Matrix theory on T^d ”, *Nucl.Phys.* **B509** (1998) 122, hep-th/9707217.
- [159] N. Seiberg, “Notes on theories with 16 supercharges”, proceedings of the Trieste Spring School (March 1997), hep-th/9705117.
- [160] N. Seiberg and S. Sethi, “Comments on Neveu-Schwarz five-branes”, *Adv. Theor. Math. Phys.* **1** (1998) 259, hep-th/9708085.
- [161] P. S. Howe, N. D. Lambert and P. C. West, “The selfdual string soliton”, *Nucl.Phys.* **B515** (1998) 203, hep-th/9709014.
- [162] C. G. Callan and J. M. Maldacena, “Brane dynamics from the Born-Infeld action”, *Nucl.Phys.* **B513** (1998) 198, hep-th/9708147.
- [163] P. S. Howe, N. D. Lambert and P. C. West, “The three-brane soliton of the M-fivebrane”, *Phys.Lett.* **B419** (1998) 79, hep-th/9710033.
- [164] E. Bergshoeff, J. Gomis and P. K. Townsend, “M-brane intersections from world-volume superalgebras”, *Phys.Lett.* **B421** (1998) 109, hep-th/9711043.
- [165] G. W. Gibbons and S. W. Hawking, “Gravitational multi-instantons”, *Phys.Lett.* **78B** (1978) 430;
T. Eguchi P. B. Gilkey and A. J. Hanson, “Gravitation, gauge theories and differential geometry”, *Phys.Rept.* **66** (1980) 213.
- [166] J. M. Maldacena and A. Strominger, “Semiclassical decay of near extremal five-branes”, JHEP 12(1997)008, hep-th/9710014.
- [167] N. Evans, C. V. Johnson and A. D. Shapere, “Orientifolds, branes, and duality of 4D gauge theories”, *Nucl.Phys.* **B505** (1997) 251, hep-th/9703210.
- [168] J. D. Blum and K. Intriligator, “New phases of string theory and 6d RG fixed points via branes at orbifold singularities”, *Nucl.Phys.* **B506** (1997) 199, hep-th/9705044.
- [169] K. Intriligator, “New string theories in six dimensions via branes at orbifold singularities”, *Adv.Theor.Math.Phys.* **1** (1998) 271, hep-th/9708117.
- [170] I. Brunner and A. Karch, “Branes and six dimensional fixed points”, *Phys.Lett.* **B409** (1997) 109, hep-th/9705022; “Branes at orbifolds versus Hanany Witten in six dimensions”, JHEP 03(1998)003, hep-th/9712143.
- [171] J. de Boer, K. Hori, H. Ooguri, Y. Oz and Z. Yin, “Mirror symmetry in three-dimensional gauge theories, $SL(2, \mathbb{Z})$ and D-Brane moduli spaces”, *Nucl.Phys.* **B493** (1997) 148, hep-th/9612131.

- [172] O. J. Ganor and S. Sethi, “New perspectives on Yang-Mills theories with sixteen supersymmetries”, *JHEP* 01(1998)007, hep-th/9712071.
- [173] W. Taylor, “D-brane field theory on compact spaces”, *Phys.Lett.* **B394** (1997) 283, hep-th/9611042.
- [174] L. Susskind, “T duality in M(atrrix) theory and S duality in field theory”, hep-th/9611164.
- [175] O. J. Ganor, S. Ramgoolam and W. Taylor, “Branes, fluxes and duality in M(atrrix) theory”, *Nucl.Phys.* **B492** (1997) 191, hep-th/9611202.
- [176] M. Rozali, “Matrix theory and U duality in seven-dimensions”, *Phys.Lett.* **B400** (1997) 260, hep-th/9702136;
M. Berkooz, M. Rozali and N. Seiberg, “Matrix description of M-theory on T^4 and T^5 ”, *Phys.Lett.* **B408** (1997) 105, hep-th/9704089.
- [177] O. Aharony, M. Berkooz, S. Kachru, N. Seiberg and E. Silverstein, “Matrix description of interacting theories in six dimensions”, *Adv.Theor.Math.Phys.* **1** (1998) 148, hep-th/9707079.
- [178] E. Witten, “On the conformal field theory of the Higgs branch”, *JHEP* 07(1997)003, hep-th/9707093.
- [179] G. T. Horowitz and A. Strominger, “Counting states of near-extremal black holes”, *Phys.Rev.Lett.* **77** (1996) 2368, hep-th/9602051.
- [180] J. M. Maldacena and A. Strominger, “Statistical entropy of four-dimensional extremal black holes”, *Phys.Rev.Lett.* **77** (1996) 428, hep-th/9603060.
- [181] C. V. Johnson, R .R. Khuri and R. C. Myers, “Entropy of 4D extremal black holes”, *Phys.Lett.* **B378** (1996) 78, hep-th/9603061.
- [182] G. T. Horowitz, D. A. Lowe and J. M. Maldacena, “Statistical entropy of nonextremal four-dimensional black holes and U-Duality”, *Phys.Rev.Lett.* **77** (1996) 430, hep-th/9603195.
- [183] T. Banks, W. Fischler, I. R. Klebanov and L. Susskind, “Schwarzschild black holes from Matrix theory”, *Phys.Rev.Lett.* **80** (1998) 226, hep-th/9709091.
- [184] I. R. Klebanov and L. Susskind, “Schwarzschild black holes in various dimensions from Matrix theory”, *Phys.Lett.* **B416** (1998) 62, hep-th/9709108.
- [185] M. Li, “Matrix Schwarzschild black holes in large N limit”, *JHEP* 01(1998)009, hep-th/9710226.
- [186] G. T. Horowitz and E. M. Martinec, “Comments on black holes in Matrix theory”, *Phys.Rev.* **D57** (1998) 4935, hep-th/9710217.

- [187] T. Banks, W. Fischler, I. R. Klebanov and L. Susskind, “Schwarzschild black holes in Matrix theory II”, *JHEP* 01(1998)008, hep-th/9711005.
- [188] N. Ohta and J.-G. Zhou, “Euclidean path integral, D0-branes and Schwarzschild black holes in Matrix theory”, hep-th/9801023.
- [189] K. Sfetsos and K. Skenderis, “Microscopic derivation of the Bekenstein-Hawking entropy formula for non-extremal black holes”, *Nucl.Phys.* **B517** (1988) 179, hep-th/9711138.
- [190] M. Bañados, C. Teitelboim and J. Zanelli, “The black hole in three dimensional space-time”, *Phys.Rev.Lett.* **69** (1992) 1849, hep-th/9204099;
M. Bañados, M. Henneaux, C. Teitelboim and J. Zanelli, “Geometry of the (2+1) black hole”, *Phys.Rev.* **D48** (1993) 1506, gr-qc/9302012.
- [191] S. Carlip, “The statistical mechanics of the (2+1)-dimensional black hole”, *Phys.Rev.* **D51** (1995) 632, gr-qc/9409052; “The statistical mechanics of three dimensional euclidean black hole”, *Phys.Rev.* **D55** (1997) 878; gr-qc/9606043.
- [192] A. Strominger, “Black hole entropy from near horizon microstates”, hep-th/9712251.
- [193] D. Birmingham, I. Sachs and S. Sen, “Entropy of three-dimensional black holes in string theory”, hep-th/9801019.
- [194] J. M. Maldacena and A. Strominger, “AdS3 black holes and a stringy exclusion principle”, hep-th/9804085.
- [195] E. J. Martinec, “Matrix models of AdS gravity”, hep-th/9804111.
- [196] K. Behrndt, I. Brunner and I. Gaida, “Entropy and conformal field theories of AdS₃ Models”, hep-th/9804159.
- [197] S. R. Das, S. D. Mathur, S. Kalyana Rama and P. Ramadevi, “Boosts, Schwarzschild black holes and absorption cross-sections in M-theory”, hep-th/9711003.
- [198] A. A. Tseytlin, “On the structure of composite black p -brane configurations and related black holes”, *Phys.Lett.* **B395** (1997) 24, hep-th/9611111.
- [199] F. Englert and E. Rabinovici, “Statistical entropy of Schwarzschild black holes”, to appear in *Phys.Lett. B*, hep-th/9801048.
- [200] C. M. Hull and B. Julia, “Duality and moduli spaces for timelike reductions”, hep-th/9803239;
E. Cremmer, I. V. Lavrinenko, H. Lü, C. N. Pope, K. S. Stelle and T. A. Tran, “Euclidean-signature supergravities, dualities and instantons”, hep-th/9803259.

- [201] J. M. Maldacena, “Black holes in string theory”, Princeton University Ph.D. Thesis (1996), hep-th/9607235.
- [202] J. M. Maldacena and L. Susskind, “D-branes and fat black holes”, *Nucl.Phys.* **B475** (1996) 679, hep-th/9604042.
- [203] S. R. Das, S. D. Mathur and P. Ramadevi, “Hawking radiation from four-dimensional Schwarzschild black holes in M theory”, hep-th/9803078.
- [204] P. K. Townsend, “M theory from its superalgebra”, hep-th/9712004.